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# The Band Spectrum of Carbon Monoxide

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UNITED STATES DEPARTMENT OF COMMERCE • John T. Connor, *Secretary*

NATIONAL BUREAU OF STANDARDS • A. V. ASTIN, *Director*

# The Band Spectrum of Carbon Monoxide

Paul H. Krupenie

Institute for Basic Standards  
National Bureau of Standards  
Washington, D.C.



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6	Chemical Kinetics
7	Colloid and Surface Properties
8	Mechanical Properties of Materials

The present compilation is in category 3 of the above list. It constitutes the fifth publication in a new NBS series known as the National Standard Reference Data Series.

A. V. ASTIN, *Director*.

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# The Band Spectrum of Carbon Monoxide\*

Paul H. Krupenie

This is an exhaustive review of the literature and a critical compilation of the observed and predicted spectroscopic data on CO, CO<sup>+</sup>, and CO<sup>2+</sup> in the gas phase.

Key Words: Carbon monoxide, spectrum, review.

## 1. Introduction

There exists no unified compilation of spectroscopic data for diatomic molecules. The standard references in use, now incomplete and out of date, are the tables of molecular constants in Herzberg's book [99]<sup>1</sup> and the more extensive compilations by Rosen et al., [213] which include band positions and perturbations. Rapid accumulation of new data has made apparent the need for a current critical review and more extensive compilation. This need has been partially met by recent tabulations of band head wavelengths by Pearse and Gaydon [187], and the compilation of these and other data for a select group of diatomic molecules of astrophysical interest by Wallace [266]. The only recent review known to the author which is devoted to a single molecule is that of Lofthus [153], "The Molecular Spectrum of Nitrogen."

The present work, first in a series on diatomic molecules,<sup>2</sup> follows the approach of Lofthus and is devoted to a single molecule. This report includes a comprehensive review of the literature on the spectrum of CO, CO<sup>+</sup>, and CO<sup>2+</sup> in the gas phase,<sup>3</sup> and a compilation of critically evaluated numerical data on band positions, molecular constants, energy levels, potential energy curves, and other molecular properties derived from the spectrum. Estimates of reliability are given where possible.

Papers from which the tabulated data have been extracted are discussed in the text of this report. Early data of presumably low precision have been included in the tabulations only where better or more recent data are not available. The references cited constitute a critical bibliography rather than an exhaustive one. The author will be

indebted to readers who bring to his attention significant references not included in the bibliography.

No spectrograms are included in the report (but see table 1 for references with reproductions of spectra). However, R. W. Nicholls of the Department of Physics, York University Toronto, Ontario, Canada has forthcoming identification atlases for various band systems including several for carbon monoxide.

Discussions of infrared intensities, absorption coefficients, and line shapes and widths are omitted here; literature on these topics can be traced from references in the recent paper by Benedict et al., [18].

Carbon monoxide is an important molecule for the following reasons: (1) It is a product of combustion of organic molecules and plays an important role in the reactions of flames; (2) it appears as an impurity in many systems and its spectrum is readily excited, especially the strong A-X Fourth Positive system in the UV; (3) it is a constituent of the solar chromosphere, stellar atmospheres, and comet tails [CO<sup>+</sup>], [the increase in rocket and satellite observations will enhance its importance]; (4) because it is a light molecule its electronic structure is amenable to theoretical analysis; (5) it is isoelectronic with N<sub>2</sub> which is important as the primary constituent in the earth's atmosphere; and (6) it can be obtained in pure form and is readily handled in the laboratory.

The spectrum of CO is dominated by three prominent band systems: A<sup>1</sup>Π—X<sup>1</sup>Σ<sup>+</sup> Fourth Positive system (2800–1140 Å), b<sup>3</sup>Σ<sup>+</sup>—a<sup>3</sup>Π Third Positive system (3800–2600 Å), and B<sup>1</sup>Σ<sup>+</sup>—A<sup>1</sup>Π Ångstrom system (6600–4100 Å). The A-X system is the most pronounced and extensive, with more than 150 bands being observed. For CO<sup>+</sup> the A<sup>2</sup>Π<sub>i</sub>—X<sup>2</sup>Σ<sup>+</sup> Comet Tail system (7200–3080 Å) is the strongest. This system may be

\*Work supported in part by the National Aeronautics and Space Administration.

<sup>1</sup> Figures in brackets indicate the literature references on page 78.

<sup>2</sup> The second report in this series will be devoted to O<sub>2</sub>.

<sup>3</sup> Information on CO<sup>2+</sup> is obtained from electron impact.

important in radiative heating of hypersonic vehicles at escape velocity or in the atmospheric fringe of the planet Venus (which contains CO<sub>2</sub>).

Study of the carbon monoxide spectrum in the visible region originated with the work of Ångström and Thalén in 1875. The earliest band head measurements in the UV were made by Deslandres in 1888, and by Lyman shortly after the turn of this century. All of the early observations on CO have been cataloged in volumes V and VII of Kayser's *Handbuch der Spektroskopie* [127], but they can most easily be traced from an early review on the band spectra of CO by Birge [23].

The modern study of the spectrum of carbon monoxide began in 1926. In that year, Birge [22, 23] obtained the vibrational analysis of the A<sup>1</sup>Π—X<sup>1</sup>Σ<sup>+</sup> Fourth Positive system, which, together with the correlation of emission and absorption measurements, enabled him to draw an energy level diagram for the observed electronic states. The impetus provided by the development of the quantum theory together with improved spectroscopic techniques resulted in extensive work in the 1930's on rotational analyses, and study of perturbations and predissociations. In the early

1940's, there appeared the first studies of Rydberg series, followed a decade later by the beginning of precise infrared vibration-rotation spectral measurements and microwave studies. Controversy of 30 years standing over the value of the dissociation energy of CO appears to have been convincingly settled by Douglas and Møller's re-examination [56] and review of the suspected predissociations of the A<sup>1</sup>Π state.

The mid-1950's ushered in renewed interest in forbidden transitions and new states. Two bands previously thought to originate from high vibrational levels of the a'<sup>3</sup>Σ<sup>+</sup> state (though questioned by Garg [68], Herzberg [99], and Gaydon [69]), have been indirectly found to originate from the f<sup>3</sup>Σ<sup>+</sup> state by more recent work of Herzberg and Hugo [101]. The I<sup>1</sup>Σ<sup>-</sup> state, previously known only from its perturbations of the A<sup>1</sup>Π state, has been observed in forbidden transition to the X<sup>1</sup>Σ<sup>+</sup> state. Complete details are not yet available.

Current effort is still centered more on the determination of energy levels (i.e., precise measurements of line positions) rather than on the more difficult intensity measurements.

## 2. Electronic Structure of CO and CO<sup>+</sup>

The early work of Mulliken [167, 168], largely intuitive and semiempirical, has laid the foundation for our understanding of the electronic structure and spectroscopic properties of carbon monoxide. Recent numerical calculations qualitatively account for a number of these features, though quantitative predictions are generally less reliable. The order of bound electronic states lying below 11 eV has been correctly predicted, but not their numerical term values. Energies of some unobserved electronic states have been estimated. Calculated ionization potentials agree fairly well with experimental values. Hartree-Fock calculations question the experimentally determined polarity of the dipole moment. Theoretical prediction of the identity of the E state as <sup>1</sup>Π has been confirmed experimentally. Poor

values are obtained for the ground state binding energy and potential energy curves of low lying states.

This section reviews what is known about the electronic structure of CO and CO<sup>+</sup> and briefly summarizes the results of numerical calculations. Elements of molecular orbital theory and definitions of terms used are given elsewhere [200, 250, 172]. Abbreviations used include: IP, ionization potential; MO, molecular orbital; LCAO, linear combination of atomic orbitals; CI, configuration interaction; SCF, self-consistent field; a.u., atomic units (=27.210 eV).

### 2.1. Electronic Structure

The relative order of the molecular orbitals for CO is as follows [46, 171, 168, 109, 173]:

(a)	σ1s	<	σ*1s	<	σ2s	<	σ*2s	<	π <sub>y</sub> 2p=π <sub>z</sub> 2p	<	σ2p	<	π <sub>y</sub> *2p=π <sub>z</sub> *2p	<	σ*2p
(b)	K		K		zσ		yσ		wπ		xσ		vπ		uσ
(c)	1σ		2σ		3σ		4σ		1π		5σ		2π		6σ
	inner		inner		b.		a.b.		b.		a.b.		a.b.		a.b.



Orbital designation (a) specifies the atomic origin of the MO's; designations (b) and (c) in current use, do not. The bonding character of the orbitals is specified above as: bonding, antibonding, or nonbonding.  $1\sigma$  and  $2\sigma$  orbitals virtually retain their purely atomic character and are often called K or inner orbitals.  $3\sigma$  is primarily O( $2s$ );  $4\sigma$  is a mixture of C( $2s$ ) and O( $2p$ ), though the actual ratio is a function of  $r$ . The  $5\sigma$  (slightly *a.b.*) and  $1\pi$  orbitals have similar energies and can be expected to have their order reversed at some value of  $r$ . Isoelectronic molecules have similar MO's and roughly similar sequence of MO's, as is the case for CO and N<sub>2</sub> up to about 11 eV. At higher energies this similar order is destroyed.

The assignment of the electronic structure and dissociation products of a number of states of CO and CO<sup>+</sup> has been given by Mulliken [168]. The ground state configuration is

$$X^1\Sigma^+(1\sigma)^2(2\sigma)^2(3\sigma)^2(4\sigma)^2(1\pi)^4(5\sigma)^2.$$

An examination of the ionization process from successively deeper lying orbitals shows why only three states are known for CO<sup>+</sup>. This is summarized below:

Electron ionized	Ionic state	I.P.(eV) <sup>4</sup>
$5\sigma$	X <sup>2</sup> Σ <sup>+</sup>	14. 013
$1\pi$	A <sup>2</sup> Π <sub>i</sub>	16. 544
$4\sigma$	B <sup>2</sup> Σ <sup>+</sup>	19. 671
$3\sigma$	<sup>2</sup> Σ <sup>+</sup>	41 ± 3

The ionization of the  $3\sigma$  electron has not been observed. There is some evidence for a state of CO<sup>+</sup> at 25.6 eV (± 0.5 eV) from electron impact [151] and photoionization [270] studies. (This has been called the C<sup>2</sup>Σ<sup>+</sup> state, but it does not correspond to the removal of a  $3\sigma$  electron.) Codling [42] has recently observed structure in the 500 Å region which appears to include a Rydberg series whose convergence limit lies above 20 eV. By analogy with N<sub>2</sub><sup>+</sup> the electron configuration of this limiting state may result from a two-electron excitation. States of CO<sup>2+</sup>, detected at 41.8 eV and 45.7 eV in electron impact

<sup>4</sup> The three lowest IP have been determined experimentally from Rydberg series limits and electronic spectra of CO<sup>+</sup> (see. 5.3). Only a calculated value is available for the IP of the  $3\sigma$  electron [109, 173]. Long ago Mulliken [169] estimated this quantity as 32 eV.

experiments [55, 263], have not been observed spectroscopically. Hurley and Maslen [111] have calculated the lower state energy to be 41.17 eV.

Ground state atoms C(<sup>3</sup>P)+O(<sup>3</sup>P) give rise to 18 possible electronic states of CO (table 2). Only eight of these have been observed: X<sup>1</sup>Σ<sup>+</sup>, a<sup>3</sup>Π, d<sup>3</sup>Δ<sub>i</sub>, e<sup>3</sup>Σ<sup>-</sup>, a'<sup>3</sup>Σ<sup>+</sup>, f<sup>3</sup>Σ<sup>+</sup>, A<sup>1</sup>Π, and I<sup>1</sup>Σ<sup>-</sup>; all the others are probably weakly bound or repulsive, though the <sup>1</sup>Δ state might be slightly stable [167]. However, a (<sup>1</sup>Π<sup>1</sup>Π) or (<sup>1</sup>Δ<sup>1</sup>Π) perturbation at low J in A<sup>1</sup>Π, v=0, 1 at about 8 eV can only come from a stable state [44, 71, 216]. Therefore, if real, these perturbations would most likely arise from one of the unobserved states. A possible (<sup>1</sup>Δ<sup>3</sup>Δ) perturbation of the d state indicates the presence of a <sup>1</sup>Δ state at about the same energy [137]. By analogy with N<sub>2</sub> such a state would lie at about 9 eV.

Kaplan [126] has reported a state of unknown type at about 4.8 eV above X. The reality of a state of CO at this energy is, however, highly questionable [168, 99, 187].

For the higher energy excited molecular states beginning with C<sup>1</sup>Σ<sup>+</sup>, at least one excited atom is required in the dissociation products, since the electronic energy alone exceeds the heat of dissociation into ground state atoms. In the case of singlet states of CO not formed from ground state atoms, both atoms must be excited. Table 1 is a summary of the known molecular constants and includes the assigned electron configuration and dissociation products where possible, for the observed electronic states. Table 2 gives the lowest lying configurations and possible molecular states, with the observed states indicated.

Mulliken [171] has recently compared the observed lower excited states of the π<sup>3</sup>π configuration with those predicted for CO. The d-state of CO was proposed to be a case (b) π<sup>3</sup>π<sup>3</sup>Δ in disagreement with the earlier experimentally determined assignment as d<sup>3</sup>Π<sub>i</sub>. Carroll [40] experimentally confirmed Mulliken's assignment of the d state as <sup>3</sup>Δ, but definitely established it to be inverted. A π<sup>3</sup>π configuration gives rise to a <sup>3</sup>Δ<sub>r</sub> state in the usual approximation. A theoretical interpretation of the anomalous multiplet splitting in the d<sup>3</sup>Δ term has been given by Kovács [137].

The π<sup>3</sup>π states are <sup>1</sup>Σ<sup>+</sup>, I<sup>1</sup>Σ<sup>-</sup>, <sup>1</sup>Δ, a'<sup>3</sup>Σ<sup>+</sup>, e<sup>3</sup>Σ<sup>-</sup>, and d<sup>3</sup>Δ<sub>i</sub>. Recent information [100, 238a] indicates the I state lies at about 66000 cm<sup>-1</sup>, and the e state at about 63000 cm<sup>-1</sup>. The <sup>1</sup>Δ state and <sup>1</sup>Σ<sup>+</sup> remain unobserved.



## 2.2. Numerical Calculations

### a. Semiempirical

Semiempirical studies have provided a qualitative understanding of several properties of the carbon monoxide molecule. The triple bond for the ground state has been accounted for by Sahni [217], Mulliken [170], and Linnett [152]. (See also ref. [154].) Moffitt [164] has correlated the decrease in bond length for states of CO and CO<sup>+</sup> with increase in  $\pi$ -bond order. Sahni [217], with neglect of 1s electrons and the use of many approximate integrals, has derived individual and total orbital electron densities which are shown in contour diagrams. The lowest ionization of CO, that is, removal of an almost nonbonding electron from the 5 $\sigma$  orbital, has been shown loosely equivalent to removal of an electron from the carbon atom [217, 170, 152]. Indeed, the dissociation products of the ground state of CO<sup>+</sup> are C<sup>+</sup>+O.

### b. Single Configuration

Several calculations for CO have been made using single configuration LCAO-MO-SCF wave functions. Ground state energies have been calculated by Ransil [199, 200] and Hurley [110] using functions constructed from minimal basis sets. Merryman et al. [159], and Brion and Moser [29] have calculated excitation energies for a number of states from similar functions. For states lying above 9 eV the calculated energies are too high. To describe such states it is necessary to include more than one configuration or to introduce atomic orbitals of higher quantum number. Lefebvre-Brion et al. [146], have compared the vertical excitation energies calculated from SCF functions built from extended basis sets. Calculated values agree fairly well with experiment, even for states lying above 11 eV.

The most accurate wave functions for the ground state of CO have recently been reported by Wahl and Huo [265] and Huo [109]. Huo has calculated single determinantal SCF wave functions from two extended basis sets, including one built by extending the basis set necessary for obtaining accurate Hartree-Fock functions for the atoms. The calculated value of the total molecular energy is -112.7877 a.u. This result was obtained from a minimal basis set supplemented by 1s<sub>0</sub>', 1s<sub>c</sub>', 2s<sub>0</sub>', 2s<sub>c</sub>', 2p $\pi$ <sub>0</sub>', 2p $\pi$ <sub>c</sub>', 2p $\sigma$ <sub>0</sub>', 2p $\sigma$ <sub>c</sub>', 3d $\sigma$ <sub>0</sub>', 3d $\sigma$ <sub>c</sub>', 3d $\pi$ <sub>0</sub>', 3d $\pi$ <sub>c</sub>', 2p $\pi$ <sub>0</sub>'', 4f $\pi$ <sub>c</sub>', 3s<sub>0</sub>, and 3s<sub>c</sub>. Expectation values

of several one-electron operators are given as well as contour diagrams for orbital and total charge densities [109(b)].

Expectation values of a number of one-electron operators were calculated, including the dipole moment whose polarity was determined as C<sup>+</sup>O<sup>-</sup>, in disagreement with the reverse assignment given by Rosenblum, Nethercot, and Townes [215] from microwave experiments. Huo [109] and Nesbet [173] have independently concluded that more accurate SCF functions should not change the sign of the dipole moment, and that the polarity should be considered undetermined.

Nesbet [173] has calculated SCF functions for the ground state of CO using a double exponential basis set with optimized exponents for each occupied atomic orbital. This basis set appreciably lowers the total energy but makes only a small change in the binding energy from that using a smaller basis set. The addition of d $\sigma$  and d $\pi$  orbitals makes only a small change in the total energy but makes a larger improvement in the binding energy.

### c. Configuration Interaction

Configuration interaction has been used to calculate the correlation energy [experimental total energy—Hartree-Fock energy]. Combined use of extended basis sets and CI results in improved values of excited state energies and gives lower calculated total energies. Lefebvre-Brion et al. [144], have found that use of a basis set which includes functions from the M shell (3s and 3p) with Slater exponents improves the calculated vertical excitation energies more than does doubling the number of L shell functions. By using orbitals constructed from a minimal basis set and including CI they have also [145] calculated potential energy curves for a number of states. These include repulsive <sup>3</sup>II and <sup>1</sup>II states which lie at 11.5 and 12 eV, respectively. The latter value agrees with the vertical excitation energy of the F state and suggests that the F state is not <sup>1</sup>II. [See also ref. 146.] The fragmentary data in section 5 suggest that the F state may be triplet. Lefebvre-Brion [147] predicts the state to be <sup>1</sup> $\Sigma^+$  (table 1).

Fraga and Ransil [66] have calculated the molecular energy and dipole moment for the ground state of CO using a limited CI with minimal SCF wave functions. Even with 14 configurations (having the same symmetry as the ground state)



$E_{\text{calc}} - E_{\text{obs}}$  is about 27 eV. These results indicate the need [66] to include MO's with different  $\lambda$  values and atomic orbitals of different  $l$  values from those utilized in the ground state.

### 2.3. Rydberg States and Rydberg Series

The observed Rydberg states in CO (table 1 supplement) are most probably singlet, and either  $\Pi$  or  $\Sigma^+$ , since they have all been detected in absorption series. There may exist others which do not combine with the ground state and could be observed in emission, but to date none have been reported. The electronic structure of the states considered is given below:

Configuration	Possible Rydberg states	Convergence limit [CO <sup>+</sup> ]
$\sigma(R\sigma)$	$^1\Sigma^+, ^3\Sigma^+$	X, B
$\sigma(R\pi)$	$^1\Pi, ^3\Pi$	
$\pi^3(R\sigma)$	$^1\Pi, ^3\Pi_i$	A
$\pi^3(R\pi)$	$^1\Sigma^+, ^1\Sigma^-, ^1\Delta, ^3\Sigma^+, ^3\Sigma^-, ^3\Delta$	

For the most likely Rydberg orbitals  $np\pi$  or  $np\sigma$ , the most likely states are  $\Pi$  or  $\Sigma$ . Huber [106] believes that the states converging to the X and B states of CO<sup>+</sup> are  $^1\Sigma^+$  and the  $\alpha$  states converging to A<sup>2</sup> $\Pi_i$  are  $^1\Pi$ , all having Rydberg orbitals  $np\sigma$ . Tanaka [254] has speculated that the sharp  $\beta$  series is composed of  $^1\Sigma^+$  states and the diffuse  $\beta$  states are  $^1\Pi$ . With present know-

ledge it is not possible to specify uniquely their symmetry types.

Rydberg states with orbital  $n$  dissociate into atoms at least one of which is configurationally excited (i.e., having principal quantum number greater than in its ground state,  $(n-1) \geq 3$ ). Huber [106] suggests that the dissociation products of series converging to states X and A are C( $n-1$ )s( $^3P^0$ ) + O( $^3P$ ), and states converging to B have products C( $n-1$ )s( $^1P^0$ ) + O( $^1D$ ).

Lefebvre-Brion, Moser, and Nesbet [147] have recently calculated the energies of Rydberg levels of CO which are of symmetry type  $^1,^3\Sigma^+$  and  $^1,^3\Pi$  and lie below the first ionization potential. The calculated energies fall within 0.2 eV of the observed values for the B $^1\Sigma^+$ , C $^1\Sigma^+$ , b $^3\Sigma^+$ , and c $^3\Sigma^+$  states which probably belong to Rydberg series converging to the ground state of CO<sup>+</sup>. The configuration interaction functions used were built from LCAO-MO-SCF orbitals used previously [146, 148] for calculations on the ground state of CO, but with extended basis sets.

The deviations between these calculated energies and experiment indicate that the E state is  $^1\Pi$ , and the F state may be  $^1\Sigma^+$ . The E state assignment has been confirmed experimentally by Tilford et al. [257]; the tentative assignment of the F state as  $^1\Sigma^+$  is untested. (See sec. 5.2f and table 1, footnote F( $^1\Sigma^+$ ).) No  $^3\Pi$  Rydberg levels of CO have been observed to date. Codling [42] has recently observed Rydberg series converging to what may be a new state of CO<sup>+</sup> above 20 eV, but the symmetry and type of Rydberg levels have not been identified as yet.

## 3. Electronic Spectrum of CO and CO<sup>+</sup>

The spectral region 8600–600 Å includes all the presently known electronic transitions of CO and CO<sup>+</sup>. Intense systems span most of the region between 7500–1100 Å. From 3800 Å to longer wavelengths are found the B–A Ångstrom bands and the *d*–*a* Triplet system, the latter system composed of many bands, only four of which have had rotational analysis. Between 7200–3000 Å is found the A–X Comet Tail system of CO<sup>+</sup>, the most prominent system of the ion. From 3800–2600 Å is the *b*–*a* Third Positive system which is most important because of the most violent perturbations and predissociation of the *b* state. The

A–X Fourth Positive system dominates the region between 2800–1140 Å, though the shorter wavelength bands are rather weak. Below 1200 Å are found, in general, weak systems, incompletely studied, for which no rotational analysis is available. In total 29 transitions are found among the 23 known states of CO. Three states of CO<sup>+</sup> partake in three observed transitions. Details of the observed transitions are discussed below, including the determination of vibrational and rotational constants, multiplet splitting, and coupling cases of the electronic states. The critically evaluated data pertaining to these transitions have been tabulated in section 12.



### 3.1. $A^1\Pi-X^1\Sigma^+$ Fourth Positive System (2800–1140 Å) R

The  $A^1\Pi-X^1\Sigma^+$  Fourth Positive system, originally attributed to carbon, is the most prominent system of carbon monoxide in the ultraviolet and vacuum UV regions. The single-headed red-degraded bands of this allowed transition are about a thousand times stronger than those due to forbidden transitions which they overlap. Wavelengths of more than 150 emission bands which have been observed in the region 2800–1280 Å are assembled in table 3a. Included among these are early measurements by Lyman and Deslandres of 30 band heads whose vibrational assignments are due to Birge [23], and which have not been reported by anyone else. Band origins for this system are given in table 4. Rotational constants for the X and A states are given in tables 36 and 37, respectively. Positions of a number of unidentified bands have been assembled in table 35a. These have been observed together with the A–X system and possibly belong to it.

Attention is focused on the A–X system by the numerous perturbations of the  $A^1\Pi$  state, a detailed discussion of which is given in section 4.1. The recently observed  $I^1\Sigma^-$  state was first identified from its interaction with the A state. Several predissociations of the A state have been reported which have led to conflicting values of the dissociation energy of CO. Re-examination of these by Douglas and Møller [56] has shown that none of them are real.

The work before 1926, including the earliest band head measurements by Deslandres and Lyman, was discussed and analyzed for the first time by Birge [23]. All observations prior to that date were in emission, except for Leifson's [149] observation of the strong  $v''=0$  progression (with  $v'=0$  to 11) in absorption. The vibrational analysis of this system [23, 22] and the correlation of Lyman's emission bands with the absorption bands of Leifson showed that they were due to neutral CO. Thus an energy level diagram could be drawn with the lower state of this system established as the ground state.

Estey [60] remeasured part of the A–X system in the region 2800–1970 Å, and observed 16 new bands which involve high vibrational levels of the ground state. His source was a low pressure discharge, viewed with a quartz spectrograph having dispersion 3.91 Å/mm–1.21 Å/mm. Each band clearly shows a single P, Q, and R branch

as is expected for a  $^1\Pi-^1\Sigma$  transition. Less extensive band head measurements by Wolter [277] agree with those of Estey to within 0.1 Å.

Headrick and Fox [90] measured bands in the region 2170–1280 Å, using a 1-m focal length vacuum spectrograph. Several new bands were reported, mostly at shorter wavelength. The source was a low voltage arc in CO at 0.5 mm pressure with a trace of hydrogen. Wavelengths obtained from the use of two different standards agreed to within 0.2 Å, and are close to the values of Estey [60] where they overlap near 2000 Å. Dispersion was 17.4 Å/mm. Intensities, originally given as densitometer readings, have been converted in this report to a scale with the maximum reading taken as 10. In table 3a, for the bands above 2000 Å, the intensities are those given primarily by Estey [60]; below 2000 Å, the intensities are those of Headrick and Fox [90]. Several strong unidentified bands at short wavelengths have been included in table 35a.

Read [209] has given the vibrational analysis of 65 heads, observed in a hollow cathode discharge in flowing CO<sub>2</sub>, and viewed with a 2-m focal length vacuum spectrograph with dispersion of 4.2 Å/mm. In addition, he resolved the rotational structure for 12 bands, but details only of the 1–1 and 4–10 bands have been published [209b]. This was the first rotational analysis of bands involving the ground state of CO.  $\Lambda$ -doubling was found to be negligible, as had been determined from earlier studies of the  $B^1\Sigma-A^1\Pi$  Ångström bands.

A formula was obtained which fit the newly reported band heads as well as those of Estey at longer wavelength:

$$\sigma_H(v'-v'') = 64756.3 + (1497.49v' - 17.1841v'^2) \\ - (2155.61v'' - 13.2843v''^2) \\ + 0.012v'^3 - 0.0804v'v''.$$

The derived vibrational constants are very close to the best values listed in table 1.

The papers of Estey [60], Headrick and Fox [90], and Read [209] include almost all of the emission band head measurements of the A–X system. Shortly after their publication there began to appear extensive rotational analyses, mostly by Schmid and Gerö.

In a Geissler discharge in CO<sub>2</sub>, Gerö [73] observed a series of bands in the region 2700–1950 Å. Numerous perturbations were easily recognized in all branches. Using a 6.5-m focal length grating



with dispersion of 1.2 Å/mm, he obtained the rotational structure of the 4-11, 5-13, 6-15, 7-16, 8-17, 9-19, 10-20, and 10-21 bands. Gerö further reported [74] rotational structure for the 9-18, 11-20, 12-22, and 13-24 bands, and, in addition, gave a Deslandres table of 84 band origins calculated from the band heads with a positive correction of  $\frac{(B'+B'')}{4(B'-B'')}$ , making use of previous data [209, 73].

Table 4 gives an extensive list of band origins which includes several additions of low accuracy which supplement the Deslandres table of Herzberg [99, p. 156-7]. Herzberg's table is mainly that of Gerö [74] with slight modifications and numerous additions of low accuracy. The A-X system origin is at 64746.5 cm<sup>-1</sup> [209].

The heads of the 14-23, 15-25, and 16-25 bands listed by Estey [60] deviate by +13.2, +14.8, and +21.7 cm<sup>-1</sup>, respectively, from Read's formula. Gerö [74] believed that the appearance of these bands was due to accidental accumulation of lines, where strong overlapping is observed. The first of these heads lies close to a band reported by Kaplan [125] at 2518 Å, and the third lies close to a band of the 3A system (*c-a*). Gerö concluded from this evidence that the *v*=13 level of the A state was the last before dissociation. However, Tanaka, Jursa, and LeBlanc [255] have since observed the A state to *v*=20.

Schmid and Gerö [228] reported the measurement and rotational analysis of many bands by Kelemen [in a dissertation], including 1-8, 2-9, 2-10, 3-9, 3-10, 3-12, 4-10, 4-12, 4-13, 5-11, 5-12, 5-14, 6-13, 6-14, 7-14, 7-15, 7-17, 8-16, 8-18, 8-19, 9-17, 9-20, and 10-19. From these data and those of Read [209] and Gerö [73, 74], they obtained formulas for the rotational constants for both the X<sup>1</sup>Σ<sup>+</sup> and A<sup>1</sup>Π states. Schmid and Gerö [228] suggested that Read should have included a higher power of *J* in his expression for *F<sub>v</sub>*(*J*), which would have resulted in slightly larger *B<sub>v</sub>* values. The calculation of *B<sub>v</sub>* for the strongly perturbed A state follows the method of Gerö [71]. Least squares fitting to the observed data gives for A<sup>1</sup>Π : *B<sub>v</sub>*=1.6116-0.02229(*v*+½)-0.000105(*v*+½)<sup>2</sup> (table 1). Brons' [31] rotational analysis of the 10-19, 10-20, and 10-21 bands gives *B'<sub>10</sub>*, *B''<sub>19</sub>*, *B''<sub>20</sub>*, and *B''<sub>21</sub>*, slightly smaller than the values of Schmid and Gerö [228].

Tschulanowsky and Stepanow [262] have given the rotational analysis of the 3-8, 3-7, 2-8, 2-7,

2-6, 1-6 and 1-5 bands, including the determination of rotational constants and a discussion of various perturbations of the A state. Twenty-four bands were observed with a dispersion of 8.4 Å/mm (with resolution of lines separated by 0.07 Å), and details of the three best are given. Branches were observed up to *J* values of 35. The three best bands for analysis were 3-8, 2-7, and 1-6. The R branches were measured very inaccurately and their details are not given.

Onaka [182] has looked at the 6-13 and 6-14 bands with a vacuum echelle spectrograph of high resolution (~170,000) and high dispersion (0.30-0.26 Å/mm). Earlier attempts at analysis of the lower rotational levels in A<sup>1</sup>Π, *v*=6 was prevented by the complicated structure due to overlapping of the *v*=6 level by levels of the *d*<sup>3</sup>Δ<sub>i</sub> and *a*<sup>3</sup>Σ<sup>+</sup> states. Several perturbations were found (see sec. 4.1). The lines of the 6-14 band are not given. A formula for the rotational structure of the *v*=6 level is given by Onaka from an unpublished work on the Fourth Positive system:

$$T_{v=6}(A^1\Pi)=74197.44+1.4616J(J+1).$$

Tanaka, Jursa, and LeBlanc [255] reported on the absorption spectrum of the Fourth Positive system in the vacuum UV for the first time since the pioneer work of Leifson [149], and Hopfield and Birge [104]. Using rare gas continua as sources, Tanaka et al., observed new bands and extended known bands to shorter wavelengths, all under low resolution, covering the spectral region of 1177-1544 Å. The (*v*'-0) progression was observed for *v*'=0 to 20; also the (*v*'-1) progression for *v*'=0 to 13. In addition, the isotopic C<sup>13</sup>O<sup>16</sup> band progression (*v*'-0) was observed up to *v*'=17 (not including 16). Tanaka et al., used pure CO in an 8-cm long absorption cell, at pressures of 0.01 to 600 mm of Hg. The spectrograph used was a 2-m focal length normal incidence type with a dispersion of about 8 Å/mm. The absorption band positions are given in table 3b; the observed isotopic bands of C<sup>13</sup>O<sup>16</sup> are given in table 5. Several unidentified bands which were found are given in table 35b.

McCulloh [157a] and McCulloh and Glockler [157b] first studied emission by C<sup>13</sup>O<sup>18</sup> in the positive column of an uncondensed discharge at low current density in CO<sub>2</sub> in the spectral region of 1900-6500 Å. They reported vibrational analysis of 35 bands of the Fourth Positive system



and rotational analysis of bands in the  $(0-v'')$  progression of the Ångström system ( $B^1\Sigma^+-A^1\Pi$ ). Band origins were obtained from the head measurements by using rotational constants of Schmid and Gerö [228] and the equations for the isotope effect. Low resolution precluded rotational analysis. The observed emission heads and origins of the isotopic A-X bands are given in table 5a.

McCulloh and Glockler [157b] believed that Read's [209] vibrational formula for the ground state was good for higher vibrational levels of  $C^{12}O^{16}$ . However, they assumed that his results were unduly influenced by errors in the short wavelength heads due to poor standards, and were not consistent with more recent infrared measurements. Hence, they fitted molecular constants for the normal and isotopic species which represented both Read's data and the infrared data of Lagemann et al., [139]. Slightly different constants have recently been calculated by Benedict [17] to fit the most recent infrared data as well as that for the A-X system. (See table 1.) Vibrational constants for the  $A^1\Pi$  state of  $C^{13}O^{16}$  were obtained [157] from Read's values for the normal isotope and the ratio of reduced masses. These are included in table 37.

High dispersion absorption spectra in the vacuum UV region have recently been photographed by Herzberg and Bass [100], and by Wilkinson and Tilford [276]. Several previously unobserved transitions are included in these studies, as well as bands of the  $v''=0$  progression of the A-X system up to  $v'=20$ , some bands of the  $v''=1$  progression, and also isotopic ( $C^{13}O^{16}$ ) A-X bands. The rotational analyses of these bands should provide definitive values of the rotational constants for the  $A^1\Pi$  state and precise vibrational quanta for this state [238a].

### 3.2. $B^1\Sigma^+-A^1\Pi$ Ångström System (6620-4120 Å) V

In 1875, Ångström and Thalén [see ref. 23]<sup>5</sup> observed a group of bands which became known as the Second Positive system of carbon. Fifty years later, analysis by Birge [23] showed that this system, which now bears the name of Ångström, arose from the allowed transition  $B^1\Sigma^+-A^1\Pi$  in carbon monoxide. This system has been a fertile ground for study because of the numerous

perturbations and suspected predissociations of the A state. These single-headed violet-degraded bands have also been observed for the isotope  $C^{13}O^{16}$ . Band heads and origins for the normal isotope are given in table 6; the isotopic ( $C^{13}O^{16}$ ) heads and origins are given in table 7. Rotational constants for the A and B states are given in tables 37 and 38, respectively.

Included among the original observations were weak bands at 4301 Å and 4581 Å which were later identified [23] as possibly belonging to the  $v'=2$  progression but which lay to wavelengths longer by more than 10 Å from the expected head positions. Johnson and Asundi [121] believed that these two bands belonged to another system. Only one other band with  $v'=2$  has been reported [255], but its identification is also uncertain since the B state is predissociated in levels  $v=0$  and 1 and presumably also in  $v=2$  at  $J=0$ . Another uncertain band, included as 0-6 by Birge [23], lies about 8 Å to longer wavelength from its expected position. A band observed by McLennan, Smith, and Peters [158] at 7246 Å, has not been reported elsewhere.

The first rotational quantum analysis of the 0-1, 0-2, 0-3, and 1-1 bands was made by Hulthén [108]. The 5610 Å band was the strongest observed, and incorrectly assigned as the 2-2 band. The correct vibrational quantum assignments are obtained by reducing Hulthén's  $v'$  by 2 and raising his  $v''$  by 1. Bands were also observed at 6078 Å, 6622 Å, and 4511 Å, but the structure of these could not be determined at that time.

Jasse [115] has given the rotational analysis for the 0-0 and 1-0 bands. His source was a high voltage electric discharge in  $CO_2$  at low pressure. Spectra were taken using a grating spectrograph of 3.25-m focal length; claimed precision is  $\pm 0.01$  Å. Hulthén [108] and Jasse [115] believed that the structure of the 0-0 and 1-0 bands was more complicated than the others of this system, having two additional branches. Birge [23] believed that the additional branches were probably spurious, while Johnson and Asundi [121] thought that the additional branches might originate from overlapping lines of  $C_2$ . Later, more accurate work by Rosenthal and Jenkins [216], and Schmid and Gerö [224] indicated that perturbations of  $A^1\Pi$ ,  $v=0$  were responsible for the confused structure of the 0-0 and 1-0 bands, resulting in displacement of the band heads from their expected positions.

<sup>5</sup> The data of Ångström and Thalén, as well as fine structure observations of Loos and head measurements by Watts and Wilkinson at the turn of the century, are also listed on p. 277 of Vol. V of Kayser's "Handbuch der Spectroscopie" [127a].



Johnson and Asundi [121] observed the 0-0 and 0-1 bands in the first order of a 21-ft focal length grating spectrograph (supplanting Jasse's measurements), and re-evaluated constants from Hulthén's data [108]. They also gave the first preliminary evaluation of the fine structure of the 0-4 and 0-5 bands, with details given by Asundi [5]. Referring to Jasse's analysis of the 4511 Å band, they considered the lines at 22172.3 cm<sup>-1</sup> and 22175.7 cm<sup>-1</sup> as extraneous to the band. Also they relabeled the line at 22171.8 cm<sup>-1</sup> as Q (1).

New measurements of fine structure of the 0-0, 0-1, 0-2, and 0-3 bands were made by Rosenthal and Jenkins [216] primarily to study perturbations, though the results are not completely reported in this work. (See sec. 4 on Perturbations.) It was found that Jasse, and Johnson and Asundi, did not correctly identify P and R lines above J=8 and Q lines above J=10 because of the large increase in  $\Lambda$ -doubling in the region of a perturbation. [See fig. 1 of ref. 216—the difference between the R and Q deviation curves.] Corrections to Jasse's assignments for several lines of the 0-0 band were made [216].

Coster and Brons [44] have given the rotational analysis of the 0-0, 0-1, 0-2, 0-3, and 0-4 bands in order to interpret the perturbations of the  $v=0, 1, 2, 3$ , and 4 levels of the A<sup>1</sup>Π state. Their discharge tube produced strong Ångström bands relative to the background of CO<sub>2</sub> bands. A value of  $\sigma_H(0-0)$  for the system of 22163.1 cm<sup>-1</sup> was obtained. (Herzberg [99] lists 22171.3 cm<sup>-1</sup>.) (See also sec. 4.1.)

Schmid and Gerö [224], from observations in second and third order using a 21-ft focal length grating spectrograph, gave improved measurements of the 0-0, 0-1, 1-0, and 1-1 bands. They obtained rotational constants for the B<sup>1</sup>Σ<sup>+</sup> state, and in addition,  $\Delta G(\frac{1}{2})=2082.07$  cm<sup>-1</sup>.

McCulloh [157a] and McCulloh and Glockler [157b] have reported the B-A Ångström bands due to C<sup>13</sup>O<sup>16</sup> emission in the positive column of an uncondensed discharge. The perturbations in the A<sup>1</sup>Π state are discussed at some length [157a] (see sec. 4.1) and details are given of the rotational analysis of the (0- $v''$ ) progression including the 0-1, 0-2, 0-3, 0-4, and 0-5 bands, and the 1-1 band. The C<sup>13</sup>O<sup>16</sup> 0-0 head lies to the red of the C<sup>12</sup>O<sup>16</sup> head. Overlapping by Third Positive bands ( $b^3\Sigma^+-a^3\Pi$ ) prevented rotational analysis of the 0-5 isotopic band. Overlapping of regular and isotopic bands required 0-1 branches to be

identified by a method analogous to the Loomis-Wood method (see pp. 191-2 in Herzberg [99]). For the 0-2 band there is almost complete superposition of the P and R branches such that  $R(J) \sim P(J+10)$ . The region beyond the tail of this band and also the 0-3 band is quite free from the general background caused by CO<sub>2</sub> bands. Therefore, these bands are useful when searching for predissociation in the B<sup>1</sup>Σ<sup>+</sup> state. The 0-4 band has a near superposition of the P and R branches.

Since only one  $v''$  progression was analyzed no information can be obtained about the vibrational constants of the B<sup>1</sup>Σ<sup>+</sup> state. (See Schmid and Gerö [224].) Values for  $\omega_e$  and  $\omega_e x_e$  for the isotopes C<sup>12</sup>O<sup>16</sup> and C<sup>13</sup>O<sup>16</sup> were approximated from a study of the isotope effect. The abnormally large value of  $\omega_e x_e$  for the B state suggests probable rapid convergence of levels. (See table 9a and footnote B<sup>1</sup>Σ<sup>+</sup> to table 1.)

Douglas and Møller [56] have photographed under high dispersion and analyzed the 0-1, 1-1, and 0-2 Ångström bands of C<sup>13</sup>O<sup>16</sup>. The bands were observed in a discharge through O<sub>2</sub> in a tube containing carbon, using a 21-ft focal length grating spectrograph with a reciprocal dispersion of 2.5 Å/mm. Their primary aim was an examination of reported predissociations of the C<sup>12</sup>O<sup>16</sup> and C<sup>13</sup>O<sup>16</sup> molecules (see sec. 5.2).

### 3.3. C<sup>1</sup>Σ<sup>+</sup>-A<sup>1</sup>Π Herzberg System (5710-3680 Å) V

The C<sup>1</sup>Σ<sup>+</sup>-A<sup>1</sup>Π Herzberg system, which partly overlaps the B-A Ångström system, consists of eight observed bands of the  $v'=0$  progression. No bands of this system have been observed with  $v' > 0$ , but the level C,  $v=1$  has recently been observed in the C-X system [276]. This suggests that there is either predissociation for  $v' > 1$ , or that a weak transition is expected from the Franck-Condon principle. Table 8 lists the heads and origins of the C-A system. The rotational constants for the C<sup>1</sup>Σ<sup>+</sup> state are given in table 39 and those for the A<sup>1</sup>Π state in table 37. Three bands of this system were first observed by Duffendack and Fox [58] who thought that the new bands belonged to the Ångström system because of their similar structure. Herzberg [97] observed eight bands and recognized these as belonging to a new system whose upper state he incorrectly designated as <sup>1</sup>Δ, by analogy with the Grottrian diagram for the Mg atom, as was done at that



time, following Mulliken. The new level lay between the  $B^1\Sigma^+$  and  $c^3\Sigma^+$  states.

Asundi [3] and Asundi and Johnson [121] observed the 0-1, 0-2, and 0-3 bands using a 21-ft focal length grating spectrograph in first order (wavelengths of the stronger lines are believed accurate to within 0.01 Å). Fine structure analysis showed the C state to be  $^1\Sigma$ . The Q branches were reported to be twice as strong as the P and R branches. For the  $A^1\Pi$  state there is no combination defect up to  $J=17$  (limit of observations here), indicating no appreciable  $\Lambda$ -doubling to high rotational quantum numbers. This is so for  $v''=1$  to 5, not for  $v''=0$ , for which combination defects occur from almost the lowest  $J$  values. [See ref. 224.] For the 0-0 band, only the calculated position is given, because of confusion with a band associated with the Third Positive system of CO ( $b^3\Sigma^+-a^3\Pi$ ). A band reported by Deslandres at 3893 Å which Wolter [277] did not observe is the 0-1 band of the C-A system.

Schmid and Gerö [224] observed the 0-0, 0-1, 0-2, 0-3, and 0-4 bands of the C-A system. From the rotational analysis, a value of  $B_0'$  of  $1.9422\text{ cm}^{-1}$  is obtained.

### 3.4. Hopfield-Birge Systems (1810-950 Å)

#### a. $B^1\Sigma^+$ , $C^1\Sigma^+$ , $E^1\Pi$ , $F(^1\Sigma^+)$ ,<sup>6</sup> $G(^1\Pi)$ , and $b^3\Sigma^+-X^1\Sigma^+$ .

In 1927, there appeared preliminary reports by Hopfield [103] and Hopfield and Birge [104] concerning the observation of seven band systems of CO both in emission and absorption. These transitions, all going to the ground state, included four whose upper states had not previously been recorded:  $a'^3\Sigma^+$ ,  $E^1\Pi$ ,  $F(^1\Sigma^+)$ , and  $G(^1\Pi)$ . Strongest absorption was to the highest lying levels (G, F, E, C). Full details of this work have never been published. Though several of the bands have since been observed elsewhere, only the  $a'^3\Sigma^+-X^1\Sigma^+$  system has been studied in detail [101] and will be discussed in section 3.4b. The other six transitions will be discussed together since the information about them is incomplete. Recently, however, high dispersion measurements have been made of the C-X, 0-0, and 1-0 bands, with  $v'=1$  observed for the first time [276].

The observed band heads of all six systems are given in table 9. The rotational constants for

the B, C, E, and  $b$  states are given, respectively in tables 38, 39, 40, and 43.

Read [209] observed several emission bands belonging to the B-X, C-X, and E-X transitions. Tanaka et al. [255] have seen some of these in absorption. Since the  $B_e$  values of the three excited states and the ground state are roughly the same, each band had the appearance of a doublet with a clear center under low resolution. Identification of the B-X, 2-0 band [255] is rather uncertain.

Recently, Tilford et al. [257] have identified the E state as  $^1\Pi$  from analysis of the violet-degraded 0-0, E-X band observed in absorption under high dispersion. This identification confirmed the theoretical prediction by Lefebvre-Brion et al. [147].  $\Lambda$ -doubling is about  $0.2-0.3\text{ cm}^{-1}$  for  $J\sim 20$ . (The  $J'=31$  level is perturbed.) The isotopic ( $C^{13}O^{16}$ ) 0-0 band was also analyzed. (Table 40.)

#### b. $a'^3\Sigma^+-X^1\Sigma^+$ (1810-1280 Å) R

Of all the band systems of CO which were originally observed by Hopfield and Birge [103, 104], only the dipole forbidden  $a'^3\Sigma^+-X^1\Sigma^+$  system has been extensively studied (by Herzberg and Hugo [101]). A long progression of bands has been observed up to  $v'=23$ , because of the large difference in internuclear separation between the two states involved. The upper state of this system is responsible for many perturbations of the  $A^1\Pi$ ,  $b^3\Sigma^+$ , and  $a^3\Pi$  states.

The band heads and origins of the  $a'-X$  system are given in table 10. The rotational constants for the X and  $a'$  states are given in tables 36 and 41, respectively.

Five bands of this system were first reported in absorption by Hopfield [103] with the one at 1696.9 Å designated as the 0-0 [104]. On the basis of an additional band found by Birge at 1731.2 Å, Estey [60] raised the vibrational quantum numbering in the  $a'$  state by one unit. The final assignment was deduced by Beer [16] from rotational perturbations in the  $b^3\Sigma^+-a^3\Pi$  Third Positive bands caused by the  $a'^3\Sigma^+$  state. This raised the quantum numbering of Hopfield and Birge by three units.

Recently, only two papers have dealt with this system [255, 101]. Tanaka, Jursa, and LeBlanc [255] have seen a long progression of these single-headed red-degraded bands under low resolution. Because a short absorption cell was used, the 0-0,

<sup>6</sup> It is not certain whether the F state is singlet or triplet. (See sec. 5.2f and table 1, footnote  $F(^1\Sigma^+)$ .)



1-0, and 2-0 bands were not observed. Herzberg and Hugo [101] observed this progression in absorption under high resolution in the region 1810-1280 Å. These forbidden bands were photographed in fourth and fifth orders of a 3-m focal length spectrograph, with reciprocal dispersions of 0.63 and 0.50 Å/mm, respectively.

In most of the bands [101] all four predicted branches, i.e.,  $^{\text{P}}\text{P}$ ,  $^{\text{P}}\text{Q}$ ,  $^{\text{R}}\text{Q}$ , and  $^{\text{R}}\text{R}$ , are observed and give the appearance of a doublet structure. The order of the sublevels of the  $a'^3\Sigma^+$  state is given as  $F_3 > F_1 > F_2$ .

The triplet splitting for the  $a'$  state is approximately constant for  $N > 5$ . [ $N$  was formerly called  $K$ .] A discussion of splittings for  $N < 5$  determined from other bands has been given by Gerö and Lörinczi [80]. The sum of splittings  $[2\lambda + \gamma]$  in corresponding members of the  $\text{P}$  and  $\text{R}$  branches is approximately constant, with small change of  $\lambda$  with  $v$ .

The rotational analysis results in the following formulas for the  $a'^3\Sigma^+$  state.

$$\sigma_0 = 55822.92 + 1230.651(v' + \frac{1}{2}) - 11.013(v' + \frac{1}{2})^2 \\ + 0.07378(v' + \frac{1}{2})^3 - 0.00115(v' + \frac{1}{2})^4 - 1081.59.$$

[Ref. 101 used 1081.74 for the ground state zero-point energy.]

$$B_v = 1.3453 - 0.01872(v + \frac{1}{2}) \\ + 0.000205(v + \frac{1}{2})^2 - 0.000051(v + \frac{1}{2})^3.$$

$D'$  was found to be about  $6.5 \times 10^{-6}$  and varied little from  $v' = 2$  to 23.

Only a perturbation in  $v' = 4$  at approximately  $J = 20$  was observed. The perturbing level is  $a^3\Pi$ ,  $v = 7$ , which has not been observed directly. Other perturbations which occur at high  $J$  (due to  $A^1\Pi$ ,  $a^3\Pi$ , and  $b^3\Sigma^+$ ) were not observed.

### 3.5. $a^3\Pi - X^1\Sigma^+$ Cameron System (5800-1770 Å) R

The  $a^3\Pi - X^1\Sigma^+$  transition consists of quintuple-headed bands, degraded to the red, covering the spectral region 5800-1770 Å. This forbidden intercombination system is overlapped by the  $B^2\Sigma^+ - X^2\Sigma^+$  First Negative bands of  $\text{CO}^+$  and the  $A - X$  system of  $\text{CO}$ . Recent high dispersion measurements by Rao [201] have provided an accurate value for the vibrational term interval  $\Delta G(\frac{1}{2})$  of the  $a^3\Pi$  state, but there is still lacking

an accurate value of  $\omega_e x_e$ . The observed band heads of the  $a^3\Pi - X^1\Sigma^+$  transition are listed in table 11, the origins in table 12. Rotational constants for the  $a^3\Pi$  state are given in table 42.

The band system bearing his name was first observed by Cameron [39] in an uncondensed discharge in neon with a trace of  $\text{CO}$ ; the following year it was seen in absorption by Hopfield and Birge [104]. Johnson [119] first identified these bands as due to transitions from the lower states of the Third Positive system to the ground state of  $\text{CO}$ , and assigned vibrational quantum numbers to the bands. Knauss and Cotton [131] also observed these bands at low pressure in an electrodeless discharge in  $\text{CO}$ .

Schmid and Gerö [232] observed the 4-8 band in a Geissler discharge in neon. Because of the coincidence of one head of this band and the intense red-degraded band reported by Kaplan [126], Schmid and Gerö believed that the level reported by Kaplan at  $38820 \text{ cm}^{-1}$  above  $X$  was not real. (See footnote (K) of table 1.)

Herman [94] reported selective excitation of the  $a - X$  system between 5000-2000 Å in a feeble discharge through  $\text{CO}$  at low temperature. Details are not given except for mention of the  $R_2$  (or  $Q_3$ ) heads of the 2-4 and 0-1 bands, at 2299 Å and 2159 Å, respectively. Comparison was made between the  $R_1$  and  $Q_2$  heads calculated from vibrational constants given by Johnson [119] and the faint low dispersion measurements of the night sky by a number of authors. The apparent agreement for a number of bands suggested to Herman the possible presence of the  $a - X$  system in the night sky. No other mention has been made of such correlation, and this agreement may be fortuitous.

Recently, Tanaka, Jursa, and LeBlanc [255] observed the  $v'' = 0$  progression in absorption under low resolution. Unlike other intercombination bands observed by them, this progression did not develop to high  $v'$  because of the Franck-Condon principle. [Note that ref. 255 lists vacuum wavelengths.]

The first high dispersion measurements of this system in absorption were those of Gerö, Herzberg, and Schmid [79] who gave the fine structure analysis of the 0-0 and 1-0 bands. This was the first example of analysis of the fine structure of a  $^3\Pi - ^1\Sigma$  transition. At least for large  $J$  the  $a^3\Pi$  state is close to Hund's case (b). Thus the quantum number  $N$  (formerly  $K$ ) has meaning for the rotational levels of both upper and lower

states in this transition. Two equivalent notations of the nine branches are given below:

$$\begin{array}{ccc} P_1 Q_1 R_1 & P_2 Q_2 R_2 & P_3 Q_3 R_3 \\ \text{OP } P^Q \text{ } ^Q R & P^P \text{ } ^Q Q \text{ } ^R R & Q^P \text{ } ^R Q \text{ } ^S R. \end{array}$$

The combination differences which give  $\Delta_1 F_1$  for the  $^3\Pi_0$  component of the  $a$  state do not coincide because of the very strong  $\Lambda$ -doubling of this component.

Gerö [76] used a Geissler discharge in neon with a trace of CO to excite (in emission) the  $a$ -X bands and, weakly, the  $\Lambda$ -X system. His close examination of the work by Gerö, Herzberg, and Schmid [79] revealed the use of two unreliable standard iron lines by Schmid and Gerö, which accounted for a discrepancy in their wavelengths. Gerö tabulated the rotational lines for the 0-0, 0-1, 0-2, 0-3, 1-3, 1-4, and 2-5 bands.

The most precise measurements of the  $a$ -X system have recently been made by Rao [201] who has given the rotational structure of the 0-0, 0-1, 0-2, 0-3, 1-3, and 1-4 bands seen in a Geissler discharge. A 21-ft focal length grating spectrograph was used in fourth and fifth order, with an iron arc providing a comparison spectrum. Errors in the combination differences for the 0-1 band by Gerö were clarified. Relative accuracy of the line measurements [201] is estimated to be  $\pm 0.003$  Å. (A reproduction of the 0-1 band is given in this reference.)

Rotational constants for the ground vibrational levels are close to those obtained from infrared measurements. Derived values for vibrational quanta for the ground state are the best available from electronic spectra of CO, but are inferior to those determined from infrared measurements. A value of  $1714.61 \text{ cm}^{-1}$  is obtained for the vibrational quantum  $\Delta G(\frac{1}{2})$  for the  $a$  state, confirming the value obtained previously [79].

### 3.6. $b^3\Sigma^+ - a^3\Pi$ Third Positive System (3820-2660 Å) V

The Third Positive system of carbon,<sup>7</sup> is composed of strong quintet-headed red-degraded bands whose structure is characterized by in-

tensity fluctuations due to perturbations of the upper state. Since the upper state is also pre-dissociated, only the  $v'=0$  and 1 progressions are known, the latter having been dubbed by Asundi [4] as 5B, and at that time thought to be a new system. This is understandable because violent perturbations of the  $b^3\Sigma^+$  state drastically alter the relative appearance of the two progressions. The original measurements by Deslandres [127a, 23] in 1888, of emission from discharge tubes containing carbon and oxygen, included some bands down to 2295 Å which turned out to be part of the 3A system ( $c^3\Sigma^+ - a^3\Pi$ ). The confirmed bands of the  $b$ - $a$  system extend from 3830-2660 Å. These vibrational heads are given in table 13. The rotational constants for the  $b^3\Sigma^+$  and  $a^3\Pi$  states are given in tables 43 and 42, respectively.

The first assignment of vibrational quantum numbers of the  $b^3\Sigma^+ - a^3\Pi$  system was made by Johnson [119] who identified the lower state with the upper state of the Cameron bands. Previously Wolter [277], using high resolution, had observed five bands of the  $v'=0$  progression. He mentioned the repeated rise and fall of intensity of the band lines and, in addition, reported the appearance of weak diffuse lines between the main lines. Later this system was observed by Duffendack and Fox [57, 58] who concluded that the group of bands labeled by Johnson [119] as the  $v'=4$  progression constituted the  $v'=0$  progression of a new system called 3A.

Asundi [4] observed the  $b$ - $a$  and 3A bands in an uncondensed discharge using a 21-ft focal length grating spectrograph. Outside the first (longest wavelength) head of these violet-degraded bands were faint lines which decreased in spacing toward higher wavelengths, formed a diffuse head, and continued further with a gradual increase in spacing. Asundi attributed the  $b$ - $a$  system to a  $^5\Sigma - ^5\Pi$  transition. Birge [23] disagreed and considered the transition as  $^3\Sigma - ^3\Pi$ , an assignment later firmly established by Dieke and Mauchly [54].

An extensive discussion of the spin splitting and  $\Lambda$ -doubling in these bands has been given by Dieke and Mauchly [54], who observed the 0-0 to 0-4 bands using dispersion of  $< 1 \text{ Å/mm}$ . Budó [32] has used these data to obtain rotational constants for the  $a^3\Pi$  state. Dieke and Mauchly's analysis established the transition as  $^3\Sigma^+ - ^3\Pi$ , with triplet (spin) separation of the  $b^3\Sigma$  state unnotice-

<sup>7</sup> Deslandres called the Third Positive group all [CO] bands in the UV region not belonging to the Fourth Positive group or Second Positive group (Ångström system). Some authors refer to the  $v'=0$  progression as Third Positive and call the  $v'=1$  progression "5B". The two together should be treated as a single system.



able for  $J < 20$ . The character of the  $\Lambda$ -doubling and the number of missing lines near the origin identified the lower state as regular  $^3\Pi$ , intermediate in coupling between case (a) and (b). From a study of the Zeeman effect on this band system Schmid [236] showed that, beginning at about  $N=22-23$ , the  $a^3\Pi$  state belonged to coupling case (b). The  $^3\Sigma$  terms have separation  $< 0.2 \text{ cm}^{-1}$  for  $N$  up to about 20. The terms in order of decreasing energy are anomalous:  $F_1 > F_2 > F_3$ .

The  $b-a$  bands represent a transition from case (a) to case (b); i.e., the  $^3\Pi$  state already departs from case (a) coupling for small  $N$ . Fifteen branches are allowed by the selection rules, of which fourteen were observed as well as the trace of a fifteenth. Nine strong bands indicate coupling close to case (b). The five head-forming branches are  $O_3$ ,  $O_2$ ,  $P_3$ ,  $P_2$ , and  $P_1$ .

Many large perturbations, mostly far from the heads, cause the characteristic intensity fluctuations first mentioned by Wolter [277], and cause clustering of lines toward the tails. At high pressure the tails show characteristic triplets especially in the 0-1 band at  $2977 \text{ \AA}$ . [See fig. 2 of ref. 54b.] In front of the heads (at shorter wavelength) are faint lines forming very weak series which do not form heads. Dieke and Mauchly showed that these are due to transitions with  $\Delta N=3$ . It is not certain whether these were the lines found by Asundi [4].

Schmid and Gerö [226] have reported clustering of Q-branch lines in the  $v'=1$  progression caused by perturbations of the  $b^3\Sigma^+$  state. The positions of these perturbations converge to a limit which seems to coincide with the predissociation limit common to the  $v'=0$  progression and to the B state. Gerö's [75] report on the fine structure of the 1-0 band has confirmed the predissociation of the upper state between  $N=42$  and 43.

Rotational analysis of the 0-4, 0-5, 1-4, and 1-5 bands has been given by Beer [16]. Perturbation of  $v''=4$  was observed and assigned to  $a^3\Sigma^+$ .

The molecular constants for the  $b$  state are somewhat uncertain. Schmid and Gerö's values [226] are quoted by Herzberg [99]; i.e.,  $B_e=2.075$ ,  $\alpha_e=0.033$ ,  $B_1=2.025 \text{ cm}^{-1}$ . Stepanov [246] has criticized the way these values were determined and his values are listed in table 1, i.e.,  $B_e=1.986$ ,  $\alpha_e=0.042 \text{ cm}^{-1}$ . For  $\Delta G(\frac{1}{2})$  from measurement of  $b-a$  bands Schmid and Gerö [226] obtained  $2109 \text{ cm}^{-1}$ . Jevons [117] had previously listed  $2214 \text{ cm}^{-1}$  from the data of Asundi [4]. Gerö's [75]

estimate of  $2198 \text{ cm}^{-1}$  is the value listed in Herzberg. This value was obtained by extrapolation to  $N=0$  of the difference between the Q lines of the 1-0 and 0-0 bands. Stepanov [246] criticized this value and used a different method to obtain  $2188 \text{ cm}^{-1}$  which is the value quoted in table 1. These constants should be accepted as only tentative and renewed attempt should be made to determine them more precisely.

### 3.7. $c^3\Sigma^+ - a^3\Pi$ 3A System (2710-2300 $\text{\AA}$ ) V

Duffendack and Fox [57, 58] measured the excitation potentials of several band systems of CO, including five bands of a new system appearing at about 11.1 eV. These bands had the same final state as the Third Positive system, [ $b^3\Sigma^+ - a^3\Pi$ ], but originated from a state about 1.02 eV above the  $b$  state. The new system was labeled 3A, and bands were identified as the  $(0-v'')$  progression for  $v''=0$  to 4. The violet-shaded bands were multiple headed.

Faint, five-headed bands of this system were produced by Asundi [4] in an uncondensed discharge in CO. The first and fifth heads are strongest; hence, under low dispersion the bands appear double-headed. The electronic multiplet separation of the measured extreme heads is largest for the 0-0 band, and, as with the Third Positive bands, decreases for higher vibrational states.

An equation for the heads is given by Asundi [4]

$$\begin{array}{rcl} \sigma & = & 43542.5 \\ & \text{---} & \\ & \text{---} & - (1726.5v'' - 14.4v''^2). \\ & \text{---} & \\ & & 43612.4 \end{array}$$

Schmid and Gerö [230] excited intense 3A bands in a discharge through neon in a Geissler tube with carbon electrodes and gave a preliminary report on the 0-1, 0-2, and 0-3 bands. They detected 12 branches in each band, the remaining branches being weak and overlapped. Six branches are head forming:  $O_3$ ,  $O_2$ ,  $P_3$ ,  $P_2$ ,  $P_1$ , and  $Q_1$  in order of diminishing wavelengths. The nature of the branches shows the upper state to be  $^3\Sigma$ , with negligible multiplet splitting.

Gerö [77] reported the detailed rotational analysis of the 0-1, 0-2, and 0-3 bands observed at a dispersion of  $1.3 \text{ \AA/mm}$ . A predissociation in this system is observed at about 11.5 eV. The



combination differences for the  $a^3\Pi$  state agree well with those of the  $b-a$  system which confirms the order of the bands and band lines and the  $\Sigma^+$  character of the  $c$  state. A value of  $B_0=1.9563\text{ cm}^{-1}$  is derived.

The last observed lines of the  $R_3$  branch of the 0-1 3A band lies in the region of a spurious pre-dissociation in the P branch of the red-degraded 9-18 Fourth Positive band. The rotational analysis of the 3A bands shows that the apparent increase of intensity of some of the Fourth Positive band lines [74] in this region are caused by superposition of certain 3A lines.

The band heads of the 3A system are given in table 14, and the rotational constant  $B_0=1.9563$  is included in tables 1 and 44.

### 3.8. $a'^3\Sigma^+-a^3\Pi$ Asundi System (8590-3900 Å) R

The multiheaded bands of the  $a'-a$  system have been studied mostly under low dispersion, with only a few rotational constants determined for the  $a'$  state from their analysis. More accurate values for these constants are available from analysis of another system [101]. Two bands which were believed to originate from levels with  $v\geq 30$  in the  $a'$  state [78] most probably belong to a new  $^3\Sigma^+$  state whose term value is uncertain. The observed vibrational heads of the  $a'^3\Sigma^+-a^3\Pi$  system are given in table 15. Rotational constants for the  $a'$  and  $a$  states are given in tables 41 and 42, respectively.

McLennan, Smith, and Peters [158], using low dispersion, observed a large number of red-degraded bands produced in a long, high-voltage discharge tube filled with CO. (A number of these bands which have not been identified are listed in table 35c.) Shortly thereafter Asundi [4] discovered 16 red-degraded bands of the system bearing his name which were overlapped by the  $d^3\Delta_i-a^3\Pi$  Triplet bands below 5700 Å. The  $a'-a$  bands appear double headed under low dispersion, but have five heads when seen under higher dispersion.

By assuming the band at 8592 Å was the 1-0, Asundi assigned vibrational quantum numbers. A new term of the  $a'-X$  system found by Estey (at  $57763\text{ cm}^{-1}$ ) [60] required the  $v'$  assignments to be raised by unity. The correct  $v'$  numbering was finally established by Beer [16] who indicated that a perturbation of the  $a^3\Pi$ ,  $v=4$  term is caused by  $a'^3\Sigma^+$ ,  $v=0$ , which required Asundi's original  $v'$  values to be raised by three. [See ref. 80.]

Knowledge of the  $a'-a$  system was considerably extended by the observations of Garg [68]. Numerous impurity CO bands were observed during the preparation of a discharge tube for the study of emission from  $\text{SnBr}_4$  vapor. The band positions are assumed accurate to  $\pm 4\text{ cm}^{-1}$ . Garg's formula, fitted to both his and Asundi's measurements, is given below with  $v'$  numbering of Gerö and Lörinczi [80].

$$\sigma_R = 7213.8 + 1218.1(v' + \frac{1}{2}) - 9.5(v' + \frac{1}{2})^2 \\ - 1740.9(v'' + \frac{1}{2}) + 14.4(v'' + \frac{1}{2})^2.$$

A few bands which cannot be fitted in this system are listed separately, and assumed members of an unknown system. (See table 35c.) Garg's table of band heads shows increasing deviation of observed-calculated differences for  $v'\geq 16$ . It is not presently known whether such deviations arise from perturbations, or are due to incorrect identification of particular subheads used to calculate these deviations.

Schmid and Gerö [233], in a note, reported several red-degraded bands near the heads of the violet degraded  $b^3\Sigma^+-a^3\Pi$  Third Positive system. An especially strong band is observed at 2670 Å in a Geissler tube discharge in neon with carbon electrodes, in the presence of very little oxygen. In another brief communication Schmid and Gerö [231] ascribed the bands as due to a  $^3\Sigma^- - a^3\Pi$  transition. The upper state, provisionally designated as  $f^3\Sigma$ , was suspected as including high vibrational levels of the  $a'^3\Sigma$  state. This was supposedly confirmed by Gerö [78], but Herzberg [99] believed the bands belonged to another state and confirmed this indirectly in later work [101].

Gerö and Lörinczi [80] have given the rotational analysis of the 8-0, 9-0, 10-1, and 11-1 bands which were observed in emission. The 8-0 band is weaker than the rest and its branches are not given. The high dispersion measurements are relatively free of overlap by other systems, and allowed measurement of lines to a few hundredths  $\text{cm}^{-1}$ . Rotational constants for  $a'^3\Sigma^+$ ,  $v=8$  to 11 are obtained. A long extrapolation of these [which is usually dangerous] gives the molecular constants  $B_v=1.331\text{ cm}^{-1}$  and  $\alpha_v=0.016\text{ cm}^{-1}$  which are the values listed by Herzberg [99]. The rotational constants for the upper levels are smaller by only about 1.5 percent than those given previously from perturbation data [234a]. From more recent work on the  $a'^3\Sigma^+-X^1\Sigma^+$  system, Herzberg and Hugo [101] obtained im-



proved constants  $B_e=1.3453$ ,  $\alpha_e=0.01872$ , and  $\gamma_e=0.000205$  cm<sup>-1</sup>.

The spin splitting of the  $^3\Sigma$  state is rather large. For the 10-1 band, the splitting is constant above  $N=5$ . The splitting (in cm<sup>-1</sup>) is almost the same for each of the bands as is seen below:

	$F_3-F_2$	$F_1-F_2$
9-0	1.40	0.97
10-1	1.37	.98
11-1	1.40	1.01

(See ref. 80 for a correction to table 2 of ref. 78).

### 3.9. $d^3\Delta_i-a^3\Pi$ Triplet System (7500-3770 Å) R

The  $d^3\Delta_i-a^3\Pi$  Triplet bands of CO are a fairly extensive system of moderate intensity spanning the visible region of the spectrum. Merton and Johnson [160], in studying the effect which the presence of helium had on band spectra of carbon, first isolated two progressions of this system. These red-degraded bands were later assigned to CO when Birge [21] identified the lower state as  $a^3\Pi$  and gave a vibrational analysis. Extension of the system to longer wavelengths by Asundi [7] required an increase in  $v'$  assignments by two units. Carroll [40] has reported that the numbering for the  $d$  state vibrational levels should be increased by one more unit. Only recently the upper state was identified [40] as  $^3\Delta_i$ , following a prediction of the  $\Delta$  character by Mulliken [171]. Kovács [137] has explained the inverted character of this state which is not expected in first order from a  $\pi^3\pi$  configuration. Band heads are given in table 16; origins in table 17. Rotational constants for the  $d^3\Delta_i$  state are given in table 45.

Asundi [7], from study of three bands at longer wavelength, concluded that the band at 15540 cm<sup>-1</sup> was not 0-0, and that the  $v'$  assignments as listed by Gerö and Szabó [82] should be raised by two. The bands were found to have five heads. He was not certain whether his band at 13303 cm<sup>-1</sup> was definitely 0-0, and speculated that the system origin might be at longer wavelength than his range of observation.

Herman and Herman [92, 96] observed emission of  $d^3\Delta_i-a^3\Pi$  Triplet bands in a low current discharge in neon with traces of CO at liquid nitrogen temperatures. Under these conditions bands with high  $v'$  predominate. Bands with  $v'=6$  [add 3 to

the  $v'$  assignments of Herman and Herman] are relatively weak compared to intense bands for  $v'=5$  and 7. The bands with  $v'=12$  are relatively weak and those with  $v'=13$  are enhanced compared with those for  $v'=14$ . A new system of triple-headed bands was also observed and is discussed further in section 3.14.

Recently Sato [219] used a 33.3 MHz discharge to produce more than 40 rather intense bands of the  $d-a$  system. His low dispersion measurements included almost all bands observed previously as well as the 12-0, 5-1, 6-1, 6-2, 7-2, 17-2, 18-2, 19-3, 20-3, 17-4, 21-4, and 22-4. His  $v'$  values should be raised by three units. Uncertainty in the shorter wavelengths is about 0.5 Å; uncertainty in the longer wavelengths is about 1 Å.

For a number of bands, Sato's measurements differ considerably from those of several other authors. This raises questions about quantum assignments for other bands observed by Sato. This system should be reexamined to clarify the band identification.

Gerö and Szabó [82] have given the rotational analysis for the 3-0, 4-0, 6-0, 7-0, and 9-1 bands. (Their original  $v'$  assignments are too low by three units.) The fine structure was interpreted as showing the transition to be  $d^3\Pi_i \rightarrow a^3\Pi_r$ . The upper state was later correctly identified as  $^3\Delta_i$  by Carroll [40]. Gerö and Szabó observed these bands in a Geissler discharge in neon with traces of oxygen. The bands also appeared under favorable conditions without the presence of a noble gas. In the latter case there was overlap with the Ångström and Asundi bands,  $B^1\Sigma^+-A^1\Pi$  and  $a'^3\Sigma^+-a^3\Pi$ , respectively.

Twenty-seven branches were observed, doubled because of  $\Lambda$ -doubling. It was established that  $\Lambda$ -doubling of the band lines agreed with that of the  $a^3\Pi$  state (from the  $b^3\Sigma^+-a^3\Pi$  Third Positive bands) and that the  $\Lambda$ -doubling of the  $d$  term was immeasurably small. Rotational constants, coupling constants  $A$ , and  $Y=\frac{A}{B}$  for the  $d$  state were calculated from combination differences and the  $F_1-F_3$  differences based on Budó's formulas [32b] for a transition between two triplet  $\Pi$  states in intermediate coupling between cases (a) and (b). Corrected values are given by Carroll [40]. The origins of four bands were calculated by Gerö and Szabó [82] using an average of nine Q-branches whose series were extrapolated to  $J=0$ . Based on these zero positions the vibrational term differ-



ences were calculated for the  $d$  term. A formula fitted to the correctly numbered  $v''=0$  progression is given here as  $\sigma_0(v'-0)=11598.50+1152.58(v'+\frac{1}{2})-7.2812(v'+\frac{1}{2})^2-0.1125(v'+\frac{1}{2})^3$  which gives  $\sigma_0(0-0)$  at  $12,172.96\text{ cm}^{-1}$ .

Carroll [40] has given the rotational analysis of the 3-0 band based on high resolution measurements and has shown the upper state to be  $^3\Delta_i$  in agreement with electron-configuration theory and not  $^3\Pi_i$  as indicated by Gerö and Szabó [82]. Mulliken [171], on the basis of LCAO-MO calculations for the lower excited states of CO, had predicted a  $^3\Delta$  state at about the observed energy. Carroll's work was initiated to clarify the nature of the upper state, and is the first reported example of a  $^3\Delta\rightarrow^3\Pi$  transition with nonnegligible (multiplet) splitting. The source used was a discharge through flowing helium with a trace of CO. Of the several bands observed, the one at  $6401\text{ Å}$  was analyzed because it extended to high  $J$  and was relatively free from overlapping by other structure. All 27 branches were observed, although sometimes the  $\Lambda$ -doublets were not resolved for lower  $J$ . There are nine subbands, each composed of three doublet branches, although the branches going to  $^3\Pi_2$  of the lower state are only resolved at higher  $J$ . The  $^3\Delta_i$  state is case (a) at low  $J$  and changes to case (b) at higher  $J$  (See fig. 1 of ref. 40.) The  $a$  state also is case (a) at lower  $J$ .

Carroll's measurements agree well with the previous ones of Gerö and Szabó for the 3-0 band, with several differences in classification of lines. In addition, some of Gerö's assignments at low  $J$  seem to be spurious. Detailed evidence is given for the assignment of the upper state as  $^3\Delta_i$ , including confirming evidence from the intensity distribution in the fine structure. At that time there had been no theoretical calculation of branch intensities for a  $^3\Delta-^3\Pi$  transition, but such formulae have since been obtained by Kovács [136], for various coupling conditions of the states involved.

The energy formulas of Budó [32b], although good to a first approximation, do not give the rotational term values of the  $^3\Delta_i$  state with sufficient accuracy [40]. This deviation between the observed and calculated energies, which affects most strongly the middle multiplet component, is explained by Kovács [137] as arising from the joint effect of two factors normally neglected: (1) spin-

spin interaction, and (2) spin-orbit interaction by mixing with a nearby state, probably  $^1\Delta$ . No  $^1\Delta$  state has been observed directly, though a possible perturbation of  $A^1\Pi$ ,  $v=0$  by such state has been mentioned by Rosenthal and Jenkins [216].

### 3.10. $e^3\Sigma^- - X^1\Sigma^+$ Transition (1540-1240 Å) R

A forbidden electronic transition,  $e^3\Sigma^- - X^1\Sigma^+$ , has been observed by Herzberg and Hugo [101] in the vacuum UV region. They observed a long  $v''=0$  progression of weak, red-degraded single headed bands in absorption by use of a 3-m focal length vacuum spectrograph, at dispersions of 0.63 and 0.50 Å/mm. These bands have also been seen under low dispersion by Tanaka, Jursa, and LeBlanc [255]. The existence of the  $e^3\Sigma^-$  state had previously been inferred from perturbations in the B-A bands [101] and in the A-X bands [234b]. Many  $e-X$  bands are overlapped by bands of the A-X,  $a'-X$ , and  $d-X$  transitions [101, 255]. Table 18 lists the band heads and origins of the  $e-X$  system. Rotational constants for the  $e^3\Sigma^-$  state are given in table 46.

The observed progression began with a relatively strong band at  $64816\text{ cm}^{-1}$  which was labeled 0-0. Since then, two bands have been found at longer wavelength [100] which suggested raising the  $v'$  values reported earlier by two units. Fine structure analysis of these bands [238a] showed conclusively that the  $v'$  values of Herzberg and Hugo should be raised by unity.

Fine structure has been determined for the 5-0, 6-0, 7-0, 9-0, 11-0, and 13-0 bands, which resemble a  $^1\Pi-^1\Sigma$  transition, but with the Q branch doubled. Electric dipole selection rules ( $+\leftrightarrow-$ ,  $\Delta J=0, \pm 1$ ) indicate five branches are expected for a  $^3\Sigma^- - ^1\Sigma^+$  transition which is possible because of spin-orbit interaction: Q,  $^oP$ ,  $^oR$ ,  $^sR$ , and  $^oP$ . Recall that the selection rule prohibiting  $\Sigma^- - \Sigma^+$  transition does not apply to intercombinations [220]. In the  $^3\Sigma$  state,  $F_1$  and  $F_3$  sublevels ( $J=N\pm 1$ ) lie close together, with the  $^oP$  and  $^oR$  branches nearly coincident and separated from the Q branch. These conclusions are confirmed in the observation of the 8-0 band, for in most other bands, the three Q-form branches are not resolved. Herzberg and Hugo [101] have made no quantitative intensity measurements, but have compared the intensity factors of the Q-form branches as given by Schlapp [220].

The formulas fitted to their data (with revised quantum numbering) are given below:

$$B_v' = 1.2848 - 0.0181(v' + \frac{1}{2}) + 0.0001(v' + \frac{1}{2})^2$$

$$\sigma_0(v' - 0) = 64236.32 + 1113.167(v' + \frac{1}{2}) - 9.596 \\ (v' + \frac{1}{2})^2 + 0.00587(v' + \frac{1}{2})^3 - 1081.59$$

where the last number is the zero-point energy of the X state.

### 3.11. $C^1\Sigma^+ - a'^3\Sigma^+$ Knauss System (3250–2930 Å) V

Knauss [130], in an abstract, reported four violet-degraded bands obtained in an electrodeless discharge through CO. These were identified as belonging to  $C^1\Sigma^+ - a'^3\Sigma^+$  (table 19). More recent data on the  $a'$  state indicates that the  $v''$  values should be two units larger than those quoted by Knauss.

### 3.12. Kaplan Bands (2750–2520 Å) V

A. In a brief note Kaplan [125] reported three bands forming a new system when a trace of CO was excited in a long atomic hydrogen tube. The bands resembled the 3A system ( $e^3\Sigma^+ \rightarrow a^3\Pi$ ), each having six heads and degraded to the violet. The transition is attributed to  $K - a^3\Pi$ , the upper state being tentatively identified by Kaplan as a metastable quintet. These bands do not appear under ordinary conditions. (Schmid and Gerö [227] have speculated that these bands may be the  $v'=2$  progression of the  $b-a$  Third Positive system.) Assuming  $v'=0$  yields  $T_0 = 89889 \text{ cm}^{-1}$  (above X) for the upper level. The bands are listed in table 20.

B. Kaplan [126], in an abstract, reported an intense single band at 2575 Å, similar in structure to the  $A^1\Pi - X^1\Sigma^+$  Fourth Positive bands, hence degraded to the red. This band was discovered in quenching mercury resonance radiation by CO. Schmid and Gerö [232] indicated that this band coincided with a head of the 4–8,  $a^3\Pi - X^1\Sigma$  Cameron band (which according to Cameron himself was the most intense band of the system) and disputed the reality of the 2575 Å band as due to a new CO transition. Schmid and Gerö [234b] indicated that the disputed band is strongly overlapped by the 0–0 band of  $A^1\Pi - X^1\Sigma$  of CS in a variety of discharges and this level of CO

would appear to be illusory, and should not be included among CO terms.

### 3.13. $d^3\Delta_i - X^1\Sigma^+$ Transition (1620–1230 Å) R

In a study of absorption spectra of CO in the vacuum ultraviolet region, Tanaka, Jursa, and LeBlanc [255] observed a weak progression of single headed bands extending to high  $v'$ , assumed to correspond to the R heads of the transition. At that time the upper state was thought to be  $^3\Pi$ , and is now known to be  $^3\Delta_i$ . The  $v'$  values given in reference [255] need to be raised by one unit. (See the discussion of  $v'$  numbering in the Triplet  $d-a$  bands.) The band heads are listed in table 21.

### 3.14. $e^3\Sigma^- - a^3\Pi$ Herman System (5430–4270 Å) R

In a study of the  $d^3\Delta_i - a^3\Pi$  Triplet bands, Herman and Herman [96, 92] reported a new progression of emission bands similar in appearance [red-degraded], but having more closely spaced heads. These bands were attributed to a new system whose upper state was assumed to be  $d$  (then thought to be  $^3\Pi$ ) and the lower state assumed to be a new triplet level lying near  $a^3\Pi$ . Long ago Merton and Johnson [160] had observed some of these bands, but had not identified them. By using the absorption data of Herzberg and Hugo [101] and that of Tanaka, Jursa, and LeBlanc [255], Barrow [12] has shown that the Herman bands constitute the  $(v'-0)$  progression of the  $e^3\Sigma^- - a^3\Pi$  transition (table 22). The  $v'$  numbering of Barrow should be raised by unity [238a]. "It is anticipated that other bands of the Herman system will probably be found at longer wavelengths" [12].

### 3.15. $E_0^1\Sigma^+ - X^1\Sigma^+$ , $^1\Pi - X^1\Sigma^+$ , and Several Unidentified Transitions in the Vacuum UV Region (1180–930 Å)

In a discharge in CO at several mm pressure Tschulanowsky and Gassilewitsch [261] excited a number of weak bands in the region 1200–930 Å. Included among the bands is the single band of the E–X Hopfield-Birge transition. Most of the bands are unclassified and are listed as such in in table 35d. Several bands belong to a new system labeled  $E_0^1\Sigma - X^1\Sigma$  by Tschulanowsky [260] (see table 23). The  $(0-v'')$  progression of



red-degraded bands has a system origin at 90866.2  $\text{cm}^{-1}$ . The relative error is estimated to be 0.01 to 0.03 Å (or less than 3  $\text{cm}^{-1}$ ). The constant  $v'$  value is uncertain and was assumed to be zero. A formula is obtained which fits the heads:  $\sigma(0-v'')=90869.45-(2171.19 v''-14.28 v''^2)$ , giving constants which deviate considerably from those of the ground state. On this basis the reality of the  $E_0$  state is questioned.

For the most intense band at 1183.62 Å, a value of  $(B'-B'')=-0.6882 \text{ cm}^{-1}$  was obtained. Both branches observed were of comparable intensity. By using  $B''=1.8702 \text{ cm}^{-1}$  one obtains  $B'_0=1.182 \text{ cm}^{-1}$  (which is given in table 47),  $r'\sim 1.44 \text{ Å}$ , and  $\sigma_0(0-0)=90866.2 \text{ cm}^{-1}=1100.52 \text{ Å}$ . This suggests a very weak bond and a state with a flat potential curve. The  $B'_0$  is not considered accurate, but a rough confirmation of its magnitude was obtained from a partial analysis of two other bands. It is estimated that  $\omega_e$  for the  $E_0$  state  $<1000 \text{ cm}^{-1}$  (comparable to the  $b'^1\Sigma_u^+$  state of  $N_2$ ).

Tschulanovsky [260] has also reported a group of weak, overlapped red-degraded bands in the region of 1100–1000 Å which form a new  $^1\Pi-X$  system. The measurements are not very reliable. The best resolved band is the 0–1 at 1034.65 Å, which shows the P and R branches longer and more intense than the Q. A value of  $(B'-B'')=-0.7659 \text{ cm}^{-1}$  gives  $B'=1.139$  (see table 48) using  $B'_0=1.9052 \text{ cm}^{-1}$ . The calculated origin of the  $^1\Pi-X^1\Sigma^+$  system is at 98836  $\text{cm}^{-1}$ . The bands are given in table 24.

### 3.16. $f^3\Sigma^+-a^3\Pi$ Transition (2980–2670 Å) R

Schmid and Gerö [233] have observed several red-degraded bands (table 25) near the heads of the 1–0 and 0–1 violet-degraded  $b-a$  Third Positive system. An especially strong band appeared at 2670 Å in a Geissler tube discharge (between carbon electrodes) in neon, in the presence of very little oxygen. Schmid and Gerö [231] ascribed the bands to a  $^3\Sigma^+-a^3\Pi$  transition. The upper state, provisionally designated as  $f^3\Sigma^+$ , was suspected as including high vibrational levels of the  $a'^3\Sigma^+$  state. The supposed confirmation of this assignment by Gerö [78] was not fully accepted by Garg [68], Gaydon [69], and Herzberg [99]. The work of Herzberg and Hugo [101] on the  $a'-X$  transition implies that the Schmid and Gerö bands do not belong to the  $a'-a$  system.

Gerö [78] has given the rotational analysis of the two  $f-a$  bands which he labeled the 30–1 and

34–0 bands of the  $a'-a$  transition. Stepanov [247] labeled these with  $v'$  increased by one unit. For the  $a^3\Pi$  state in coupling case (a), 27 branches are allowed. Relatively large triplet splitting of the  $f^3\Sigma^+$  terms is indicated since most of these were observed. Spin splitting ( $\text{cm}^{-1}$ ) is found to be roughly independent of  $N$ :

	2979.9 Å	2669.7 Å
$F_1-F_2$	0.80	1.05
$F_3-F_2$	1.08	1.43

(See p. 461 of ref. 80 for a correction to Gerö's paper.) Perturbations are indicated near  $N=0$  of the  $f$  state, presumably caused by  $b^3\Sigma^+$ . Rotational constants for the  $f$  state are estimated using the unperturbed levels with higher  $N$  (table 49). The unperturbed term values of the upper state vibrational levels are estimated to lie at 83744 and 85969  $\text{cm}^{-1}$  (36  $\text{cm}^{-1}$  has been added to Gerö's figures [78] to account for revised  $a-X$  data). Stepanov [247] obtained values close to those of Gerö which, when revised, would give 83755 and 85960  $\text{cm}^{-1}$ .

The small  $B$  values (0.83 and 0.75  $\text{cm}^{-1}$ ) for the  $f$  state terms imply either high vibrational quantum states or a weakly bound electronic state with large  $r_e$ . This rather peculiar structure should be re-examined experimentally.

### 3.17. $I^1\Sigma^--X^1\Sigma^+$ Transition (1520–1460 Å) R

Three red-degraded bands consisting of a single Q branch have been observed in absorption [100] under high dispersion and correspond to the forbidden  $I^1\Sigma^--X^1\Sigma^+$  transition. Rough measurements of wavelength indicate a vibrational frequency near 1000  $\text{cm}^{-1}$ . Full details are not yet available.

The  $I^1\Sigma^-$  state had previously been known only from its perturbation of the  $A^1\Pi$  state (see table 54a). A number of the perturbations caused by the  $I$  state have been assembled by Schmid and Gerö [225, 234b], Kovács [135], and from more recent measurements of the  $B^1\Sigma^+-A^1\Pi$  Ångström bands by Deutsch and Barrow [52]. These authors all assumed that the  $v=0$  level of the  $I$  state perturbed  $A^1\Pi$ ,  $v=1$ . The above measurements [100] suggest that it is probably the  $v=1$  level of the  $I$  state that is involved.

Earlier estimates of  $B_e$  of about 1.48  $\text{cm}^{-1}$

implied a value of  $\alpha_e$  of about 0.046 which is rather large for any state of CO. [But the  $b$  state has an  $\alpha_e$  of 0.042 or 0.033]. Deutsch and Barrow [52], by assuming the value of  $\alpha_e$  as 0.018  $\text{cm}^{-1}$  which is similar to other states having the configuration  $\pi^3\sigma^2\pi^*$ , have derived a term value for the electronic state and vibrational and rotational constants. These values, adjusted by raising their  $v$  values by one unit, are listed in table 1. The state lies at about 65630  $\text{cm}^{-1}$ .

A summary of the observed  $1^1\Sigma^-$  perturbations is given below:

$v(A^1\Pi)$	$v(1^1\Sigma^-)$	$J$	$T_{v=0}(1^1\Sigma^-)$
1	1	35	66678
2	2	7	67704
3	4	40	69699
4	5	23	70667
5	7	44	72547
6	8	30	73459
8	11	34	76080
9	12	13	76915
11	15	16.5	79360 (ref. 234b)

All data points but the last are those of Deutsch and Barrow [52]. (Term values are in  $\text{cm}^{-1}$ ).

### 3.18. Unidentified Bands

A. Herman [95] has found two new groups of triplets attributed to CO, which had earlier been found in comets, using as source a 90-cm long discharge tube with carbon electrodes, filled with xenon to a pressure of 10 mm of Hg. The two groups of triplets have different structure. Group (a) is red-degraded as is evident even under low dispersion, group (b) has branches which look like lines (as may be seen in fig. 1 of ref. 95). (See table 35e.) No further details are given.

B. Johnson [118] has observed five double-headed red-degraded bands together with other bands belonging to  $\text{CO}^+$ . They are described as similar to the  $A^2\Pi_i-X^2\Sigma^+$  comet tail bands, but with different separation, and each head is not a close doublet. These are given in table 35j.

C. Barbier [11] has observed 26 bands in the night sky (5200–3200 Å) which were attributed to a new transition to the ground state of CO. The upper state is at 21962  $\text{cm}^{-1}$ , with  $\omega'_e=2610 \text{ cm}^{-1}$  and  $\omega_e x'_e=23.8 \text{ cm}^{-1}$ . Pearse [186] has questioned

this assignment since  $\omega'_e \gg \omega''_e$ . The identity of these bands as belonging to CO is highly questionable.

D. Huffman et al. [107] have observed numerous unclassified bands (table 35f) in their study of the absorption coefficients of CO between 1000–600 Å. These are in addition to many previously observed Rydberg states, and other bands seen by Tanaka [254] and Henning [91].

### 3.19. P, Q, R, S, T– $X^1\Sigma^+$ Tanaka Systems (800–630 Å) R

Among numerous absorption bands which make up several Rydberg series, Tanaka [254] observed five new progressions with roughly constant frequency difference. These progressions, representing transitions from the ground state to the P, Q, R, S, and T states, are listed in table 26. Weissler et al. [270] have observed the P state and the S (or T) state by photoionization. Kaneko [124] has observed these in electron impact.

### 3.20. Rydberg Series (940–630 Å)

Rydberg series are of special interest because their convergence limits give the ionization potentials of the molecule. Absorption measurements have been reported of such series whose convergence limits coincide with the three spectroscopically observed states of  $\text{CO}^+$ . Recently, series converging to what appears to be a new state of  $\text{CO}^+$  have been observed [42]. The series converging to the B, A, and X states of  $\text{CO}^+$  are given, respectively, in tables 27, 28, and 29. The observed Rydberg states of CO are most likely  $1^1\Sigma^+$  or  $1^1\Pi$ , though no fine structure analyses have yet confirmed the tentative assignments. Details of their possible electronic structure have been given in section 2. The spectroscopic observations are reviewed below.

A number of bands which reportedly belong to Rydberg series have been observed under low resolution by Henning [91] and Anand [2]. Uncertainties in the measurements make their series limits unreliable. Henning [91] observed 22 bands in absorption, including 10 which formed two progressions having nearly equal frequency difference ( $\approx 1550 \text{ cm}^{-1}$ ). The positions of these broad bands are uncertain by  $\pm 50 \text{ cm}^{-1}$ . Anand [2] observed six headless bands in emission, and in the same spectral region (900–850 Å) saw a series of four bands in absorption. Both series were



fitted by a Rydberg formula with a common limit of 14.5 eV, which falls between the term values of vibrational levels of  $\text{CO}^+$ ,  $\text{X}^2\Sigma^+$ . Henning's bands are listed in table 35g. Anand's are listed in table 35h.

Tanaka [254], using a 3-m focal length grazing incidence vacuum spectrograph, observed numerous bands, including most of those reported by Henning, as well as the F and G bands of Hopfield and Birge [104]. A band at 783.1 Å (127,698  $\text{cm}^{-1}$ ), reported to be very diffuse by Henning, was seen under higher dispersion to have three close heads, each being sharp and degraded to the red. Tanaka found two sets of absorption bands of equal intensity, one sharp and one diffuse, each forming a Rydberg series labeled the "β" series and converging to a common limit at 630.12 Å = 19.675 eV. The sharp series in the region 700–620 Å can be represented by the formula

$$\sigma = 158692 - \frac{R}{(n-1.68)^2}, n=5, \dots, 12.$$

More recently, Huffman, Larrabee, and Tanaka [107] have observed the  $n=4$ , 0–0 terms of the "β" sharp and diffuse Rydberg series in their study of absorption coefficients of CO in the region 1000–600 Å. (See also table 27.)

Another Rydberg series labeled "α" was found [by Tanaka] to span the region 800–730 Å. Its convergence limit was at 749.74 Å = 16.536 eV. The "α" series can be represented by the formula

$$\sigma = 133380 - \frac{R}{(n-1.69)^2}, n=5, \dots, 9.$$

The "β" series includes transitions from the  $\text{X}^1\Sigma^+$  state of CO to electronic states which converge to the  $\text{B}^2\Sigma^+$  state of  $\text{CO}^+$ . A diffuse band always appeared on the short wavelength side of each sharp band. The "α" series converges to the  $\text{A}^2\Pi$  state of  $\text{CO}^+$ . This latter series is strongly overlapped by other bands. Both "β" and "α" series show vibrational structure which is close to that of the ionic states which are their series limits. Tanaka [254] tentatively ascribed the sharp and diffuse series to  $(^1\Sigma^+) - \text{X}^1\Sigma^+$  and  $(^1\Pi) - ^1\Sigma^+$  transitions, respectively.

Takamine, Tanaka, and Iwata [253] have observed two Rydberg series converging to 884.73 Å and 868.13 Å, respectively. These series are the 0–0 and 1–0 vibrational series whose convergence limit is the lowest ionization potential of CO:

$113029 \pm 32 \text{ cm}^{-1}$  or  $14.013 \pm 0.004 \text{ eV}$ . This convergence limit is within the experimental error of that derived from the limits of the "β" and "α" series and the origin of the B–X system of  $\text{CO}^+$ . (The values in eV quoted here differ from those originally given in ref. 253 because of a slight difference in conversion factor.) The series has been represented by a formula

$$\sigma = 113029 - \frac{R}{(n-1.88)^2}, n=6, \dots, 14.$$

### 3.21. $\text{B}^2\Sigma^+ - \text{X}^2\Sigma^+$ First Negative System of $\text{CO}^+$ (3150–1800 Å) R

The  $\text{B}^2\Sigma^+ - \text{X}^2\Sigma^+$  First Negative system of carbon, originally observed by Deslandres (see also ref. 127a, p. 233), consists of an extensive group of single-headed, red-degraded bands. The upper state of this system is the highest lying state of  $\text{CO}^+$  which has been observed spectroscopically, though a more energetic state has been tentatively identified from electron impact and photoionization measurements. Vibrational heads of the B–X system are given in table 30, origins in table 31. Rotational constants for the  $\text{X}^2\Sigma^+$  and  $\text{B}^2\Sigma^+$  states of  $\text{CO}^+$  are given in tables 50 and 52, respectively.

A number of bands observed by Schniederjost [ref. 127a, p. 213] were later correctly identified by Jevons [116] as belonging to Deslandres' system. Additional bands of this system were reported by Johnson [118] and Blackburn [26] who gave the first vibrational assignments. This system was extended to shorter wavelengths by Biskamp's report [24] of 22 new bands in an intense high frequency discharge in helium containing a trace of CO. The dispersion used was 8.8 Å/mm [wavelengths are uncertain to  $\pm 0.05 \text{ Å}$ ]. A formula was fitted to the observed band heads:

$$\sigma = 45651.5 + (1695.6v' - 24.25v'^2) - (2196v'' - 15.12v''^2)$$

Deviation from this formula of levels with  $v'=7$  suggested [24] perturbation of this level as due to crossing of the  $^2\Sigma$  states.

Early fine structure measurements on lower vibrational levels of both states did not provide reliable rotational constants because of overlapping lines and poor resolution. These include studies by Coster, Brons, and Bulhuis [45] and by Schmid [237]. The latter observed measurable

spin splitting only for high rotational quantum number, typically about 30–34. The 3–5 band, however, showed doubling (about  $1.3\text{ cm}^{-1}$  width) for P(23) and R(24).

The rotational analysis of the 0–0 band by Schmid and Gerö [222] established the ground state as the common lower state of the Comet Tail and First Negative bands.

Rao [202a] has made a comprehensive study of the B–X system of  $\text{CO}^+$  under high resolution which has provided the best values of molecular constants of both states, band origins, rotational constants, and spin doubling coefficients. [For the latter see table 53.] Previously Woods [278] had reported values of  $|\gamma' - \gamma''|$  for the B and X states which are about half those of Rao.

Rao [202a] has given the fine structure of the 0–0, 0–1, 0–2, 0–3, 0–4, 1–2, 1–3, 1–4, 1–5, 2–3, 2–4, 2–5, 2–6, 3–5, and 4–7 bands. Observations were made viewing a hollow cathode discharge with a 21-ft focal length grating spectrograph in fourth order. Six branches are possible for a  $^2\Sigma - ^2\Sigma$  transition, with each line of the P and R branches split into a triplet by the selection rule  $\Delta J = 0, \pm 1$ . The satellite branches  $^PQ_{12}$  and  $^RQ_{21}$  were very weak except at very low N. The bands consist of a doublet P branch and a doublet R branch, with doublet splitting resolved only for higher N values.

### 3.22. $A^2\Pi_i - X^2\Sigma^+$ Comet Tail System<sup>8</sup> of $\text{CO}^+$ (7200–3080 Å) R

Important characteristics of comet tail spectra are bands of carbon compounds. Among these are found the strongest bands of  $\text{CO}^+$ ,  $A^2\Pi_i - X^2\Sigma^+$ , which were first discovered in these celestial sources [189], and shortly thereafter weakly produced in low pressure Geissler discharges in CO by Fowler [64]. These bands which extend from the UV through the visible region have two double heads degraded to the red. Each band shows eight branches (out of a possible 12), four originating from each  $\Pi$  doublet component. The other four are not resolved because of unresolved spin splitting of the  $X^2\Sigma$  state. Rao [202b] has found that previously used  $v'$  numbering should be lowered by three units. Perturbations of levels  $A^2\Pi_i, v=5$  and 10 indicate crossing by the ground state. Table 32 lists the observed

A–X band heads; precise origins are listed in table 33. Rotational constants for the X and A states of  $\text{CO}^+$  are given in tables 50 and 51, respectively.

Merton and Johnson [160], Johnson [118], and more recently Asundi [6] have observed bands of this system in a discharge through helium with a very small partial pressure of CO ( $\sim 10^{-4}$  mm). Baldet [9] observed four heads of each of 40 bands (in the region 6400–3080 Å) produced by electron bombardment of CO in the most extensive single observation of this system. He drew a Fortrat parabola for the eight observed branches of the band at 3997 Å, but the more extensive rotational analysis of many bands to which he alluded has not been published.

Birge [21] showed that the First Negative and Comet Tail bands had a common lower state. His vibrational assignments, as given in an early review by Herzberg [97], were based on a fit to all of Baldet's data with neglect of a weak band at 5281 Å. This analysis did include three comparably weak bands at 6400 Å, 5800 Å, and 6000 Å, whose intensities were estimated as  $\frac{1}{2}$  by Baldet. These were assigned as 0–0, 1–0, and 2–1 transitions. Rao [202b] did not see these weak bands, nor have any B–A bands been observed with  $v''$  values corresponding to the above assignments. In addition, the above quantum assignments gave abnormal Condon parabolas. By lowering the  $v'$  assignments by three units and omitting the four weak bands (listed in table 35i), these difficulties were eliminated [202b].

Much work has been devoted to fine structure analysis of lower lying vibrational levels. Blackburn [25] reported seeing the Comet Tail bands under high resolution, but gave no quantitative data. Sometime later Coster, Brons, and Bulthuis [45] observed the fine structure of the 2–0, 3–0, 4–0, and 5–0 bands produced in a low pressure discharge in flowing CO. Relative accuracy for sharp lines was claimed to be  $0.002\text{ Å}$ . The  $^2\Pi_i$  state was found to be in Hund's coupling case (a) for low rotational quantum numbers. Schmid and Gerö [222] have given the fine structure of the 0–0, 1–0, 1–1, 2–1, 3–2, and 4–2 bands. Rotational analyses of many more bands have been given by Bulthuis [34a, 35] who has published details only of the 6–0 band and the most intense branches of the 7–0, 8–1, 9–1, 10–1, 10–2, and 11–2 bands. Perturbations of the upper level were found in the 10–1 and 10–2 bands.

<sup>8</sup> Earlier referred to as the Third Negative bands of carbon.



The fine structure of the 0-2, 0-3, and 0-4 bands has been given by Rao [202b] who, in addition, reduced previous data [34, 45, 222] to evaluate rotational constants for the  $A^2\Pi_i$  state. Accuracy of  $\pm 0.00025 \text{ cm}^{-1}$  is claimed. The following formula was obtained:

$$B'_v = 1.58940 - 0.01942(v' + \frac{1}{2}), \quad D'_v \sim 6.60 \times 10^{-6} \text{ cm}^{-1} \text{ (est).}$$

An average value of  $-117.5 \text{ cm}^{-1}$  was derived for the coupling constant  $A$  (which is negative because the  $A$  state is inverted). A useful comparison of the branch designations used by previous authors can be found in Rao's paper [202b].

### 3.23. $B^2\Sigma^+ - A^2\Pi_i$ Baldet-Johnson Intercombination System of $\text{CO}^+$ (4240-3310 Å) V

This system of double double-headed, violet-degraded bands has been much less extensively studied than the other two known transitions for  $\text{CO}^+$ . Spin splitting of the upper state is observable; for the lower state it is negligible. Vibrational quantum numbering is based on the work of Rao [202b]. Observed band heads are given in table 34.

Merton and Johnson [160] mentioned a number of faint violet-degraded bands among their  $A-X$  Comet Tail bands and similar in appearance. Johnson [118] reported the band heads for the 0-1, 1-2, 0-0, 1-0, 2-0, and 3-0 transitions. Several additional bands were found which have not been identified (table 35j). Meanwhile,

Baldet [8] had previously reported the 0-1, 0-0, and 1-0 transitions from electron bombardment of low pressure  $\text{CO}$ .

Baldet [10] mentioned high dispersion measurements of the 0-1, 0-0, and 1-0 bands, but his fine structure analysis has not been published. Bulthuis [34b] has given the rotational analysis of the 0-0 and 0-1 bands. These bands extended to moderate  $J$  ( $\sim 25$ ), were very weak, and were overlapped by the much stronger  $A-X$  Comet Tail bands. The bands were observed in a hollow cathode discharge. The spin splitting of the  $B^2\Sigma$  level was observable, and Bulthuis concluded that the spin splitting in the  $B-X$  band for higher  $J$  [34b] is due essentially to the  $B$  state. Rotational constants for each  $^2\Pi$  sublevel were obtained which, when averaged, are in reasonable agreement with more recent results of Rao [202b]. For higher  $J$  the doublets  $P_2-Q_{12}$  and  $R_1-Q_{21}$  are clearly separated, but the doublets  $R_{12}-Q_2$  and  $P_{21}-Q_1$  show only one component even for the highest  $J$  values ( $\sim 25$ ).

Recently, Rao and Sarma [205] have summarized the work on all three observed transitions for  $\text{CO}^+$  stressing the revision of  $v'$  numbering for some  $B^2\Sigma - A^2\Pi_i$  Baldet-Johnson bands, and changes in branch designation because earlier fine structure analysis incorrectly labeled the  $A$  state as  $^2\Pi$ , instead of inverted. Making use of the value for the spin splitting constant  $\gamma = 0.0192$  for  $B^2\Sigma$ ,  $v=0$  and  $|\gamma' - \gamma''|$  obtained previously [202a], they obtained values of  $\gamma$  for the  $X$  state which are given in table 53. Blending of branches of the intercombination bands prevented accurate determination of rotational constants.

## 4. Perturbations

A number of states of  $\text{CO}$  and  $\text{CO}^+$  undergo perturbations, most notably the  $A^1\Pi$  and  $b^3\Sigma^+$ . The  $A$  state has the greatest number of perturbations; the  $b$  state suffers perhaps the most severe disruptions in its rotational structure.

For  $\text{CO}$ , the identity of most perturbing states is known. The existence of the  $I^1\Sigma^-$  state, not directly observed until recently, was revealed from its interactions with the  $A^1\Pi$  state. Supposed predissociations of the  $A$  state [227,74], indicated by intensity drops in the band fine structure, are not real. However, perturbations may occur at these positions. Table 54 summarizes the

data on reported maximum perturbations, some of which are uncertain.

This section which supplements table 54 is concerned only with the reported rotational perturbations. Multiplet splitting and  $\Lambda$ -doubling are discussed in section 3; predissociations, a special case of perturbations, are discussed in section 5.

### 4.1. $A^1\Pi$ State

#### a. $v=0$

Stepanov [244] has determined the matrix elements of interaction of  $^3\Sigma_{J=N}$  with  $^1\Pi_b$  (in inter-

mediate coupling between cases (a) and (b)), and has applied the results to perturbations of  $A^1\Pi$ ,  $v=0$ ,  $J=9, 12, 16$ . By using the data of Gerö [224a, 71], he obtained numerical values of the matrix elements which were found to be functions of  $J$ , and not constant as had been found by Budó and Kovács [33]. The almost linear increase with  $J$  led to the conclusion that the  $A^1\Pi$  state was in intermediate coupling close to case (b), as had been previously established (see sec. 3.1). Stepanov [245] has also mentioned the importance of including the distorting affect of a third level at the common point of intersection for low  $J$ .

Coster and Brons [44] have also reported a perturbation affecting all branches at  $J=1$ ; the perturber is  $\Pi$  or  $\Delta$  since both  $\Lambda$ -components are affected together, most probably  $^1\Pi$  since the perturbation is already large at  $J=1$ . This perturbing  $^1\Pi$  state has not been observed directly. In addition, all lines below  $J=16$  showed deviations of  $6\text{--}10\text{ cm}^{-1}$ ; most of these being due to crossing by  $^3\Sigma$  levels. The  $^3\Sigma$  term value was estimated at  $64815\text{ cm}^{-1}$ ; from the data of Herzberg and Hugo [101]  $e^3\Sigma^-$ ,  $v=1$  was found at  $64803\text{ cm}^{-1}$ . For high  $J$  (presumably  $>30$ ) [44], a  $^3\Pi$  perturbation is mentioned.

Crawford [50] has studied the Zeeman effect in the B-A Ångström bands. [His fig. 5 shows the Zeeman pattern for the P(1) and R(1) triplets of the 0-0 band at  $4511\text{ Å}$  and also the asymmetry of the Q doublets at low  $J$  for the 0-3 band at  $5610\text{ Å}$ .] He has tabulated a number of magnetically sensitive lines which show large Zeeman patterns. These normally occur near points of large rotational perturbations. For the 0-0 band this includes Q(25) and probably others. (See sec. 4.1, b to d for additional magnetically sensitive lines.) Crawford described the appearance of the 0-0 B-A band as having weak and fragmentary extra branches (see sec. 3.2; note particularly the same difficulty of Jasse [115]).

Watson [268, 269] has reported additional faint lines for the 0-0 B-A bands including the occurrence of extra lines at the perturbations. Some of the P, Q, and R lines at the perturbation of  $J\sim 28$  become sharp doublets. The Zeeman pattern indicated that the perturber was not a singlet state, and was assumed to be  $d^3\Pi_4$  (now known to be  $d^3\Delta_4$ ). Extra lines appear for  $J=8, 10$ , and  $17$ .

The perturbed lines show large irregular Zeeman patterns at medium and high  $J$  values. Neighboring lines are apparently insensitive to the

magnetic field. The perturbing state is assumed to be case (a)  $^3\Pi$ . Watson [269] has given a table of Zeeman patterns of the perturbed band lines in the 0-0 B-A band.

#### b. $v=1$

At  $J\sim 0$  Rosenthal and Jenkins [216] found a perturbation in all branches which they ascribed to  $a^3\Pi$ ,  $v=1$ . Schmid and Gerö [225] believed, however, that the perturbing state was  $^1\Pi$ , an unobserved state.

Coster and Brons [43] have found the following perturbations in the 0-1 B-A band: R(26), P(26), Q(23), Q(26), Q(29), and Q(35), with extra lines corresponding to R(26) and Q(29). These seem to include the  $a^3\Sigma^+$  perturbations as well as  $^1\Sigma^-$  perturbations reported elsewhere [262, 225]. The observed deviations do not exceed  $2\text{ cm}^{-1}$ , with the largest deviations accompanied by intensity reductions. Two additional  $\Sigma$  perturbations are found at Q(25) and Q(35) which may also be due to the  $^1\Sigma^-$  state.

Read [209] has observed a perturbation at  $J\sim 26$  in the 1-0 A-X band. Gerö [74] reported a  $^3\Sigma^-$  perturbation at  $J=44$  in another A-X band.

Gerö [71] has found small perturbations at small  $J$  in  $v=1$  with both  $\Lambda$ -components affected the same.

Tschulanowsky and Stepanow [262] observed perturbations in A-X bands, including the appearance of extra lines. In all cases the P and Q branches were perturbed differently and perturbations result from interaction with the  $e^3\Sigma^-$  state. For the 1-5 A-X band an extra line was incorrectly attributed to Q(28) and should be revised as follows: Q(28)= $55571.07$ , Q(29)= $55554.84$ , with the extra line for  $J=29$  at  $55559.10\text{ cm}^{-1}$ . A  $^1\Sigma$  term perturbs both  $J=26$  and  $35$ . An additional perturbation at  $J=26$  [44] is due to one component of a  $^3\Sigma$  state.

Crawford [50] has listed as magnetically sensitive lines Q(23) and Q(24) in the 0-1 B-A band. Near the perturbation at  $J\sim 28$  in this band Watson [268, 269] found that several P, Q, and R lines appeared as doublets.

McCulloh [157a] reported a perturbation of the lower state in the 0-1 B-A band for  $C^{13}O^{16}$ . Perturbed lines include P(16), Q(13), Q(19), and R(16). The perturbing state is  $^3\Sigma^+$  with  $B\sim 1.13\text{ cm}^{-1}$ . (This is the isotopic analog to what Coster and Brons [43] found.)



c.  $v=2$

Rosenthal and Jenkins [216] observed perturbations at  $J=25, 29$ , and  $33$  in the  $0-2$  B-A band. Extra lines appeared in the P and R branches for  $J=25$  and  $33$ . Tschulanowsky and Stepanow [262] saw the same perturbations in the A-X bands including extra lines in the P branch for  $J=25$  and  $26$ . The perturbing state is  $e^3\Sigma^-$ ,  $v=4$  (from the data of Herzberg and Hugo [101]). From perturbations [44], the term value of  $67953\text{ cm}^{-1}$  is obtained.

Gerö [73] observed doubling of the Q(7) line in A-X bands. The perturber is  $^1\Sigma^-$ . Deutsch and Barrow [52] observed this perturbation in the B-A bands, and have tabulated the perturbations caused by the  $I^1\Sigma^-$  state for all vibrational levels of the  $A^1\Pi$  state. Crawford [50] gives the following magnetically sensitive lines in the  $0-2$  B-A band: P(25), P(26), P(34), and P(35).

d.  $v=3$

Rosenthal and Jenkins [216] reported perturbations at  $J=28, 31$ , and  $34$  in B-A bands. Coster and Brons [44] showed that these levels are crossed by the  $a'^3\Sigma^+$  state. The estimated term value for the perturber is  $69459\text{ cm}^{-1}$ ;  $a'^3\Sigma^+$ ,  $v=12$  is found at  $69602\text{ cm}^{-1}$  from the data of Herzberg and Hugo [101]. Coster and Brons [44] found another small perturbation of Q(38) which is due to the  $^1\Sigma^-$  state.

Gerö [74] found  $J=50$  perturbed by the  $e^3\Sigma^-$  state. Tschulanowsky and Stepanow [262] have seen the  $^3\Sigma$  perturbation in the A-X bands. In the  $3-7$  band extra lines are found for Q(27) and Q(28). The extra line reported for Q(29) may not be real, and cannot be assigned to  $J=28, 29$ , or  $30$ . Crawford [50] includes the Q(28) line as magnetically sensitive in the  $0-3$  B-A bands. Deutsch and Barrow [52] found  $I^1\Sigma^-$  perturbing  $J=40$ .

McCulloh [157a] found the maximum perturbation at Q(21) in the  $0-3$  B-A band of  $C^{13}O^{16}$ . P(21) was unperturbed. Either the  $^3\Sigma^+$  state or  $^1\Sigma^-$  is responsible for the perturbation; the proper assignment can only be established from examination of high rotational levels.

e.  $v=4$

Coster and Brons [44] have found a perturbation at  $J\sim 31$  in the  $0-4$  B-A band, the reality of which is disputed by Gerö [73]. Another perturbation

at  $J=34, 37$ ,—, is due to  $e^3\Sigma^-$ . The estimated term value of  $^3\Sigma$  is  $70878\text{ cm}^{-1}$ ; from Herzberg and Hugo [101], one obtains  $e^3\Sigma^-$ ,  $v=7$  at  $70966\text{ cm}^{-1}$ .

Gerö [73, 71] has reported a  $^3\Sigma^+$  perturbation at  $J\sim 0$  and two others at  $J=23$  and  $27$ . The one at  $J=23$  is due to  $^1\Sigma^-$  [52]. Asundi [5], in addition, observed a very small perturbation for  $J=18$ . Gerö [73] has reviewed the work on a number of these perturbations and gives the following identifications:  $J\sim 0$ ,  $^3\Sigma^+$ ;  $J=23$ ,  $^1\Sigma^-$ ;  $J=37$ ,  $^3\Sigma^-$ . Additional perturbations include  $J=46$ ,  $^3\Sigma^+$ ;  $J=50$ ,  $^3\Pi$ .

f.  $v=5$

Gerö [73] has referred to close crossing of  $^3\Sigma^+$  and  $^3\Pi$  states which overlap perturbations of A,  $v=5$ . The first perturbation is at Q(30) in A-X bands, and is due to  $^3\Sigma^+$ . The last perturbation of this pair of overlapping states is at R(35), Q(36), and P(37), due to  $^3\Pi$ . A  $^1\Sigma^-$  perturbation is found at Q(44) [73, 52], and the onset of a  $^3\Sigma^-$  perturbation at R(50).

g.  $v=6$

Gerö [73] has observed overlapping  $^3\Pi$  and  $^3\Sigma$  perturbations near the heads of A-X bands. An additional perturbation was observed at Q(52). Since the other branches were not followed to such high quantum numbers the nature of the perturbing state could not be determined with certainty, and was assumed by Gerö to be  $^3\Pi$ . Several additional perturbations at  $J=29, 30$  have also been found [73, 52].

Onaka [182] later unraveled the structure near the head in his observations of the  $6-13$  and  $6-14$  A-X bands. Details of the  $6-14$  band are not given. A perturbation near  $J=0$  which acts on all branches, indicates a  $\Pi$  or  $\Delta$  perturber with  $B_v$  larger than that of the A state. (A similar perturbation has been found of A,  $v=0$  by Coster and Brons [44], indicating the possibility of a common state perturbing both the 0 and 6 levels.) Several extra lines are observed for  $v=6$ , which include P(7), R(7), Q(11), P(13), Q(13), and Q(16). The perturbations observed are:  $J'=7\sim 8, 11, 14, 16$  in Q; and  $J'\sim 0$ , all branches.

h.  $v=7$

Read [209] has observed perturbations of R(25), Q(27), and P(27) in the  $7-1$  A-X band. Another

perturbation at  $J=29$  was designated as  $^3\Sigma^-$  by Gerö [73]. In addition, Gerö found a  $^3\Sigma^+$  perturbation at  $J=26$ , and a  $^3\Pi$  perturbation at  $J=39$ . Gerö [74] mentions a sudden intensity drop from R(46) to R(47) in the 7-16 A-X band with no simultaneous line displacement.

#### i. $v=8$

Gerö [73] found the following perturbations in A-X bands:  $J=18$ ,  $^3\Sigma^+$ ;  $J=27$ ,  $^3\Pi$ ;  $J=34$  [52],  $^1\Sigma^-$ . Gerö [74] mentions the following sudden intensity drops (no line displacement)

8-16, A-X R(35)  $\gg$  R(36)  
8-17, A-X R(35)  $\gg$  R(37); R(36) overlapped.

#### j. $v=9$

Gerö [73] found a  $^1\Sigma^-$  perturbation at  $J=13$ . (See also ref. 52.) At  $J=37$ , a  $^3\Sigma^+$  perturbation is expected [74]. Gerö [74] lists the following intensity drops unaccompanied by line displacements:

A-X

9-17, P(22)  $>$  P(23); R branch overlapped  
9-18, P(22)  $\gg$  P(23); R(19)  $\gg$  R(22);  
R(20), R(21) overlapped; Q(27)  $\gg$   
Q(28); P(28)  $>$  P(29)  
9-19, P(22)  $>$  P(23); R(19)  $\gg$  R(21); R(20)  
overlapped  
9-20, P(23) and R branch overlapped; Q(27)  
 $\gg$  Q(28).

#### k. $v=10$

Gerö [73] has studied numerous perturbations of the  $A^1\Pi$  state in the A-X bands. It is determined that for  $v'=10$ , at  $J=16$  and 18, there are perturbations of P(17) and Q(18) by  $^3\Sigma^-$ . The strongest perturbations are overlapped by the  $^3\Sigma^+$  state. The branch lines, however, are of normal intensity beyond the perturbed level and can be followed to  $J=34$ . At  $J=32$  there is a  $^3\Pi$  perturbation.

#### l. $v=11$

Gerö [74] found a perturbation of  $J=16$  by  $^3\Pi$  and  $^1\Sigma^-$ .

#### m. $v=12$

Gerö [74] found a perturbation of  $J=24$  by  $^3\Sigma^+$ .

#### n. $v=13$

Gerö [74] has analyzed lines only up to  $J=15$ . He indicates a probable  $^3\Pi$  perturbation near that limit. For the bands with  $v=11, 12, 13$  only one band was observed, so that localizing the perturbations is less certain than for the other bands.

### 4.2. $E^1\Pi$ State

In the 0-0 E-X band Tilford et al., [257] observed a perturbation at  $J'=31$  in both R(30) and P(32). These lines appear as doublets, and, in addition, are anomalously more intense than their neighbors. A state of type  $^1\Sigma^+$  is the suspected perturber.

### 4.3. $a^3\Pi$ State

Beer [16] has studied the perturbations of the  $a^3\Pi$ ,  $v=4$  level in the 0-4 and 1-4 bands of the  $b^3\Sigma^+ - a^3\Pi$  Third Positive bands. Eight observed crossing points are attributed to a perturbing  $^3\Sigma^+$  term for which  $B_v=1.31$  cm $^{-1}$  was estimated. The perturbing state, which lies at an energy of 55380 cm $^{-1}$  above the reference level  $X^1\Sigma^+$ ,  $v=0$ ,  $J=0$ , has been identified as  $a'^3\Sigma^+$  [235]. From previous  $b-a$  studies by Dieke and Mauchly [54], and Gerö [75] it is shown that the  $a^3\Pi$ ,  $v=0, 1, 2$ , and 3 terms are regular (unperturbed). Dieke and Mauchly [54] tried an analysis of the 0-4 Third Positive band, but could only conclude that strong perturbations were probably present. Beer found no perturbations of  $v''=5$  from the 0-5 and 1-5 bands. [In these bands the  $b^3\Sigma$ ,  $v=0$  and 1 levels are subject to strong and consecutive perturbations [54, 75]. (See sec. 4.6.)

The eight perturbations are listed in table 54c. Numerous extra lines appear in the 0-4 and 1-4 bands.

Since the  $a^3\Pi$  state is no longer in coupling case (a) for medium values of rotational quantum number [50], one can allow for violation of the selection rule for perturbations  $\Delta\Omega=0, \pm 1$ . In fact, in two places, the perturbations occur with  $\Delta\Omega=2$ . The absence of a perturbation at the ninth crossing point where  $\Delta\Omega=3$  results from the rules for intermediate coupling. From the perturbations,  $B \sim 1.31$  cm $^{-1}$  and  $\sigma_0 \sim 55,380$  cm $^{-1}$  (above X) are calculated for the perturbing state. These values correspond closely to those of the  $a'^3\Sigma^+$ ,  $v=0$  level. The 0-6 and 1-6 bands should show the onset of perturbations sooner, but these were not analyzed due to insufficient light inten-



sity. The lowest perturber level of  $a'^3\Sigma^+$  is most certainly  $v=0$ , since  $a^3\Pi$ ,  $v=0, 1, 2$ , and  $3$  show no perturbation. The Third Positive 0-6 and 1-6 bands should show perturbations at  $J\sim 47$ . These have not yet been observed.

Beer's estimates of the constants for the  $a'$  state agree fairly well with the accurate determinations by Herzberg and Hugo [101].

#### 4.4 $d^3\Delta_i$ State

Gerö and Szabó [82] have reported perturbations in all components of the  $v=6$  term of the  $d$  state in the  $d^3\Delta_i-a^3\Pi$  Triplet bands at  $J=8$ ,  $J=13-14$ , and  $J=14-15$ .

#### 4.5. $f^3\Sigma^+$ State

In a study of the 2980 Å and 2670 Å bands of the  $f^3\Sigma^+ \rightarrow a^3\Pi$  transition, Gerö [78] observed perturbations of the upper levels by the  $v=0$  and  $v=1$  levels of the  $b^3\Sigma^+$  state. The perturbations are seen in the 2670 Å band at  $N=13$  in the  $^PQ_{12}(N)$ ,  $P_2$ ,  $R_2$ , and  $^RQ_{32}$  branches. The  $N=13$  level for the other band also seems perturbed.

A plot of  $\frac{\Delta_2 F(N)}{4(N+\frac{1}{2})}$  versus  $N$  indicates dependence on  $N$  for small  $N$ ; hence, a perturbation at  $N\sim 0$  by another  $^3\Sigma^+$  term. (See table 54e.) These bands had at one time been thought to originate from high vibrational levels of the  $a'^3\Sigma^+$  state.

#### 4.6. $b^3\Sigma^+$ State

Dieke and Mauchly [54] have indicated that numerous large perturbations of the upper state of the  $b^3\Sigma^+-a^3\Pi$  Third Positive bands cause the characteristic intensity fluctuations previously observed by Wolter [277]. The perturbations were ascribed to successive high vibrational levels of the  $^3\Sigma$  perturber, presumably  $a'^3\Sigma^+$ . These are mixed with levels of the state  $f^3\Sigma^+$ . The first large perturbation in  $v=0$  is at  $N=19$ , hence the head region is not affected. However, there is no apparent regularity of lines in the middle of the bands. The perturbations result in clustering of lines in some places and gaps in others. Similar effects which are expected for lower  $N$  values near the head region for bands with  $v'=1$  would make the  $v'=1$  progression appear to belong to another system. In the  $v'=0$  progression for  $v''=0$  to  $4$ , the last few lines in all branches showed perturbations of increasing magnitude with increasing  $N$  values, and the lines did not return to their regular

sequence after perturbation. This indicates a  $\Pi$  or  $\Delta$  perturber. [Fig. 2 of ref. 54 shows the irregularly spaced triplets.] Deviations of a single perturbation exceed  $50\text{ cm}^{-1}$  [53].

Gerö [72] determined numerous perturbations whose spacing decreased with increasing  $N$ .

Gerö [75] observed the disrupted structure of the 1-0  $b-a$  band. Perturbations made it impossible to isolate branches below  $N=5$ . Often lines did not follow in order of the quantum assignments. Additional perturbations were found for  $^3\Sigma$  levels  $N=17, 25, 30, 34$ , and  $40$ . The perturber is the same  $^3\Sigma$  state which affects the  $b^3\Sigma^+$ ,  $v=0$  level.

The perturber would appear to be the  $f^3\Sigma^+$  state [78, 247, 101], incorrectly identified previously as high vibrational levels of the  $a'^3\Sigma^+$  state. (See further details on the  $f-a$  transition in the section on electronic transitions.) The term values of the perturbing levels lie at  $83744\text{ cm}^{-1}$  and  $85969\text{ cm}^{-1}$ .

#### 4.7. $A^2\Pi_i$ State of $\text{CO}^+$

##### a. $v=5$

The  $A^2\Pi_{1/2}$  ( $v=5$ ) level is perturbed by the  $X^2\Sigma$  ( $v=14$ ) level [45]. In the 5-0,  $A^2\Pi_{1/2}-X^2\Sigma^+$  Comet Tail band, the perturbations include a number of successive extra lines in both the  $Q_1$  and  $Q_2$  branches (most of them being displaced towards higher frequency, one displaced towards lower frequency). The perturbing state is  $\Sigma$ . The perturbations occur for a large number of lines, because very similar  $B_v$  values for both interacting levels cause both sets of levels to stay close together. At  $J=14\frac{1}{2}$  is the maximum perturbation of one  $\Lambda$ -component of the  $\Pi_{1/2}$  level by the  $\Sigma$  level with  $J=N-\frac{1}{2}$ ; at  $J=22\frac{1}{2}$ , the other  $\Lambda$ -component is perturbed by  $\Sigma$ ,  $J=N+\frac{1}{2}$ ; the  $\Pi_{3/2}$  level is unperturbed. The perturbing  $\Sigma$  level lies between  $\Pi_{1/2}$  and  $\Pi_{3/2}$  for low  $J$ . (See fig. 4 of ref. 45.) This, together with the fact that the peak perturbations lie so far apart, indicates a relatively slight angle between the potential curves of the interacting  $\Pi$  and  $\Sigma$  levels.

For small  $J$ , the  $\Pi_{3/2}$  level is probably not perturbed. From extrapolation of data for the 2-0, 3-0, and 4-0 bands, any displacement would not exceed several tenths of a  $\text{cm}^{-1}$ . Both  $\Lambda$ -components of the  $\Pi_{1/2}$  level are displaced about  $2\text{ cm}^{-1}$  for low  $J$ . The perturber is  $X^2\Sigma^+$ . From the two crossing points  $X_{v=14} A_{v=5}$ , a value of  $B_{\Sigma}\approx 1.7\text{ cm}^{-1}$  is determined which is in fair agreement with the value for  $X^2\Sigma$ ,  $v=14$ . The

extra lines are due to the  $X^2\Sigma(v=14)-X^2\Sigma(v=0)$  transition. It seemed strange that perturbations were not also observed in the other bands considered here (2-0, 3-0, and 4-0). This may be explained by the fact that the observed crossing of  $\Pi$  and  $\Sigma$  here indicates that the potential curves separate rapidly from one another, and have different  $r_e$ . The  $\Sigma$  levels would appear to fall between the  $\Pi$  levels for low  $v$  and low  $J$  so that a crossing with  $\Pi_{3/2}$  is not expected till very high  $J$ . In fact, there appears to be a perturbation of  $\Pi_{3/2}-\Sigma(4-0)$  for  $J>30_{1/2}$ . (Probably due to  $X$ ,  $v=10$  affecting the low energy  $\Lambda$ -component).

#### b. $v=10$

In the 10-1 and 10-2  $A^2\Pi-X^2\Sigma^+$  Comet-Tail bands, Bulthuis [35] found a perturbation of the  $R_{12}Q_{11}$  and  $Q_{12}P_{11}$  branches. The perturbed level is  $A^2\Pi_{1/2}$ ,  $v=10$ . From a knowledge of the  $v'=9$  and 11 levels the unperturbed branches of  $A^2\Pi_{1/2}-X^2\Sigma^+$  for  $v'=10$  can be accurately calculated by addition of the mean values of  $A^2\Pi_{1/2}-A^2\Pi_{3/2}$  differences of these two known levels to the measured branches of  $A^2\Pi_{3/2}-X^2\Sigma^+(v'=10)$ . The O-C (observed-calculated) differences for the lines are the energy shifts due to the perturbation. The shifts are small (about  $3\text{ cm}^{-1}$  at most) (see fig. 3 of ref. 35) and include  $J=3\frac{1}{2}$  to  $21\frac{1}{2}$ , the

range observed. This can be understood from the results of Ittmann's [112] theory where the matrix elements determining the magnitude of the shifts are small for small  $J$ .

The perturbations occur for numerous  $J$  values as can be seen from figure 3 of reference 35. Here, for  $v=10$ , the crossing of perturbing levels occur at very low  $J$  values. Hence, the perturbation is observed for many  $J$  values because for small  $J$  the two perturbing potential energy curves are more parallel than in the case of high  $J$ , since the energy differences for the same  $\Delta J$  increases with  $J$ . This follows also because the interacting levels have very similar  $B$  values.

All perturbing energies are negative (O-C) as seen from figure 3 of reference 35; hence, the perturbing level lies entirely above the perturbed  $A^2\Pi_{1/2}$  level. By making use of Biskamp's [24] band head formula, the perturber is found to be  $X^2\Sigma^+$ ,  $v=18$ . Extrapolating from figure 3 of reference 35, for  $J=0$ , the ratio of perturbations in the two  $\Lambda$ -components is 2.5:3.6. This ratio can be used to calculate the energies of the two perturbing  $^2\Pi-^2\Sigma$  levels using Ittmann's [112] formulas. See figure 4 of reference 35 for energy curves of the perturbing levels.

The details of the perturbations are given in table 54g.

## 5. Dissociation Energies, Predissociations, and Convergence Limits

### 5.1 Dissociation Energy of CO

After a long history of controversy, the dissociation energy of carbon monoxide  $D^0(\text{CO})$ , having a value of  $89460\text{ cm}^{-1}$  or  $11.091\text{ eV}$  [28], can now be regarded as generally accepted, though the triplet levels of the atomic dissociation products are not known. From the noncrossing rule the products would be  $\text{C}(^3\text{P}_0)+\text{O}(^3\text{P}_2)$ , but this has not been established as yet. Since the  $X^1\Sigma^+$  state of CO dissociates to ground state products [69, 56] it is possible to make a unique correlation between a number of excited molecular states and their dissociation products (table 1).

Reviews concerned with the dissociation energy of CO have favored a variety of values. (Refs. 98, 235, 70, 87, 84, 69, 151, 88, 28, 56 and others.) The work of Gaydon [69], Douglas and Møller [56], Brewer et al. [27], and Brewer and Searcy [28] has resolved this question in favor of the "high" value, with several minor inconsistencies still to be ironed out.

In his review a decade ago, Gaydon [69] favored a value of  $D^0$  near  $11.1\text{ eV}$  based on application of the noncrossing rule and an exhaustive comparison of the evidence from spectroscopy, electron impact, photoionization, and determinations of the heat of sublimation of carbon  $L(\text{C})$ . Since then, Douglas and Møller [56] have made high dispersion observations ( $0.23$  to  $2.5\text{ \AA/mm}$ ) of the  $B^1\Sigma^+-A^1\Pi$  Ångstrom bands (including those of isotope  $\text{C}^{13}\text{O}^{16}$ ) and  $A^1\Pi-X^1\Sigma^+$  Fourth Positive bands. They also reviewed the data on suspected predissociations of the  $A^1\Pi$  state concluding that none of these were real. This laid to rest predissociation data in support of values of  $D^0$  lower than  $11.1\text{ eV}$  and confirmed the value obtained from predissociations of the  $B^1\Sigma^+$  and  $b^3\Sigma^+$  states. The results [56] also denied the existence of any appreciable maxima of the potentials in CO, though Hagstrum [87] had claimed one of about  $0.2\text{ eV}$  at  $r>2.9\text{ \AA}$  for the  $a'^3\Sigma^+$  state.

Brewer and Searcy [28] list as best value for  $D^0(\text{CO})$   $89460\pm150\text{ cm}^{-1}=11.091\pm0.019\text{ eV}$ .



The probable error is  $\pm 92 \text{ cm}^{-1}$ . These uncertainties are based on the value  $89595 \pm 30 \text{ cm}^{-1}$  obtained by Douglas and Møller [56] as an upper limit assuming the dissociation products are  $\text{C}(^3\text{P}_0) + \text{O}(^3\text{P}_2)$ . This was in close agreement with the value  $89620 \pm 50 \text{ cm}^{-1}$  obtained earlier [72,224] from a limiting curve of dissociation. The lower limit is  $89325 \text{ cm}^{-1}$  if the dissociation products are  $\text{C}(^3\text{P}_2) + \text{O}(^3\text{P}_0)$ . The mean of these extreme values is that given by Brewer and Searcy.  $D^0$  is defined as the energy difference between the ground electronic state of the molecule and the energy of the atomic dissociation products in their ground states (e.g.,  $\text{C}(^3\text{P}_0) + \text{O}(^3\text{P}_2)$ ). Although it is most likely that the dissociation products are in their ground states, the mean value given by Brewer and Searcy is listed here as best value in order to underline the uncertainty. For thermodynamic calculations, however, ground state products are assumed to correlate with the observed predissociation. Thus Evans and Wagman [61] have used  $89595 \text{ cm}^{-1}$  for  $D^0(\text{CO})$  to be consistent with previous usage.

Recent electron impact work [88, 62] and shock wave experiments [259, 132] also support a value of  $D^0(\text{CO})$  of about 11 eV.

The only spectroscopic evidence supporting a lower value of  $D^0$  is the fact of missing bands of  $b^3\Sigma^+ - a^3\Pi$  with  $v'=2$ , discussed by Gaydon [69] and Barrow et al., [13] which suggests  $D^0 \leq 10.94 \text{ eV}$ . This missing structure is presently an anomaly which requires further consideration. The reported observation of the 2-0 B-X band [255] which lies above the dissociation limit is somewhat uncertain, though the  $v'=2$  level could be populated by inverse predissociation. Only the early work of Ångström and Thalén (in 1875) [see footnote in ref. 23] mentions a band for B,  $v=2$ ; but this too is uncertain.

## 5.2. Predissociations and Convergence Limits

The most precise determination of  $D^0(\text{CO})$  has been obtained from the data on predissociations of  $\text{B}^1\Sigma^+$ ,  $v=0,1$  and  $b^3\Sigma^+$ ,  $v=0,1$  by the common state  $a'^3\Sigma^+$  (or  $f^3\Sigma^+$ ) [72, 56]. In these instances, where the rotational structure is broken off for at least two successive vibrational levels [235], the actual dissociation limit lies close to the predissociation limit. Furthermore, the state causing the predissociations must be bound. Glockler [84] is therefore incorrect in his assertion that a repulsive state predissociates the B state.

Termination of vibrational structure in other states may indicate predissociation [235, 221] by a state or states as yet unspecified, and possibly repulsive. Such is the case for the states  $\text{C}^1\Sigma^+$ ,  $c^3\Sigma^+$ , E, and F, for which only  $v=0$  or 1 is known.

A dissociation limit may be obtained from convergence of a state's vibrational structure or from a Birge-Sponer extrapolation (which should only be made with caution for excited states). In this regard, the  $a'$  and F states raise the most interesting questions (sec. 5.2e, f).

Discussion of the known and suspected predissociations, and convergence limits follows. A summary of the predissociations is given in table 55.

### a. Predissociation of the $\text{B}^1\Sigma^+$ state

Coster and Brons [43, 44] reported a sudden drop in intensity by a factor of two for the lines P(39), Q(38) and R(37) of the 0-1  $\text{B}^1\Sigma^+ - \text{A}^1\Pi$  Ångström band. A similar intensity drop was observed by Read [209] in the 0-0 B-X band. This effect was attributed to predissociation of  $\text{B}^1\Sigma^+$ ,  $v=0$ ,  $J=38$  by some triplet state whose dissociation limit was near the predissociation limit. A forbidden transition was indicated because the rotational structure was weakened, not terminated.

Hulthén [108] observed lines of the 1-1 B-A band up to  $J'=14$  in the P branch and  $J'=17$  in the Q and R branches, but made no mention of intensity irregularities. Schmid and Gerö [224] reported a sudden intensity drop for  $J' \geq 18$  in the 1-0 and 1-1 B-A bands, i.e., 1-0, P(18)  $\gg$  P(19); Q(17)  $\gg$  Q(18)]. A dissociation limit of  $89620 \pm 47 \text{ cm}^{-1}$  was obtained [74,226].

Douglas and Møller [56] reexamined the B-A Ångström bands and confirmed these predissociations. In addition, they established these predissociations for  $\text{C}^{13}\text{O}^{16}$  at  $\text{B}^1\Sigma^+$ ,  $v=0$ , between  $J=39$  and 40, and  $v=1$ , between  $J=19$  and 20. (A portion of the 0-2 band of  $\text{C}^{13}\text{O}^{16}$  showing the weakening of the lines is reproduced in fig. 1 of ref. 56.)

### b. Predissociation of the $b^3\Sigma^+$ state

A convergence limit for the perturbations of  $b^3\Sigma^+$ ,  $v=0$  is found for  $55 < N < 56$ , and considered by Dieke and Mauchly [54] to be a result of predissociation. It is not possible to detect the predissociation by a sudden intensity drop because



the numerous perturbations already confuse the intensity distribution of the lines. Indirect evidence supporting the predissociation is that the analysis of Dieke and Mauchly [54] stops at  $N'=55$ . Of all the branches observed by them 23 of the strongest 36 break off just at  $N'=55$ : Q(55) twelve times, P(56) five times, and R(54) six times. No line was found for  $N' > 55$ .

It is assumed that the cause of this predissociation as well as that of  $B^1\Sigma^+$  is the same—the  $f^3\Sigma^+$  state (or  $a'^3\Sigma^+$  state). (Ref. 72, fig. 2 is obtained from the method of Buttenbänder and Herzberg.)

Brons [30] confirmed the predissociation of  $b^3\Sigma$ ,  $v=0$ ,  $J=55$  in his observation of an intensity drop in the 0–1 Third Positive band. The suspected predissociation [54,226] is found by Gerö [75] in the 1–0  $b^3\Sigma^+ - a^3\Pi$  Third Positive band at  $N \sim 43$ . The branch lines  $R_{1,2,3}(41)$ ,  $Q_{1,2,3}(42)$ , and  $P_{1,2,3}(43)$  all have regular intensity; the next lines are unobserved.

The 1–0 Third Positive band reported by Schmid and Gerö [226] at 2673 Å [it should be 2665 Å] shows successive perturbations [clusters of lines looking like triplets] converging to a limit at 2628.3 Å (38036.7  $\text{cm}^{-1}$ ), which seems to be the same predissociation limit as that for the B state in the Ångström bands [72]. If it is assumed that the perturbations are caused by successively closer terms blending into a continuum, then a transition from this lower limit to the  $a^3\Pi$  state corresponds to an energy of 38036.7  $\text{cm}^{-1}$  ( $v=0$ ) which corresponds to Q(43). (The limit at 38036.7  $\text{cm}^{-1}$  corresponds to Q( $N=43$ ). See also ref. 75.)

#### c. Predissociation of the $C^1\Sigma^+$ state

Schmid and Gerö [224b, 221] have interpreted as due to predissociation a sudden intensity drop at  $J=29$  in band lines of the  $C^1\Sigma^+ - A^1\Pi$  Herzberg bands originating from level  $v'=0$ . This limit is about 3900  $\text{cm}^{-1}$  or 0.48 eV above the limit obtained from the  $B^1\Sigma^+$  predissociation, and hence provides only an upper limit to the dissociation energy. The  $C^1\Sigma^+$  state lies below  $c^3\Sigma$ ; for neither state are there bands with  $v' > 1$ . This would appear to confirm the reality of the predissociation. (See ref. 235.) The C–X, 1–0 band has been observed in absorption [276] and shows no evidence of predissociation. However, predissociation is difficult to detect in absorption.

Gerö [74] included in his tabulation of suspected predissociations the breaking off in the C state of rotational structure between 93486–93608  $\text{cm}^{-1}$  (average 93552  $\text{cm}^{-1}$ ). What has proved to be an incorrect assignment of  $C(^1S) + O(^1D)$  at 93335  $\text{cm}^{-1}$  above X led Gerö to conclude that a repulsive state caused the predissociation because of the 217  $\text{cm}^{-1}$  difference between the atomic product and predissociation limits (i.e., kinetic energy of the dissociation products).

#### d. Suspected Predissociation of the $c^3\Sigma^+$ state

Schmid [221] reported a possible predissociation of the  $c^3\Sigma^+$  state at 11.5 eV because only the  $v'=0$  progression is known for the  $c^3\Sigma^+ - a^3\Pi$  3A bands, and the upper level lies close to the  $C^1\Sigma^+$ ,  $v=0$  level which undergoes predissociation. A further reason for suspecting predissociation [77] is that the few observed bands of the  $c^3\Sigma^+ - a^3\Pi$  3A system are short and drop off in intensity rather abruptly about 10 Å from the head, while similar exposures of other bands stretch for 50 to 100 Å. ( $J_{\text{max}}$  in this case, assuming  $T=300^\circ\text{K}$  and  $10^4^\circ\text{K}$ , gives  $\sim 7$  and  $\sim 13$ , respectively.)

Because of overlapping of the 3A and  $A^1\Pi - X^1\Sigma$  Fourth Positive bands, it is not possible to determine the exact rotational quantum number of the intensity drop in any branch. However, from the predissociation limit of  $C^1\Sigma^+$  at 93554  $\text{cm}^{-1}$ , the position of the  $c^3\Sigma^+$  state at 92076  $\text{cm}^{-1}$ , and  $B_0[a^3\Pi]=1.9563 \text{ cm}^{-1}$ , the first weakened level is found to be at  $J \sim 27$ .

#### e. Convergence Limit of the $a'^3\Sigma^+$ state

The  $a'^3\Sigma^+$  state is known up to  $v=23$  from  $a' - X$  bands [101]. Perturbations of  $A^1\Pi$  identify  $a'$ ,  $v=10$  to 26; perturbations of  $b^3\Sigma^+$  have been identified as due to  $a'$ ,  $v=32$  to 41 [234a]. Some uncertainty exists [69,99] in the identity of the higher lying levels which are probably confused with terms of the  $f^3\Sigma^+$  state, about which little is known. Rotational constants for these levels go to zero at a term value of about 89600  $\text{cm}^{-1}$  above X [234a], which corresponds to the known dissociation limit.

Gaydon and Penney [70] have applied the non-crossing rule to the  $a'^3\Sigma^+$  state and concluded that this state must go to ground state atoms. If not, then the unlikely possibility arises that two  $^3\Sigma^+$  states dissociating to ground state atoms must lie completely below the  $a'$  state (and have been



unobserved to date). The regularity of the  $a'$  state's vibrational levels (up to  $v=23$ ) does not indicate avoidance of crossing.

Two bands earlier believed to originate in high vibrational levels of the  $a'^3\Sigma$  state most probably originate from the otherwise unobserved  $f^3\Sigma^+$  state [78]. Some interaction to avoid crossing by these two  $^3\Sigma^+$  states appears unavoidable. Perturbations of the  $b^3\Sigma^+$  attributed to the  $a'$  and  $f$  states may arise from their joint interaction (e.g., perturbation of three states).

#### f. Convergence Limit of $F(^1\Sigma^+)$ state

The  $F(^1\Sigma^+)$  state with  $\omega_e=2112\text{ cm}^{-1}$  and extraordinarily large  $\omega_ex_e=198\text{ cm}^{-1}$  probably has the flattest potential curve and smallest dissociation energy of all stable states of CO [224b, 105]. (But see the footnote to F state in table 1.) A linear extrapolation gives a dissociation energy of 0.5 to 0.6 eV. Gaydon cautions against Birge-Sponer extrapolations for excited states, especially when some may have potential maxima. However, the rather large value of  $\omega_ex_e$  for the F state implies rapid convergence of the vibrational levels.

If the dissociation products are known for F, then  $D^0$  for the X state can be obtained without knowledge of predissociations. Schmid and Gerö [224b] correlated this dissociation limit at  $\sim 12.9$  eV with  $C(^1S)+O(^1D)$ , which led to a low value of  $D^0$ .

The details of the F state have not been published, and its identity as  $^1\Pi$  is uncertain [69, 99, 147]. It was thought to be definitely singlet [224b] because of the high intensity of the  $F\leftarrow X^1\Sigma$  transition so that it would dissociate into either two triplet or two singlet atoms. Howell [105] rules out the former [but doesn't say why] and offers most probably products as  $C(^1D)+O(^1D)$ . This would mean a dissociation energy of the F state of about 4.4 eV which appears much too high. On this basis he obtained  $D^0(X)$  as 9.6 eV from the energy difference  $[C(^1D)+O(^1D)]-[C(^3P)+O(^3P)]=3.3\text{ eV}$ . However, using  $D^0$  as 11.09 eV and the F state dissociation limit as 12.9 eV it follows that the dissociation products are most likely  $C(^3P)+O(^1D)$ , indicating a triplet molecular state. This could imply a dissociation energy of the F state of about 0.9 eV. Lefebvre-Brion et al., [147] predict the F state to be  $^1\Sigma^+$ .

The branches of a  $^3\Sigma^--^1\Sigma^+$  transition are  $^P P$ ,  $^P Q$ ,  $^R Q$ , and  $^R R$  which under low resolution may look like the P, Q, and R branches of a  $^1\Pi-^1\Sigma$  transition. The intensity of the F-X transition suggests that probably a singlet state mixes with the F state (if triplet) to make strong an otherwise forbidden transition. The identity of the F state remains uncertain.

### 5.3. Dissociation Energy of $CO^+$ , Ionization Potentials and Rydberg Series

The dissociation energies of the ground states of CO and  $CO^+$  are related as follows:

$$D^0(CO^+) = D^0(CO) + IP(C) - IP(CO).$$

Substitution of known values into the right hand side gives

$$\begin{aligned} D^0(CO^+) &= 11.091 + 11.267 - 14.013 \\ D^0(CO^+) &= 8.345\text{ eV} \end{aligned}$$

$IP(C)$  was taken from Moore [166];  $IP(CO)$  was obtained from the Rydberg series limit of Takamine et al. [253], and  $D^0(CO)$  was taken from the predissociation limit in  $B^1\Sigma^+$  of CO [28]. (The conversion factor used is given in appendix B.) If the dissociation products of the ionic molecule are known, then the dissociation energies of the various states can be given unequivocally.

Gaydon [69] has given a succinct review of earlier estimates for the dissociation energies of the three known states of  $CO^+$ , based mostly on Birge-Sponer extrapolations of the vibrational levels of these states and the noncrossing rule. These arguments will be mentioned briefly for the sake of completeness. For the ground state of  $CO^+$ , a linear extrapolation by Biskamp [24] beyond  $v=13$  gave  $D\sim 9.9$  eV, which Gaydon and Penney [70] believed was not accurate because of its length. The vibrational levels of the A state extrapolate beyond  $v=14$  to  $\sim 9.2$  eV [70]. The levels of the B state beyond  $v=10$  extrapolate to about 9.3 eV above the X state [24,70].

There is no direct evidence as to the + or - symmetry of the X and B states [70]. By assuming that they are both  $\Sigma^+$  states [99, 69], it must necessarily be concluded that the X and B states have different dissociation limits to avoid crossing. The lowest dissociation products of  $CO^+$  are

$C^+(^2P)+O(^3P)$  which can form  $^2\Sigma^+$ ,  $^2\Sigma^-(2)$ ,  $^2\Pi(2)$  and others. Both the  $X^2\Sigma^+$  and  $A^2\Pi_i$  states give ground state atomic products. The  $B^2\Sigma^+$  state dissociates to  $C^+(^2P)+O(^1D)$ . Hence an unambiguous energy level diagram can be drawn for the observed states of  $CO^+$  (ref. 69, pp. 186-7).

$D_x(CO^+)=8.345$  eV,  $D_A=5.815$  eV,  $D_B=4.654$  eV, using  $O(^1D)-O(^3P)=15867.7$   $cm^{-1}=1.967$  eV [166]. These values are obtained from the ionization potentials of CO using data from Rydberg series [253, 254] and electronic transitions in  $CO^+$  [202b, 202a, 205]. The values given in section 2 are repeated for convenience: 14.013 eV, 16.544 eV, 19.671 eV.

There have been numerous other determinations of the ionization potentials of CO [129] mostly from electron impact and photoionization [65, 151, 267]. Wilkinson's review [275] compares the uncertainties in the three methods.

## 6. Vibration-Rotation Spectrum of CO

The vibration-rotation bands of CO lie at wavelengths shorter than  $5\mu$ . Observed bands of the  $\Delta v=1$  sequence include those with  $v'\leq 5$ ; the  $\Delta v=2$  sequence includes those with  $v'\leq 7$ . Overtone bands 3-0 and 4-0 have also been observed [102, 191]. Table 56 lists the observed band origins. Rotational constants derived from infrared measurements, including some isotopic data, are included in table 36.

Recently it has been proposed [203a] that calculated positions of lines of the 1-0 and 2-0 bands based on measurements by Rank et al. [197, 198, 196] and Plyler et al. [192-3, 190, 194], be adopted as secondary standard wavelengths in the infrared region. (The relative accuracy of the standards is  $\pm 0.0003$   $cm^{-1}$ ; absolute accuracy is  $\pm 0.002$   $cm^{-1}$  [203b, 196].) From these data, Rank et al. [197, 198] have derived molecular and rotational constants. Benedict [17] has recalculated slightly different values of molecular constants for the ground electronic state (table 1) which are consistent with both the infrared measurements and the data from electronic spectra. Analysis of new measurements on the fundamental band [17, 63] to  $J$  values up to at least 70 will provide a good test of the uncertainty in these constants.

### 6.1. $\Delta v=1$ Sequence

The earliest observations of CO vibration-rotation

### 5.4. Dissociation Energy of $CO^{2+}$

$CO^{2+}$  has not been observed spectroscopically but has been observed in electron impact experiments [263, 55] at 41.8 eV and 45.9 eV above the ground state of CO. Vaughen [263] has shown that the preferred dissociation products of  $CO^{2+}$  should be  $C^{2+}+O$  corresponding to a bound state with dissociation energy of 4.8 eV. This value is close to that of the isoelectronic molecule BeO (4.4 eV according to Gaydon [69]; Herzberg [99] gives 3.9 eV). Hurley and Maslen [111] have developed approximate expressions for the potential energy curve for the lowest bound state of  $CO^{2+}$  which use a scale factor  $<1$  multiplying the curve for CO,X. Dorman and Morrison [55] have assumed slightly larger  $r_e$  than for CO,X and have sketched the curve schematically. No further information is available on  $CO^{2+}$ .

tion bands were by Lowry [155] and Snow and Rideal [241]. Fine structure of the fundamental and overtone bands was first resolved by Whitcomb and Lagemann [272], and for the isotopic fundamental ( $C^{13}O^{16}$ ) by Lagemann et al., [139]. Mills and Thompson [162] have since measured the fundamental bands of  $C^{13}O^{16}$  and  $C^{12}O^{18}$  (with line positions good to  $\pm 0.03$   $cm^{-1}$ ). Plyler et al. [193], in addition to 1-0, measured some lines of the 2-1 band of these isotopes.

Benedict et al., [18] have measured the fundamental bands of the isotopic molecules  $C^{13}O^{16}$ ,  $C^{13}O^{18}$ , and  $C^{12}O^{18}$  relative to the 1-0 band of  $C^{12}O^{16}$ , in addition to making measurements on the 2-1 band of  $C^{12}O^{16}$ . (Individual line measurements were reproducible to  $\pm 0.02$   $cm^{-1}$ ; isotopic line positions relative to  $C^{12}O^{16}$  are accurate to  $\pm 0.01$   $cm^{-1}$  and are in close agreement with positions calculated from the molecular constants of Goldberg and Müller [85].)

Laser action on P branch lines in the 6-5, 7-6, 8-7, 9-8, and 10-9 vibrational bands has been reported by Patel and Kerl [185]. (See also references listed therein.)

### 6.2. Overtone Sequence and other Overtone Bands

Plyler et al., [191] have observed lines of the 2-0, 3-1, 4-2, and 5-3 bands in the overtone



region. Isotope shifts for the origin of the 2-0 bands were reported for  $\text{C}^{12}\text{O}^{16}$ — $\text{C}^{13}\text{O}^{16}$  and  $\text{C}^{12}\text{O}^{16}$ — $\text{C}^{12}\text{O}^{18}$ . Goldberg and Müller [85] identified nearly 300 lines of CO between 2.29–2.50  $\mu$  in the spectrum of the solar limb to  $J$  values above 70. The overtone sequence included 2-0 to 7-5, with measured wavelengths referred to the center of the solar disk. Derived molecular constants are in close agreement with the current best values. (See refs. 161, 165 for a listing of solar infrared CO lines.) Plyler et al., [190] have measured the same sequence. Transitions till 11-9 were seen,

but overlapping allowed precision measurements only on 2-0, 3-1, and 4-2 to  $J$  values near 70. Recently, St. Pierre [273, 218] has remeasured lines of the 2-0, 3-1, and 4-2 bands with improved precision. Rank [197] has determined the 3-0 band origin from interferometric measurement of the R(6) line and use of precisely determined molecular constants.

Herzberg and Rao [102] have photographed the 4-0 band with wavelengths of unblended lines having relative accuracy of  $\pm 0.01 \text{ cm}^{-1}$ , and absolute accuracy of  $0.03 \text{ cm}^{-1}$ .

## 7. Rotational Spectrum of CO (Microwave and Far Infrared)

The pure rotational spectrum of CO has been studied in the microwave region (wavelengths  $< 2.6 \text{ mm}$ ) and in the far infrared (100–600  $\mu$ ). Observed transitions, rotational constants, electric dipole moment,  $g$  factor, quadrupole moment  $Q$ , and quadrupole coupling constant ( $eqQ$ ) and other properties derived from the microwave spectra are summarized in tables 57–9. Calculated positions of the pure rotational lines of  $\text{C}^{12}\text{O}^{16}$  have been reported by Rao et al., [204].

### 7.1. Rotational Transitions in CO

The first microwave study of CO was that of Gilliam et al., [83] who measured the  $J=1 \leftarrow 0$  rotational transition of  $\text{C}^{12}\text{O}^{16}$  and  $\text{C}^{13}\text{O}^{16}$ . Shortly thereafter, Bedard et al., [15] measured the  $J=1 \leftarrow 0$  and  $J=2 \leftarrow 1$  transitions and obtained a value for  $D_0$  from their frequency separations. Cowan and Gordy [48] measured the  $J=3 \leftarrow 2$  line in the 0.867 mm region. The first three rotational lines were seen together, and their separations caused by centrifugal stretching were measured directly giving a value of  $D_0$  close to that of Bedard et al., [15]. Gordy and Cowan [86] later measured the three lowest rotational transitions. Recently, Jones and Gordy [122] have measured the  $4 \leftarrow 3$ ,  $5 \leftarrow 4$ , and  $6 \leftarrow 5$  transitions in the submillimeter wave region. The observed rotational lines are listed in table 58.

Rosenblum et al., [215] have made the best measurements of the  $1 \leftarrow 0$  lines for a number of isotopes of carbon monoxide. The measurement of the magnetic moment [215] made possible three corrections to  $B_e$  for the moment of inertia of the electrons: (a) for the nonspherical electron dis-

tribution, (b) wobble of the nuclei caused by rapid precession of electronic angular momentum in the molecule, and (c) the Dunham correction between  $Y_{01}$  and  $B_e$ . (The magnetic moments and rotational constants are included in table 58.)

Palik and Rao [184] have observed the pure rotational absorption lines of CO for  $J=3$  to 23 in the 100 to 600  $\mu$  region using a far infrared spectrometer with echelette gratings of 90 lines/in. and 45 lines/in. Positions of the lines from  $J=0$  to 30, calculated from the most precise infrared data [197, 85] have recently been proposed by Rao [203a, 204] as secondary standards in the far infrared.

### 7.2. Stark Effect

Burrus [36] has determined the electric dipole moment of CO to be  $0.112 \pm 0.005$  Debye from the observed Stark shift of the  $J=1 \leftarrow 0$  transition. Mizushima [163] has calculated a small correction for the polarizability Stark effect making the final value 0.114 Debye. (See table 57.) Rosenblum et al., [215] claimed to have determined the sign of the dipole moment from the relative magnetic moments as corresponding to the charge distribution  $\text{C}^-\text{O}^+$ . This is opposite in sign to that predicted from electronegativity differences alone. The molecular orbital calculations of Huo [109] and Nesbet [173] have shown that the assigned polarity  $\text{C}^-\text{O}^+$  is to be questioned. (See sec. 2.)

### 7.3. Zeeman Effect

Cox and Gordy [49] have observed the Zeeman effect in several linear and symmetric top mole-

cules. For CO the rotational  $g$  factor was found to be  $0.268 \pm 0.005$  nuclear magnetons  $\mu_N$ . In a linear molecule without nuclear coupling, the rotational magnetic moment is directed along the rotational momentum vector  $J$  but may point parallel or antiparallel as  $g_J$  is  $+$  or  $-$ . Here only the magnitude of  $g_J$  was determined, but it is normally considered negative (except for hydrides). For a magnetic field of  $10^4$  G the  $\sigma$  component separation is 4.08 MHz.

Rosenblum et al., [215] have determined the molecular magnetic moments for various isotopic species of carbon monoxide. For  $C^{12}O^{16}$   $g_J$  is  $0.2691 \pm 0.0005 \mu_N$ . (See table 58.) Burrus [37] determined somewhat lower values from observation of the Zeeman splitting of the first rotational transitions. In a field of 10080 G, the separation was  $4.115 \pm 0.025$  MHz for the  $\sigma$  components of  $J=1 \leftarrow 0$  and  $4.095 \pm 0.050$  MHz for  $J=2 \leftarrow 1$ . Neither case showed  $\pi$  splitting or

broadening which indicates no change in  $g$  factor with  $J$ . This appears to be the first check of  $J$  independence of  $g_J$ .

## 7.4. Quadrupole Hyperfine Structure

Rosenblum and Nethercot [214] have observed the hyperfine structure of the  $J=1 \leftarrow 0$  transition of  $C^{12}O^{17}$ . The observed frequencies for the fully resolved  $\Delta F = \pm 1$  transitions are given in table 59. The value of  $(eqQ)$  is  $+4.43 \pm 0.40$  MHz.  $Q$  equal to  $-0.0265 \pm 0.003$  barn is obtained from paramagnetic studies on  $O^{17}$  [123] confirming a microwave value of  $0.026 \pm 0.009$  [249].

The rotational constants  $B_0$ ,  $D_0$ , obtained from microwave measurements are given in table 58.

The electric dipole moment and hyperfine structure obtained from microwave measurements are given in tables 57 and 59, respectively.

# 8. Raman, Stark, and Zeeman Effects in Electronic Spectra of CO and $CO^+$

## 8.1. Raman Effect

Rasetti [206, 207] observed the fundamental vibrational frequency which was excited in 1 atm of CO by mercury radiation. Observation at 35 atm was made by Bhagavantam [20]. Cabannes and Rousset [38] observed the fundamental frequency and, in addition, determined the depolarization ratio  $\rho$  to be 0.29. Amaldi [1] reported the rotational Raman effect at 6 atm pressure, including the observation of 10 Stokes and 12 anti-Stokes lines.

## 8.2. Stark Effect

All experimental attempts to observe the Stark effect in the electronic spectra of CO and  $CO^+$  have yielded negative results because of insufficient resolution. Svensson [252] observed no splitting in the B-A bands of CO in fields up to 115 kV/cm using a dispersion of 4 Å/mm. Steubing [248] failed to observe splitting in the  $CO^+$  B-X bands with fields up to 67 kV/cm, and dispersion of 5-9 Å/mm. Rave [208] observed no shift in fields up to 250 kV/cm for the A-X bands ( $CO^+$ ) with dispersion of 10 Å/mm. From an assumed slit width of less than 0.02 mm and estimates of the resolving power of these authors, Kopelman

and Klemperer [134] have estimated the dipole moments of several states (table 57). To observe the Stark splitting resolution of better than 0.5  $cm^{-1}$  is necessary.

## 8.3. Zeeman Effect

The Zeeman effect, while important of itself, reveals in addition the existence of perturbations of rotational levels and the coupling case of electronic states. Zeeman patterns have been observed for three systems:  $B^1\Sigma^+ - A^1\Pi$  (CO),  $b^3\Sigma^+ - a^3\Pi$  (CO), and  $A^2\Pi_i - X^2\Sigma^+$  ( $CO^+$ ). A general review in 1934 by Crawford [51] on the Zeeman effect in diatomic spectra summarizes most of the work that has been done on CO.

Theory predicts the largest width patterns for low  $J$ , with a maximum for  $J=\Lambda$ . For large  $J$  values Zeeman displacements are generally immeasurably small, except in the neighborhood of perturbations where the lines are very sensitive to the magnetic field. Such Zeeman patterns are large and irregular.

Kemble, Mulliken, and Crawford [128] and Crawford [50] have observed symmetric patterns with displacements proportional to field strength (up to 35,000 G) for several bands of the B-A system. Thirteen patterns were observed out of a



possible 18 for the first two lines of the P, Q, and R branches. (Greater intensity was observed for the low-frequency components of the Q doublets ( $\parallel$  polarization,  $\Delta M=0$ ) and of high-frequency components of P doublets ( $\perp$  polarization,  $\Delta M=\pm 1$ ) for high J.) Large displacements due to perturbations of the  $A^1\Pi$  state were observed for  $J=23$  to 35.

Watson [268, 269] has observed large and irregular Zeeman patterns for a number of magnetically sensitive lines in the 0-0 and 1-0 B-A bands using fields up to 30,000 G. Very violent perturbations were observed, especially in the 0-0 band. Numerous extra lines appeared. (See also ref. 216.) The  $B^1\Sigma^+$  state, having no magnetic moment, is insensitive to the field.

Watson studied in detail the large displacements of lines for  $J > 23$  previously observed by Crawford [50]. In the 0-0 band, for Q(12), the displaced lines are strongly perturbed though fields of  $10^4$  G only increase their displacement (but do not produce a Zeeman pattern). A similar perturbation is observed in the 0-1 band. The state perturbing  $A^1\Pi$  is either case (a)  $a^3\Pi$  [216] or

$d^3\Delta_i$  (then thought to be  $^3\Pi_7$ ) in intermediate coupling between (a) and (b).

Schmid [236] has observed the Zeeman effect in the  $b^3\Sigma^+ - a^3\Pi$  bands in fields up to 29,000 G. Greater splitting or broadening was found near the heads than at higher J.  $R_2$  branch lines (with lower state  $^3\Pi_1$ ) are seen as sharp triplets with twice the normal component separation.  $R_1$  and  $R_3$  branch lines are weakened and broad. These effects confirmed the assignment of the  $a^3\Pi$  state to case (b) by  $N=23$ . Also, the  $b^3\Sigma^+$  state already displays a Paschen-Back effect for 15,000 G.

Schmid and Gerö [223] observed all lines split into symmetric doublets in the  $A^2\Pi_i - X^2\Sigma^+$  bands of  $CO^+$  (fields 14,000 to 28,000 G). For  $\Pi_{1/2} - \Sigma$ , all doublets reached the maximum value  $2\Delta\sigma_{normal}$  for small J. The separation was smaller for large J in this subband and for all J in the  $\Pi_{3/2} - \Sigma$  subband. Splitting was proportional to the field, and independent of vibrational quantum number. It is noteworthy that no deviation was detected to indicate perturbation of the  $A^2\Pi_{1/2}$   $v=5$  level (sec. 4.7).

## 9. Molecular Energy Levels and Potential Energy Curves

The potential energy curve for an electronic state of a diatomic molecule represents the sum of the coulomb repulsion of the nuclei and the electronic energy as a function of internuclear distance. This function (or curve) is defined when the electronic motion and nuclear motion do not interact directly, that is, away from the region of a perturbation or predissociation. Reliable potential energy curves can be calculated from the spectroscopically determined vibrational and rotational energy levels by the method of Rydberg-Klein-Rees (RKR).<sup>9</sup>

RKR curves for several states of CO have been calculated by Tobias et al. [258]. Krupenie and Weissman [138] have recalculated these using a numerical method [271] which avoids discontinuities in some integrals. Due account was taken of revised quantum numbering in some states, intermediate coupling, and better data in several instances. There is close agreement with previous calculations [258]. Potential curves for the three observed states of  $CO^+$  are included for the first time.

Table 74 lists the term values, vibrational quanta, and turning points. The potential curves are displayed in figure 1. An energy level diagram is given in figure 2.

<sup>9</sup> The foundations of this method can be traced through reference 258.

## 10. Transition Probabilities and Lifetimes (Electronic Spectra)

The integrated band intensity in emission is proportional to the relative vibrational transition probability or band strength which can be approximated by  $p_{v',v''} = R_e^2(r)q$ , where  $R_e$  is the electronic transition moment, and  $q$  is the vibrational overlap integral square. The conditions under which this equation is valid have been discussed by Fraser [67]. It is often assumed that the electronic transition moment  $R_e$  is a slowly varying function of  $r$  and that  $R_e$  can be replaced by an average  $R_e(\bar{r}_{v',v''})$ , where the  $r$ -centroid is given by

$$\bar{r}_{v',v''} = \frac{\int \psi_{v'} r \psi_{v''} d\tau}{\int \psi_{v'} \psi_{v''} d\tau}.$$

When  $R_e$  fluctuates considerably (perhaps even by a factor of 5) then the Franck-Condon factor  $q$  becomes a poor approximation to the band strength, and the validity of the Born-Oppenheimer approximation breaks down.

The origin, properties, and methods of computation of  $r$ -centroids are discussed fully by Nicholls and Jarman [180]. Detailed discussions of the concepts involved and the work done on numerous molecules can be found in a series of reviews [133, 242, 174, 175, 181, 178].

Evaluation of the Franck-Condon factor, the dominant term in the band strength, depends on the potential energy curves of the states involved. In virtually all cases (for CO and CO<sup>+</sup>) Morse functions have been assumed. Tables 60 to 72 summarize the numerical data on Franck-Condon factors, band strengths,  $r$ -centroids, and  $I_\infty$  ( $= \frac{p_{v',v''}}{\lambda_{v',v''}^4}$ ) for a number of electronic transitions in CO and CO<sup>+</sup>. Lifetimes and oscillator strengths  $f$  for several transitions have been measured (table 73). The experimental work and calculations are discussed below (including for some transitions, determination of the functional form of  $R_e$ ). The few quantitative measurements of band intensities are mentioned.

### 10.1. A<sup>1</sup>Π—X<sup>1</sup>Σ<sup>+</sup> Fourth Positive System

Franck-Condon factors  $q$  have been calculated by Nicholls [178] (table 60a), and  $r$ -centroids have been calculated by Jarman et al., [113] (table 60b).

Other less extensive calculations of  $q$  are in substantial agreement with the above values, except for some rough estimates [195] for high quantum numbers which differ by more than a factor of 10. Nicholls [176] has shown from visual band intensity observations that  $R_e$  should be only weakly dependent on  $r$  for a number of bands with  $v'=0$  to 4, so that, for these,  $p$  is well represented by  $q$ . Silverman and Lassettre [238] have reported good agreement between the experimental relative intensity for the  $v''=0$  progression observed by electron impact with those obtained from Franck-Condon factors [113] corrected for band overlap. Lassettre [141] has indicated that more recent measurements with lower energy electron bombardment (and greater resolution) give poorer agreement.

### 10.2. B<sup>1</sup>Σ<sup>+</sup>—A<sup>1</sup>Π Ångström System

Ortenberg [183] and others [150, 14, 211] have calculated Franck-Condon factors (table 61) in substantial agreement with one another, by assuming Morse functions with  $\omega_e x_e$  estimated (23–27 cm<sup>-1</sup>). These values (some with  $v'$  up to 11) are all rather uncertain because no level of the B state has been identified with certainty for  $v>1$  because of predissociation. Robinson and Nicholls [211] have calculated  $r$ -centroids,  $p$ , and  $I_\infty$  (table 61). Robinson [210] and Robinson and Nicholls [211] have measured relative intensities for the 0–0 (8.2), 0–1 (10), 0–2 (3.2), 0–3 (5.3) bands of the  $v'=0$  progression.  $R_e(r)$  is given const. (0.768 $r^2$ –1.745 $r$ –1) for 1.12 Å <  $r$  < 1.19 Å and is almost independent of  $r$  in this range.

### 10.3. C<sup>1</sup>Σ<sup>+</sup>—A<sup>1</sup>Π Herzberg System

Leskov [150] has estimated  $q$  for several bands (table 62) assuming a Morse function with  $\omega_e x_e'$  estimated as 18.07 cm<sup>-1</sup>. These values are uncertain since only the  $v=0, 1$  levels are known for C<sup>1</sup>Σ<sup>+</sup>.

### 10.4. b<sup>3</sup>Σ<sup>+</sup>—a<sup>3</sup>Π Third Positive System

The absence of any recorded bands of the  $b$ - $a$  system for  $v'>2$  prompted Barrow et al., [13] to look into their intensity distribution. From assumed Morse functions, and the use of the



Pekeris relation (ref. [99], p. 108) an estimate was made of  $q$  relative to  $q(0-0)$  and compared with their measured relative integrated band intensities. Comparison of intensities with  $q$  showed that  $R_e$  varies significantly with  $r$ . Using a crude vibrational temperature and the ratio  $\frac{q(2-0)}{q(0-0)}$ , an estimate was made of the relative intensity of the 2-0 band to be 0.023 times that of the 0-0. A very long exposure at the expected wavelength of the 2-0 band showed no sign of a violet-degraded band, though some overlap by A-X bands prevented a categorical denial of the existence of the  $v'=2$  level.

The 2-0 band is thus much weaker here than expected. If this be due to predissociation, then  $D^0(\text{CO}) < 10.94$  eV. If the perturbations of  $b$ ,  $v'=0, 1$  extend to  $v'=2$  it is not clear how this could suppress the band completely [13]. It is suggested [13] that further study of  $b$  levels be made from the  $b-a$  and  $b-X$  systems in  $\text{C}^{12}\text{O}^{16}$  and  $\text{C}^{13}\text{O}^{16}$  to clarify this contradiction with the accepted dissociation energy.

Values of  $q$ ,  $r$ -centroids,  $p$ , and  $I_\infty$  have been calculated [210, 211, 179, 256, 183] (table 63).  $R_e(r)$  is given [256] as const.  $(0.943r-1)$  for  $1.08\text{\AA} < r < 1.16\text{\AA}$ , which confirms the large fluctuation observed by Barrow et al., [13]. In view of the crude approximations (estimates of  $\omega_e x_e$ ) involved in determining the potential functions, it is suggested that the tabulated parameters be accepted with considerable reservation, since the  $b$  state is predissociated. Schwenker [237a] has measured the lifetime of the  $b-a$  transition and calculated the oscillator strength (table 73a).

#### 10.5. $a'^3\Sigma^+ - a^3\Pi$ Asundi System

Jarmain et al. [114], have calculated Franck-Condon factors assuming modified Morse functions (table 64).

#### 10.6. $a^3\Pi - X^1\Sigma^+$ Cameron System

Jarmain et al., [114] have calculated  $q$ 's assuming modified Morse functions (table 65).

#### 10.7. $b^3\Sigma^+ - X^1\Sigma^+$ Hopfield-Birge System

Nicholls [178] has calculated  $q$ 's for the first two  $v'$  progressions (table 66).

#### 10.8. $a'^3\Sigma^+ - X^1\Sigma^+$ Hopfield-Birge System

Nicholls [178] has calculated Franck-Condon factors for a large array of these transitions (table 67).

#### 10.9. $d^3\Delta_1 - a^3\Pi$ Triplet System

Pillow and Rowlatt [188] have calculated  $I_\infty$  ( $=p_{v'v''}/\lambda^4$ ) from the data of Herzberg [99] by using Pillow's distortion method and implicitly assuming that  $R_e$  is constant. Approximate mean wavelengths are quoted for the Triplet bands. Singh and Jain [239] have calculated  $q$  and  $I_\infty$  from the same data as Pillow and Rowlatt, but assuming Morse functions. The two sets of values of  $I_\infty$  diverge from one another for  $v' > 5$ , with the Morse-based values preferred. Herman and Rakotoarijimy [93] have made measurements of the relative intensity of emission bands of the  $v''=0$  progression observed in a high voltage discharge in CO diluted in xenon, and have calculated the  $N'_v$  distribution. It is shown that, in this discharge, resonance excitation due to collision with metastable xenon atoms selectively populates the  $v'=6$  level. [In the above, corrected vibrational assignments have been used.]

Singh and Jain [240] have calculated  $q$  and  $r$ -centroids after having determined that the Morse function closely approximated the Rydberg-Klein-Rees (RKR) potential over a small range of  $v$ . (Their  $q$ 's are essentially those of ref. 239.) Assuming  $R_e$  is approximately constant, they obtained  $I_\infty$  from the Franck-Condon factors and estimates of  $N'_v$  comparable to those of Herman [93]. All the above references use  $v'$  one unit lower than the revised numbering used here. This makes the numerical results somewhat uncertain. In any event, the assumption of constant  $R_e$  is untested. The values of  $q$ ,  $r$ -centroids and  $I_\infty$  are given in table 68.

#### 10.10. $A^2\Pi_t - X^2\Sigma^+$ ( $\text{CO}^+$ ) Comet Tail System

Bennett and Dalby [19] have experimentally determined the oscillator strength [ $f$ ] for this transition, which was derived from the observed spontaneous emission lifetimes of several vibrational levels of the upper state. These quantities are given in table 73. The lifetimes decrease slightly with increasing  $v'$ , but are independent of CO pressure over the range  $0.2\text{ }\mu$  to  $6\text{ }\mu$ . Constant  $R_e$  is assumed here.

Arrays of  $q$ ,  $r$ -centroids,  $p$ , and  $I_\infty$  have been calculated [212, 177, 243, 114, 41] (table 69). Robinson and Nicholls [212] have measured relative band intensities photoelectrically. Absolute variation of  $R_e(r)$  is obtained [212, 177] as  $43(-1+1.73r-0.74r^2)$  over the range  $1.10 \text{ \AA} < r < 1.21 \text{ \AA}$  (a fluctuation of 20%). The Einstein A coefficients, absolute band strengths, and band oscillator strengths were estimated (table 73) [177] for the A-X bands by combining the lifetime ( $\tau_v$ ) measurements of Bennett and Dalby [19] with the relative band intensities of Robinson and Nicholls [212] for a number of bands.

Spindler and Wentink [243] have obtained  $R_e = 1.077 \pm 0.021$  Debye independent of  $r$  between  $1.074 \leq r \leq 1.184 \text{ \AA}$  by using the lifetime measurements of Bennett and Dalby [19]. The disagreement between these two determinations of  $R_e$  is attributed [243] to possible systematic error in the intensity measurement by Robinson and Nicholls [212] due to regions of strong overlap. This discrepancy together with the nonconstant lifetimes  $\tau'_v$  from the 2-0 and 2-1, and also the 1-1 and 1-0 bands [19] makes desirable a recheck of the A-X Comet Tail system. (The calculated Einstein functions of ref. 243 differ by a factor of 2 from those of ref. 177.) The calculated absorption  $f$  numbers disagree by a factor of three.

#### 10.11. $B^2\Sigma^+ - X^2\Sigma^+(CO^+)$ First Negative System

Nicholls [177] has calculated a large array of Franck-Condon factors and  $r$ -centroids (table 70). Lawrence [143] has measured lifetimes by means of the phase shift method and obtained Einstein coefficients and oscillator strengths for this system (table 73). Schwenker [237a], using a more conventional technique, has also measured the life-

time of the B state which is about a factor of 2.5 smaller than that of Lawrence (table 73a). The discrepancy between these two values has not been explained.

#### 10.12. $B^2\Sigma^+ - A^2\Pi_i(CO^+)$ Baldet-Johnson System

Nicholls [177] has calculated Franck-Condon factors and  $r$ -centroids (table 71).

#### 10.13. Ionization of CO

Wacks [264] has calculated Franck-Condon factors for ionizing transitions from  $X^1\Sigma^+(CO)$ ,  $v=0$  to the  $X^2\Sigma^+$ ,  $A^2\Pi_i$ , and  $B^2\Sigma^+$  states of  $CO^+$  using Morse functions (table 72). In his consideration of ionization efficiency curves, it has been assumed that  $R_e(r)$  does not fluctuate. Halmann and Laulicht [88a] have since made a more extensive calculation for three isotopes of CO (in all cases assuming Morse functions).

#### 10.14. Miscellaneous

Integrated oscillator strengths for the A-X, B-X, and C-X transitions in CO have been obtained by Lassettre and Silverman [142] from inelastic electron scattering (table 73a). Absolute  $f$ -values for bands of the A-X and B-X systems have been determined by Hesser and Dressler [102a, b] from measured radiative lifetimes and previously calculated Franck-Condon factors [178] (tables 73a, c). The sums  $\sum_v f_{v'v''}$  confirm the magnitude of the integrated  $f$ -values [142].

Fox and Hickham [65] have measured the relative ionization probability to be  $\frac{I(A)}{I(X)} = 0.2$ . For the CO continuum between  $880 \text{ \AA} - 374 \text{ \AA}$ , the  $f$ -value is 2.8 [251].

## 11. Summary and Conclusion

The aims of molecular spectroscopy include the determination of energy levels, molecular constants, potential energy curves, elucidation of the electronic structure, and the determination of transition probabilities. The status of our knowledge of these properties of CO,  $CO^+$ , and  $CO^{2+}$  will now be summarized.

Precise measurements were made years ago on bands of the B-A system, and more recently on the  $a' \leftarrow X$ ,  $e \leftarrow X$ ,  $d \rightarrow a$ ,  $E \leftarrow X$  systems and the three transitions of the ion. However, for several of these, only a few vibrational levels have been carefully studied. Work is in progress on analysis of high dispersion measurements on the A-X,



$e-X$ ,  $d-X$ ,  $I-X$ , and  $a'-X$  transitions [100, 238a]. Definitive measurements have been made of the lowest pure rotational levels ( $X^1\Sigma^+$ ,  $v=0$ ) and for the lowest vibrational quanta (IR) in the electronic ground state. From these data, precise molecular constants have been determined which, together with the well-established dissociation energy, characterizes the  $X^1\Sigma^+$  state as the best known for CO. However, even for this state, the vibrational levels (and potential energy curve) are known only halfway to the dissociation limit.

Much is known about the states lying below the first dissociation limit (11 eV), the singular exceptions being the  $f^3\Sigma^+$  and  $I^1\Sigma^-$  states. For most states lying above 11 eV, including all Rydberg states, the symmetry type, multiplicity, electronic structure, and dissociation products are unknown or uncertain. No Rydberg series have been observed in emission. Quantitative data are not available for any repulsive state.

Isotopes should be used to make unambiguous all vibrational quantum assignments. Critical re-examination of perturbations is needed, especially those pertaining to  $A^1\Pi$  and  $b^3\Sigma^+$ , to unravel the overlapping structure of various

states. Terms of the  $a'^3\Sigma^+$  and  $f^3\Sigma^+$  states near the dissociation limit have not been separated. Franck-Condon factors based on Morse functions have been calculated for many transitions, but these values should be considered tentative. Quantitative intensity measurements are virtually nonexistent.

Two states of  $CO^{2+}$  have been observed only in electron impact experiment.

*Note added in proof.* Bands of the  $T\leftarrow X$  and  $S\leftarrow X$  Tanaka progressions have been reclassified as belonging to new B-X Rydberg series III and IV, respectively (Ogawa, M., Rydberg series converging to the  $B^2\Sigma^+$  state of  $CO^+$ , J. Chem. Phys. **43**, 2142-4L (1965).

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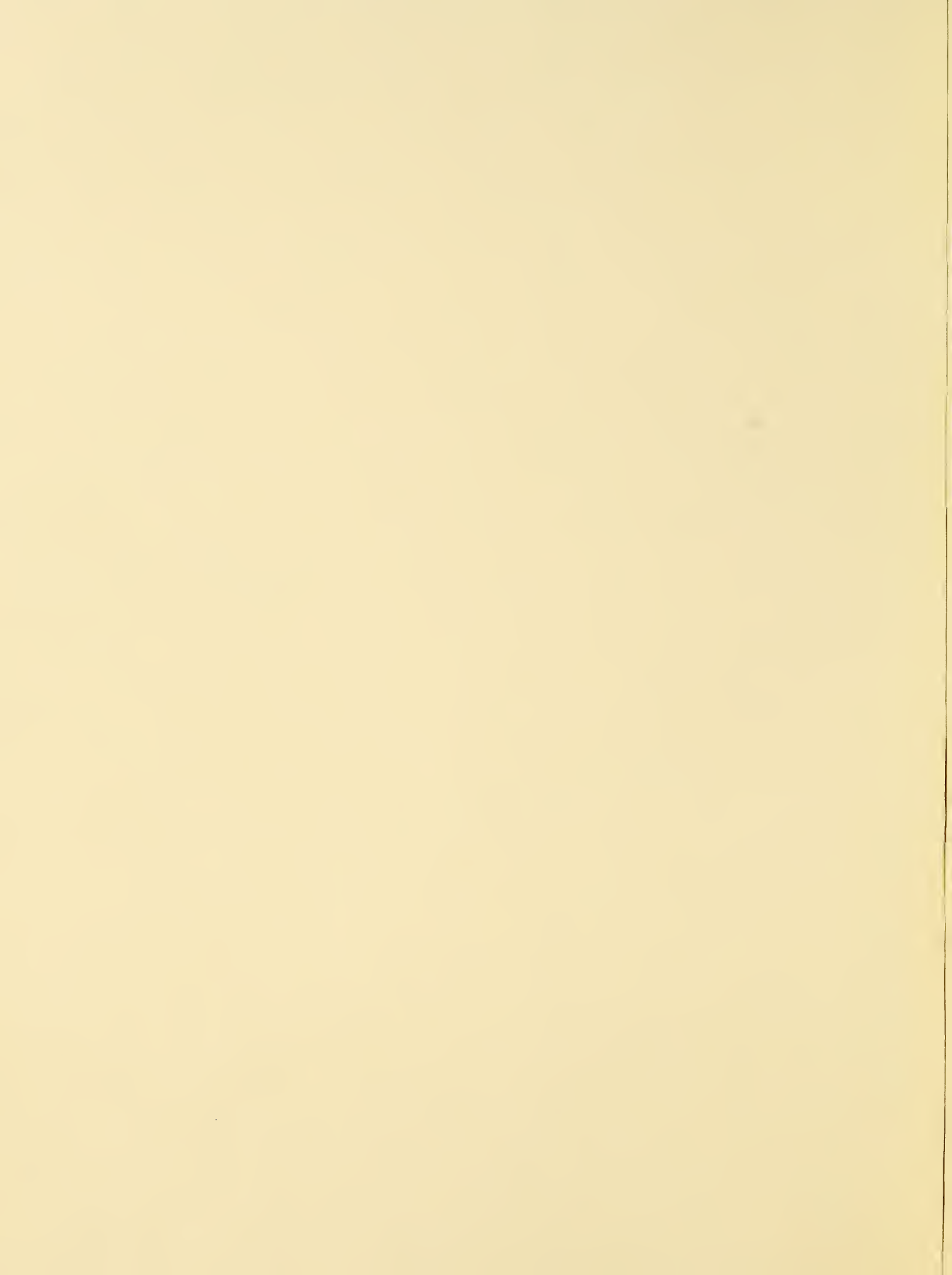


TABLE 1. Molecular constants, electron configurations, and dissociation products for the electronic states of CO and CO<sup>+</sup>

TABLE 1. Molecular constants, electron configurations, and dissociation products for the electronic states of CO and CO <sup>+</sup>																													
State	T <sub>0</sub>	T <sub>e</sub>	MO configuration		Diss. products C+O	Dissociation energy		k <sub>0</sub>	ω <sub>e</sub>	ω <sub>e</sub> x <sub>e</sub>	ω <sub>e</sub> y <sub>e</sub>	ω <sub>e</sub> z <sub>e</sub>	r <sub>e</sub>	B <sub>e</sub>	α <sub>e</sub>	γ <sub>e</sub>	D <sub>e</sub> (10 <sup>-4</sup> )	β <sub>e</sub>	Zero pt. energy	Observed transition	System name	Spectral region	Hund's coupling case	r <sub>max</sub> (Å)	References	Remarks	Reproductions of spectra		
			(3σ)(4σ)(1π)	(5σ)(2H)(6σ) Ry		D <sup>a</sup>	D <sup>b</sup>																						
								10 <sup>4</sup> dyne/cm					Å									Å							
12Σ <sup>+</sup>	158063	158884	1 4 2		<sup>1</sup> P <sub>1</sub> + <sup>1</sup> D	37538	38308	12.147	1734.18	27.927	0.3283		1.16878	1.79002	0.03025			7.75	0.22×10 <sup>-3</sup>	860.15	B→A	V	Baldet-Johnson	4240-5310		205, 100, 118-10	e <sub>g</sub> (0-0)=25220.05, e <sub>g</sub> =25143.55		
A <sup>1</sup> Π <sub>1</sub>	133136	133740	2 3 2		<sup>2</sup> P <sub>1</sub> + <sup>2</sup> P <sub>1</sub>	46002	47080	9.850	1562.06	13.532	0.0131		1.24378	1.58940	0.01942			6.60	~0	775.65	B→X	R	First Negative	3150-1800	b	10	202a, 205	e <sub>g</sub> (0-0)=45635.52, e <sub>g</sub> =45870.70	[202a]
N <sup>2</sup> Σ <sup>+</sup>	113020	113007	2 4 1		<sup>2</sup> P <sub>1</sub> + <sup>2</sup> P <sub>1</sub>	67300	68412	19.804	2214.24	15.164	-0.0007		1.11516	1.07720	0.01806	-0.000037		6.37	0.01	1103.33	A→X	V	Comet Tail	7200-3080	b	11	202b, 205	e <sub>g</sub> (0-0)=20405.47, e <sub>g</sub> =20732.16	[202b]
T	151362 H	(151760)						7.405	(1354) H	(9)										(675)	T→X	R	Tanaka	650-610	b	13	202a, 205		
H	144735 H	(144967)						10.877	(1641) H	(4.8)										(819.3)	S→X	R	Tanaka	690-670	b	2	254		
R	128860 H	(129167)						9.931	1568 H	11.6											R→X	R	Tanaka	780-730	b	3	251		
Q	128738 H	(129043)						9.805	1558 H	10.6											Q→X	R	Tanaka	780-740	b	5	251		
P	129410 H	(129712)						9.918	1507 H	13.6											P→X	R	Tanaka	790-750	b	7	251		
G <sup>1</sup> Π <sub>1</sub>	105260 H	(105790)	3 2 0 1					4.861	(1097)											(549)	G→X	R		950		0	99, 104	e <sub>g</sub> (0-0)=105200	
F <sup>1</sup> Σ <sup>+</sup>	90730 H	(90805)	4 1	3 7				18.017	2112 H	198										(1006.5)	F→X	R	Hopfield-Hirgo	1000	b	0	101, 117	e <sub>g</sub> (0-0)=99730	
II	98836																				II→X	R		1130-1010	b	0	200	e <sub>g</sub> (0-0)=98830, e <sub>g</sub> (0-0)=100830	
E <sup>1</sup> Π <sub>1</sub>	92030 H	(92045)	4 1	3pπ				18.305	[2134] H					[1.469]	[1.139]					(1067)	E→X	V	Hopfield-Hirgo	1080-1050	b	1	104, 257	e <sub>g</sub> (0-0)=92023, e <sub>g</sub> (0-0)=92030 (4)	[257]
C <sup>2</sup> Σ <sup>+</sup>	92070 H		4 1	3pσ	Pr									[1.1211]	[1.9503]						C→a	V	3A	2710-2300	b	0	77, 90	e <sub>g</sub> (0-0)=135602.1	
																					C→A	V	Herzberg	5710-3680	b	0	221	e <sub>g</sub> (0-0)=27170.17	
																					C→a'	V	Knapen	3250-2930	b	1	130, 141	e <sub>g</sub> (0-2)=30100	
																					C→X	V	Hopfield-Hirgo	1110-1060	b	1	209, 274, 270	e <sub>g</sub> (0-0)=91010	
																					F <sub>g</sub> →X	R		1310-1100	b	0	260	e <sub>g</sub> (0-0)=90860	
																					K→a	V	Kaplan	2750-2520	b	0	125	e <sub>g</sub> (0-0)=41415 (calc.)	
																					B→A	V	Ångström	6020-4120	b	221	209, 255	e <sub>g</sub> (0-0)=2217	[50, 209]
																					B→X	V	Hopfield-Hirgo	1150-1100	b	(2)	209, 255	e <sub>g</sub> (0-0)=86917.7	
																					b→a	V	Third Positive	3830-2600	b	1	101, 75, 99	e <sub>g</sub> (0-0)=85658.5	[51, 220, 55]
																					b→X	V	Hopfield-Hirgo	1190-1140	b	1	101, 99	e <sub>g</sub> (0-0)=83831	
																					f→a	R		2980-2670	b	7	78, 233		
														</															

State  
N<sup>2</sup>Σ<sup>+</sup> (a) D<sup>a</sup>=89160±160 cm<sup>-1</sup>=44.00 eV; uncertainty in D<sup>a</sup> is due to uncertainty in the multiplet components of the dissociation products [28]. Observed predissociation limit is 89935±30 cm<sup>-1</sup> [56]. Binding energies of the excited electronic states are referred to the value 89935 cm<sup>-1</sup>. (See sec. 6.1.)  
(b) H<sub>1</sub>~H<sub>2</sub>=5.66N<sup>2</sup>10<sup>-11</sup> [12]  
(c) H<sub>1</sub> has been calculated assuming atomic point masses. Corrections for nonspherical electron distribution and wobble stretching reduce this value by ~0.00001 cm<sup>-1</sup> [47, 218]. Uncertainty in the value of A probably makes any determination of r<sub>e</sub> uncertain in the sixth figure.  
(K) Herzberg [99] considers the state reported at 58820 cm<sup>-1</sup> to be spurious. Pearse and Gaydon [167] include it in their tabulations. (See sec. 7.12.)  
a<sup>1</sup>Π<sub>1</sub> (a) ω<sub>e</sub> is calculated from a precise value of ΔG(1) based on band origins [70] and ω<sub>e</sub>x<sub>e</sub> obtained from band heads [77]. The value of ω<sub>e</sub> listed by Herzberg [99] is based on band heads. There appears to be a typographical error in the value of D<sub>e</sub> listed by Herzberg. It can not be 1.8510 cm<sup>-1</sup> since it is 1.8603 cm<sup>-1</sup>.  
(b) A=41.3 cm<sup>-1</sup> [241]  
(c) ΔG(1)=724.61 cm<sup>-1</sup> from band origins [70]  
a<sup>1</sup>Σ<sup>+</sup> (a) T<sub>0</sub>=92070 is calculated [101]  
(b) A<sub>e</sub>=-5.1×10<sup>-3</sup> cm<sup>-1</sup> [101]  
(c) A+γ=-1.3 [101]  
(d) Schimid and Herz [234] list term values obtained from perturbation data to e=41. It is not certain that all these perturbations are due to the a<sup>1</sup> state, some may originate from the f state.  
(e) e<sub>g</sub>(0-0) is calculated to be 68799 cm<sup>-1</sup>.  
d<sup>1</sup>Δ<sub>1</sub> A<sub>e</sub>=-5.6 cm<sup>-1</sup>; Y<sub>e</sub>=-11.8 cm<sup>-1</sup> [60, 82]

12Σ<sup>+</sup> (a) e<sub>g</sub>(0-0) calculated from band origins with r<sub>e</sub> values [101] raised by unity [234]. Rotational constants have also been recalculated in this way.  
(b) A+γ=+0.8 cm<sup>-1</sup> [101]  
12Σ<sup>+</sup> (a) Provisional constants for this state are based on perturbation data [59] and recent observation of the I→X bands [101]. Analysis of these bands is not yet available.  
(b) Howell [105] assumed that the I state dissociates to 1D+1D which gives D<sup>a</sup>~6.5 eV (too large).  
II or 1Δ<sup>+</sup> Reality of a state at this energy (<6230 cm<sup>-1</sup>) is uncertain [101].  
Vibrational assignments for the I→X bands attributed to this state are unknown. The term value of this state is uncertain. It seems unlikely that the bands observed in ref. 78 originate from the a<sup>1</sup>Σ<sup>+</sup> state. (See sec. 3.16.)  
12Σ<sup>+</sup> Herzberg [99] lists ΔG(1)=216 cm<sup>-1</sup> which is taken from [72]. Jevois [117] lists 2214 cm<sup>-1</sup>. Schimid and Herz [234] list 2109 cm<sup>-1</sup>. Arnold's data [4] gives 2113 cm<sup>-1</sup> from band head measurements. The value listed here [2188] is taken from [246] who also obtained the rotational constants listed above. Herzberg lists D<sub>e</sub>=2.075 cm<sup>-1</sup>, ω<sub>e</sub>=0.03 cm<sup>-1</sup> which are taken from [234]. All constants for the I state are obtained from the perturbed structure and are uncertain.  
12Σ<sup>+</sup> (a) If the identification [235] of e=2 is correct then the I state would tend to dissociate to 1D+1D with an apparent D<sup>a</sup> of ~4 eV. It is, however, predissociated by a 3<sup>1</sup>+1<sup>1</sup> state.  
(b) Jevois [117] lists ω<sub>e</sub>=2161, ω<sub>e</sub>x<sub>e</sub>=50, but gives no reference. Only the ΔG(1) given above is precisely known. Using the data of [224, 235] gives approximate values ω<sub>e</sub>=2164, ω<sub>e</sub>x<sub>e</sub>=41 from which the ZPE was calculated.  
C<sup>2</sup>Σ<sup>+</sup> ΔG(1) has been determined from band origins. ω<sub>e</sub>(~2162) and ω<sub>e</sub>x<sub>e</sub>(~4.8) are estimates from which the ZPE has been obtained. ω<sub>e</sub> and D<sub>e</sub> are tentative values [72].  
F<sup>1</sup>Σ<sup>+</sup> (a) Herzberg [99] lists ω<sub>e</sub>=2134 which is the same as Jevois [117], but no reference is given for e=1. Thilker et al. [237] have observed a band head about this distance from their C→X, 0-0 band.  
(b) B<sub>e</sub> for this state is obtained only from the 11<sup>1</sup> levels (A-doubling is very small) [237].

F<sup>1</sup>Σ<sup>+</sup> (a) Though only one band has been reported in the open literature for the F→X transition (446 (Herzberg [99]) and 447 (H<sub>1</sub> [101]) cm<sup>-1</sup>), as given in the above table and listed in the table as F<sup>1</sup>Σ<sup>+</sup>.  
(b) The only vibrational bands ω<sub>e</sub>, indicate that the vibrational levels of this state converge rapidly, the 11<sup>1</sup> by Herzberg [99] being calculated to be ~0.5 eV which is a reasonable dissociation energy (1.47 eV). This would require the f state to be higher. Dissociation to 1D+1D, the only single state, seems unlikely for this energy (1.47 eV) (see above). (c) Herzberg [99] lists 446 cm<sup>-1</sup> which is taken from [72]. Jevois [117] lists 447 cm<sup>-1</sup>. Schimid and Herz [234] list 447 cm<sup>-1</sup>. Arnold's data [4] gives 447 cm<sup>-1</sup> from band head measurements. The value listed here [2188] is taken from [246] who also obtained the rotational constants listed above. Herzberg lists D<sub>e</sub>=2.075 cm<sup>-1</sup>, ω<sub>e</sub>=0.03 cm<sup>-1</sup> which are taken from [234]. All constants for the I state are obtained from the perturbed structure and are uncertain.  
(c) If Herzberg [99] lists 446 cm<sup>-1</sup> which is taken from [72]. Jevois [117] lists 447 cm<sup>-1</sup>. Schimid and Herz [234] list 447 cm<sup>-1</sup>. Arnold's data [4] gives 447 cm<sup>-1</sup> from band head measurements. The value listed here [2188] is taken from [246] who also obtained the rotational constants listed above. Herzberg lists D<sub>e</sub>=2.075 cm<sup>-1</sup>, ω<sub>e</sub>=0.03 cm<sup>-1</sup> which are taken from [234]. All constants for the I state are obtained from the perturbed structure and are uncertain.  
(c) If Herzberg [99] lists 446 cm<sup>-1</sup> which is taken from [72]. Jevois [117] lists 447 cm<sup>-1</sup>. Schimid and Herz [234] list 447 cm<sup>-1</sup>. Arnold's data [4] gives 447 cm<sup>-1</sup> from band head measurements. The value listed here [2188] is taken from [246] who also obtained the rotational constants listed above. Herzberg lists D<sub>e</sub>=2.075 cm<sup>-1</sup>, ω<sub>e</sub>=0.03 cm<sup>-1</sup> which are taken from [234]. All constants for the I state are obtained from the perturbed structure and are uncertain.  
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(c) If Herzberg [99] lists 446 cm<sup>-1</sup> which is taken from [72]. Jevois [117] lists 447 cm<sup>-1</sup>. Schimid and Herz [234] list 447 cm<sup>-1</sup>. Arnold's data [4] gives 447 cm<sup>-1</sup> from band head measurements. The value listed here [2188] is taken from [246] who also obtained the rotational constants listed above. Herzberg lists D<sub>e</sub>=2.075 cm<sup>-1</sup>, ω<sub>e</sub>=0.03 cm<sup>-1</sup> which are taken from [234]. All constants for the I state are obtained from the perturbed structure and are uncertain.  
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(c) If Herzberg [99] lists 446 cm<sup>-1</sup> which is taken from [72]. Jevois [117] lists 447 cm<sup>-1</sup>. Schimid and Herz [234] list 447 cm<sup>-1</sup>. Arnold's data [4] gives 447 cm<sup>-1</sup> from band head measurements. The value listed here [2188] is taken from [246] who also obtained the rotational constants listed above. Herz





TABLE 1. Supplement. Rydberg states

Series	State (n)	T <sub>0</sub>	$\omega_e$	$\omega_e x_e$	Reproduction of spectra
B(CO <sup>+</sup> )←X	<sup>a</sup> 4	137760	-----	-----	<b><math>\beta</math></b> series [254]
	5	148712	(1739)	(22.8)	
	6	152812	(1737)	(20.3)	
	7	154820			
Sharp	8	155943			
	9	156637			
	10	157082			
	11	157416			
	12	157627			
	Ref. [254]; (a) Ref. [107]				
Diffuse	<sup>a</sup> 4	139530			
	5	149105			
	6	153004			
	7	154943			
	8	156053			
	9	156728			
	Ref. [254]; (a) Ref. [107]				
A(CO <sup>+</sup> )←X	5	123335			
	6	127436			
	7	129431			
	8	130659			
	9	131323			
	Ref. [254]				
X(CO <sup>+</sup> )←X	6	106576	[2185]		
	7	108849	[2171]		
	8	110084	[2181]		
	9	110865	[2161]		
	10	111357	[2181]		
	11	111707	[2175]		
	12	111957	[2172]		
	13	112145	[2167]		
	14	112284	[2167]		
	Ref. [253]				

All above Rydberg series observed in absorption.

TABLE 2. Electron configurations for states of CO and CO<sup>+</sup><sup>c</sup>

	Electron configuration					Ry	States
	(1 $\sigma$ ) (2 $\sigma$ )	(3 $\sigma$ ) (4 $\sigma$ )	(1 $\pi$ ) (5 $\sigma$ )	(2 $\pi$ ) (6 $\sigma$ )			
CO <sup>+</sup>	-----	1 2	4 2	-----	-----	-----	Unobserved
	-----	2 1	4 2	-----	-----	-----	B <sup>2</sup> $\Sigma^+$
	-----	2 2	3 2	-----	-----	-----	A <sup>2</sup> $\Pi_i$
	-----	2 2	4 1	-----	-----	-----	X <sup>2</sup> $\Sigma^+$
CO	-----	-----	4 1	-----	3 or 4 <sup>a</sup>	F(1 $\Sigma^+$ or 1 $\Pi$ )	
	-----	-----	4 1	-----	3 $p\pi$	E <sup>1</sup> $\Pi$	
	-----	-----	4 1	-----	3 $p\sigma$	C <sup>1</sup> $\Sigma^+$ , c <sup>3</sup> $\Sigma^+$	
	-----	-----	4 1	-----	3s	B <sup>1</sup> $\Sigma^+$ , b <sup>3</sup> $\Sigma^+$	
	-----	-----	3 2	0 1	-----	G <sup>1</sup> $\Pi$ , <sup>3</sup> $\Pi_i$	
	-----	-----	4 1	0 1	-----	1 $\Sigma^+$ , f <sup>3</sup> $\Sigma^+$	
	-----	-----	3 2	1	-----	I <sup>1</sup> $\Sigma^-$ , 1 $\Sigma^+$ , 1 $\Delta$ , a' <sup>3</sup> $\Sigma^+$ , e <sup>3</sup> $\Sigma^-$ , d <sup>3</sup> $\Delta$ <sup>b</sup>	
	-----	-----	4 1	1	-----	A <sup>1</sup> $\Pi$ , a <sup>3</sup> $\Pi$	
	2 2	2 2	4 2	-----	-----	X <sup>1</sup> $\Sigma^+$	
	-----	-----	-----	-----	-----	-----	-----

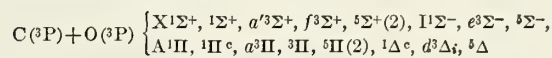
<sup>a</sup> These Rydberg states all converge to the ground state of CO<sup>+</sup> [147].<sup>b</sup> The d state is d<sup>3</sup> $\Delta_i$  [40, 137].<sup>c</sup> 1 $\Pi$  or 1 $\Delta$  possibly lies below 66230 cm<sup>-1</sup>.



TABLE 3. Band heads of the A<sup>1</sup>Π - X<sup>1</sup>Σ<sup>+</sup> Fourth Positive system (R)

(a) Emission														
λ	σ	I	v'-v''	Ref.	λ	σ	I	v'-v''	Ref.	λ	σ	I	v'-v''	Ref.
Å	cm <sup>-1</sup>				Å	cm <sup>-1</sup>				Å	cm <sup>-1</sup>			
2799.7	35708.2	9	9-22		2137.0	46779.9	5	7-14		1603.3	(62371.4)	1	9-7	L
2785.4	35890.3	8	4-18		2128.3	46971.1	8	3-11		1597.14	62611.8	8	0-1	
2742.6	36450.7	6	11-23		2113.07	47309.51	9	6-13	182	1595.60	62672.4	2	6-5	
2740.0	36486.0	4	7-20		2107.2	47440.3	7	2-10		1576.67	63425.0	8	2-2	
2712.1	36860.5	4	6-19		2089.9	47834.0	10	5-12		1560.14	64096.9	7	<sup>a</sup> 1-1	
2698.30	37049.38	6	13-24	74	2086.9	47901.5	1	1-9		1559.47	64124.4	8	4-3	
2684.0	37247.4	3	5-18		2067.71	48347.17	10	4-11	73	1545.3	64712.8	4	<sup>b</sup> 6-4	90
2680.8	37291.5	5	9-21		2046.3	48852.9	10	3-10		1544.2	64758.	-----	<sup>c</sup> 0-0	255
2662.9	37541.8	4	12-23		2034.35	49140.0	4	6-12	90	1542.34	64836.4	8	3-2	
2661.5	37561.7	4	15-25		2031.7	(49204.)	1	9-14	L	1534.2	(65180.)	2	8-5	L
2659.6	37587.9	4	4-17		2025.8	49347.8	9	2-9		1527.52	65465.5	3	5-3	
2630.0	38011.4	6	11-22		2011.8	49690.6	8	5-11		1525.75	65541.6	6	2-1	
2598.31	38475.05	4	10-21	73	2005.8	49839.8	5	1-8		1520.4	(65772.)	1	12-7	L
2594.5	38531.5	1	16-25		1990.89	50228.8	2	4-10		1515.7	65977.9	3	7-4	90
2567.8	38932.8	5	9-20		1970.05	50759.3	3	3-9		1510.4	66206.3	2	4-2	90
2556.03	39111.51	3	12-22	74	1950.07	51280.2	3	2-8		1509.63	66240.0	6	1-0	
2538.6	39379.7	4	8-19		1939.12	51569.9	1	5-10		1506.8	(66366.)	2	9-5	L
2524.1	(39606.)	4	11-21	DE	1930.70	51794.8	3	1-7		1497.8	(66764.)	3	6-3	L
2521.8	39642.7	3	14-23		1918.08	52135.4	2	4-9		1493.60	66952.3	6	3-1	
2509.9	39830.1	8	7-18		1912.8	52279.9	1	0-6	90	1488.0	(67204.)	2	8-4	L
2492.86	40102.42	8	10-20	73	1897.83	52691.6	4	3-8		1480.2	67556.6	2	5-2	90
2483.8	40248.8	3	6-17		1891.2	(52876.5)	6	6-10	L	1477.48	67682.9	3	2-0	
2463.22	40585.12	10	9-19	73	1878.33	53238.9	7	2-7		1475.4	(67778.)	1	12-6	L
2458.0	40671.5	2	5-16		1870.3	(53467.4)	3	5-9	L	1473.0	(67889.)	1	14-7	L
2433.9	41057.5	9	8-18		1859.41	53780.4	5	1-6		1463.4	68333.3	10	4-1	90
2424.20	41238.16	5	11-20	74	1850.1	54050.8	1	4-8	90	1462.17	68862.5	5	6-2	
2407.6	41552.0	7	7-17		1846.7	(54150.6)	2	7-10	L	1447.27	69095.4	1	3-0	
2394.2	41754.2	3	10-19		1841.47	54304.6	1	0-5		1443.7	(69266.)	1	8-3	L
2393.1	41773.9	4	13-21		1829.81	54650.6	4	3-7		1435.28	69672.9	7	5-1	
2381.6	41975.3	6	6-16		1825.38	54783.1	1	6-9		1425.78	70137.2	5	7-2	
2365.47	42261.95	5	9-18	74	1810.82	55223.8	8	2-6		1418.91	70475.6	1	4-0	
2356.5	42422.3	4	5-15		1804.71	55409.8	2	5-8		1414.0	(70721.)	1	11-4	L
2337.95	42759.48	7	8-17	73	1792.38	55791.7	8	1-5		1411.4	(70852.)	1	13-5	L
2332.5	42860.0	3	4-14		1784.9	56026.8	1	4-7	90	1408.86	70979.6	5	6-1	
2311.47	43249.27	8	7-16	73	1774.90	56341.3	3	0-4		1401.02	71376.5	3	8-2	
2309.7	(43282.)	4	3-13	DE	1747.20	57234.4	4	2-5		1395.7	(71649.)	2	10-3	L
2301.7	(43433.)	2	10-18	DE	1743.12	57368.4	1	5-7		1391.07	71813.8	1	5-0	
2286.1	43728.5	7	6-15		1729.25	57828.7	7	1-4		1384.00	72254.2	3	7-1	
2273.9	43963.4	3	9-17		1723.79	58011.9	2	4-6		1377.75	72582.3	2	9-2	
2272.3	43994.0	1	12-19		1712.19	58404.7	4	0-3		1373.7	72796.1	1	11-3	90
2261.7	44200.0	9	5-14		1705.16	58645.7	3	3-5		1371.8	(72897.)	2	13-4	L
2247.2	44486.6	7	8-16		1704.30	58675.2	2	6-7		1368.0	(73099.)	1	6-0	L
2238.3	44663.8	9	4-13		1684.9	59350.0	1	5-6	90	1361.3	(73459.)	2	8-1	L
2221.5	45001.4	10	7-15		1669.68	59891.8	4	1-3		1355.84	73755.0	1	10-2	
2215.8	45115.9	3	3-12		1653.02	60495.5	7	0-2		1353.6	(73877.)	1	12-3	L
2196.8	45506.4	10	6-14		1647.90	60683.4	4	3-4		1339.0	(74683.)	1	9-1	L
2194.0	(45565.)	6	2-11	DE	1630.40	61334.6	8	2-3		1335.0	(74906.)	1	11-2	L
2188.1	(45687.)	2	9-16	DE	1629.61	61364.4	8	5-5		1316.0	75986.7	1	12-2	90
2173.01	46004.67	9	5-13	73	1623.4	(61599.1)	1	10-8	L	1299.3	76965.7	1	11-1	90
2161.6	(46248.)	8	8-15	DE	1615.1	(61915.7)	2	7-6	L	1280.5	78091.4	1	12-1	90
2150.2	46492.6	8	4-12		1611.26	62063.3	3	4-4						

Air wavelengths above 2000 Å mostly from Estey [60]; vacuum wavelengths below 2000 Å mostly from Read [209]. Inaccurate head measurements by Deslandres=DE, and Lyman=L are taken from ref. 23 and indicated here by parentheses ( ). Original measurements of Read [209] and Headrick and Fox [90] were all in vacuum.

<sup>a</sup> Position calculated by Read [209], head obscured.

<sup>b</sup> Quantum assignment by Read [209]; band observed by Headrick and Fox [90].

<sup>c</sup> Absorption [255].

## (b). Absorption

$v'$	$v''=0$	$\sigma$	$I$	$v''=1$	$\sigma$	$I$
	$\lambda$			$\lambda$		
	$\text{\AA}$	$cm^{-1}$		$\text{\AA}$	$cm^{-1}$	
0	1544.2	64 758	6	1597.3	62 606	3
1	1509.8	66 234	7	1560.1	64 098	2
2	1477.5	67 682	7	1526.8	65 496	3
3	1447.4	69 089	8	1493.9	66 393	4
4	1419.1	70 467	10	1463.5	68 329	7
5	1392.6	71 808	8	1435.3	69 672	10
6	1367.7	73 115	7	1409.1	70 967	9
7	1344.1	74 399	5	1384.2	72 244	9
8	1322.1	75 637	3	1360.8	73 486	6
9	1301.4	76 840	2	1338.8	74 694	5
10	1282.0	78 003	1	1317.8	75 884	3
11	1263.6	79 139	1	1298.6	77 006	2
12	1246.2	80 244	0.7	1280.1	78 125	1
13	1229.8	81 314	0.4	1263.4	79 151	1
14	1214.3	82 352	0.3			
15	1199.8	83 347	0.1			
16	1186.1	84 310	0.07			
17	1173.2	85 237	0.05			
18	1161.2	86 118	0.03			
19						
20	1139.7	87 742	0.02			

Data taken from Tanaka, Jursa, and LeBlanc [255]. (Birge, *Nature* **124**, 182-3 (1929) referred to measurement of the 0-0 to 15-0 transitions in absorption by Hopfield and Birge; details were never published.)



TABLE 4. Deslandres table for the band origins of the  $A^1\Pi - X^1\Sigma^+$  Fourth Positive system

$v'' \backslash v'$	0	1	2	3	4	5	6	7	8	9	10	11
0	64,748.8	62,601.8	60,484.7	58,393.2	56,329.4	54,291.8	(52,266)					
1	66,231.3	64,087.6	-----	59,881.6	57,818.3	55,780.6	53,768.5	51,782.4	49,823.4	(47,887)		
2	67,674.8	65,533.1	63,416.1	61,325.2	-----	57,224.2	55,212.9	53,227.7	51,268.2	49,334.5	47,427.3	(45,529)
3	69,087.8	66,944.3	64,828.1	-----	60,674.5	58,636.3	-----	54,640.3	52,680.6	50,747.2	48,839.7	46,959.2
4	70,469.5	68,323.4	(66,199)	64,116.5	62,055.3	(59,990)	58,002.9	(56,017)	(54,041)	52,125.2	50,218.2	48,335.4
5	71,807.2	69,666.1	(67,550)	65,458.2	-----	61,356.6	(59,342)	57,360.0	55,400.9	(53,457)	51,560.0	49,678.2
6	(73,093)	70,973.2	68,855.9	(66,754)	(64,714)	62,665.1	-----	58,667.1		54,774.5	(52,866)	
7	-----	72,248.3	70,131.1	-----	(65,971)	-----	(61,908)				(54,142)	
8	-----	(73,453)	71,370.8	(69,260)	(67,198)	(65,174)						
9	-----	(74,678)	72,576.9	(70,466)	-----	(66,360)	-----	62,365.1				
10	-----	-----	73,750.0	(71,644)								
11	(79,134)	77,001	(74,902)	(72,791)	(70,717)							
12	(80,240)	78,121	75,982.2	(73,873)	(71,824)	-----	(67,773)	(65,765)				
13	(81,310)	79,147	-----	-----	(72,893)	(70,848)	(68,846)					
14	(82,348)	-----	-----	-----	-----	-----	-----	(67,884)				
15	(83,343)											
16	(84,307)											
17	(85,234)											
18	(86,115)											
19												
20	(87,739)											

$v'' \backslash v'$	12	13	14	15	16	17	18	19	20	21	22	23	24
0													
1													
2													
3	45,101.9	(43,249)											
4	46,480.7	44,649.9	(42,845)	-----	-----	(37,567)	(35,866)						
5	47,821.9	45,992.8	44,187.4	42,408.2	40,655.3	-----	(37,228)						
6	49,130.0	47,300.3	45,496.1	43,716.6	(41,962)	(40,234)	-----	(36,842)					
7	-----	-----	46,770.2	44,990.5	43,238.1	41,509.0	(39,817)	-----	(36,469)				
8	-----	-----	-----	(46,214)	44,476.5	42,748.6	41,045.5	39,367.7					
9	-----	-----	(49,196)	-----	(45,657)	43,953.9	42,251.3	40,574.0	38,920.2	(37,278)	35,693.8		
10	-----	-----	-----	-----	-----	-----	(43,403)	41,744.6	40,092.2	38,464.0			
11	-----	-----	-----	-----	-----	-----	-----	41,229.0	(39,580)	38,000.9	36,435.2		
12	-----	-----	-----	-----	-----	-----	-----	-----	-----	-----	37,532.0		
13	-----	-----	-----	-----	-----	-----	-----	(43,986)	-----	-----	-----	37,040.2	
14	-----	-----	-----	-----	-----	-----	-----	-----	-----	(41,766)	39,102.3	-----	
15	-----	-----	-----	-----	-----	-----	-----	-----	-----	-----	-----	(39,635)	
16													
17													
18													
19													
20													

This table is essentially that given by Herzberg [99] with the additions he mentioned which have large quantum numbers. This is supplemented by

origins at low quantum numbers based mostly on data of Tanaka, Jursa, and LeBlanc [255].

( ) Signifies data of much lower accuracy than the rest.

TABLE 5. *Band heads and origins of the  $A^1\Pi - X^1\Sigma^+$  Fourth Positive system of  $C^{13}O^{16}$  (R)*

(a) Emission			
$\lambda_H$	$\sigma_H$	$\sigma_0$	$v'-v''$
$\text{\AA}$	$cm^{-1}$	$cm^{-1}$	
2613.8	38247.3	38234.1	8-20
2564.5	38982	38972	10-21
2559.5	39058.8	39043.7	6-18
2535.3	39431.4	39420.0	9-20
2507.2	39873.4	39861.3	8-19
2463.2	40586	40576	10-20
2454.2	40733.4	40719.7	6-17
2429.5	41148.0	41133.4	5-16
2405.7	41555.9	41540.4	4-15
2368.3	42211	42202	10-19
2356.0	42431.3	42418.7	6-16
2340.8	42707	42697	9-18
2331.9	42870.2	42856.8	5-15
2314.3	43197	43187	8-17
2308.6	43302.2	43288.1	4-14
2264.1	44153.4	44141.8	6-15
2240.6	44616.7	44604.4	5-14
2226.7	44896.2	44886.5	8-16
2217.9	45073.2	45060.2	4-13
2201.9	45402	45391	7-15
2178.0	45900.0	45889.2	6-14
2155.0	46388.2	46376.8	5-13
2133.0	46868.6	46856.5	4-12
2120.3	47148.9	47139.2	7-14
2111.7	47340.4	47327.7	3-11
2097.0	47672.3	47662.1	6-13
2091.3	47803	47790	2-10
2074.7	48184.2	48173.5	5-12
2053.2	48689.5	48678.3	4-11
2032.4	49186.4	49174.5	3-10
2020.8	49469	49459	6-12
2012.5	49674.5	49662.1	2-9
1999.8	50005.2	49995.1	5-11
1978.8	50535.2	50524.6	4-10
1958.6	51056.7	51045.6	3-9

Data from McCulloh and Glockler [157b].

(b) Absorption (R)

$\lambda_H$	$\sigma_H$	$I$	$v'-v''$
$\text{\AA}$	$cm^{-1}$		
1510.5	66203	6	1-0
1478.8	67622	6	2-0
1449.3	68999	8	3-0
1421.4	70353	10	4-0
1395.4	71664	9	5-0
1370.8	72950	8	6-0
1347.7	74200	5	7-0
1326.0	75415	4	8-0
1305.3	76611	3	9-0
1286.2	77748	2	10-0
1267.7	78883	1	11-0
1250.6	79962	1	12-0
1234.2	81024	0.8	13-0
(1218.8)	82048	0.6	14-0
1204.4	83029	0.5	15-0
-----			
1177.8	84904	0.1	16-0
			17-0

Data from Tanaka, Jursa, and LeBlanc [255].

TABLE 6. *Band heads and origins of the  $B^1\Sigma^+ \rightarrow A^1\Pi$  Angstrom system (V)*

$\lambda_H$	$\sigma_H$	$I$	$v'-v''$	Ref.	$\sigma_0$	Ref.
$\text{\AA}$	$cm^{-1}$				$cm^{-1}$	
6620.34	15100.8	7	0-5	121	15107.2	121
6298.5	15872	2	1-6	23	-----	-----
6079.94	16443.0	9	0-4	121	16449.8	121
5817.0	17186	2	1-5	23	-----	-----
5610.15	17819.9	10	0-3	121	17827.2	121
5397.5	18522	2	1-4	23	-----	-----
5198.22	19232.0	10	0-2	121	19240.3	121
5015.0	19935	1	1-3	23	-----	-----
4835.24	20675.74	10	0-1	224	20682.1	224
4697	21284	2	1-2	23	-----	-----
4510.86	22162.65	10	0-0	224	22171.2	224
4392.93	22757.48	8	1-1	224	22763.9	224
4123.55	24244.13	7	1-0	224	24253.7	224

Data from Johnson and Asundi [121], Schmid and Gerö [224], and Birge [23]. The latter reference lists the band head measurements of Ångström and Thalén which date from 1875. Intensities are primarily those of Pearse and Gaydon [187]. Duffendack and Fox [58] list a band at 4209.0 Å as previously given, but this band is not further identified.

TABLE 7. *Band heads and origins of the  $B^1\Sigma^+ \rightarrow A^1\Pi$  system of  $C^{13}O^{16}$  (V)*

$\lambda_H$	$\sigma_H$	$v'-v''$	Ref.	$\sigma_0$	Ref.
$\text{\AA}$	$cm^{-1}$			$cm^{-1}$	
6559.4	15241.0	0-5	157b	15246.8	157b
6038.08	16556.98	0-4	157a	16563.6	157b
5583.06	17906.34	0-3	157a	17913.5	157b
5182.71	19289.55	0-2	56	19297.00	56
4828.50	20704.57	0-1	56	20712.50	56
4396.06	22741.28	1-1	56	22749.76	56

Data from McCulloh [157a], McCulloh and Glockler [157b], and Douglas and Möller [56].

TABLE 8. *Band heads and origins of the  $C^1\Sigma^+ \rightarrow A^1\Pi$  Herzberg system (V)*

$\lambda_H$	$\sigma_H$	$I$	$v'-v''$	Ref.	$\sigma_0$	Ref.
$\text{\AA}$	$cm^{-1}$				$cm^{-1}$	
5705.9	17520.9	0	0-7	97	17526.6	(*)
5318.4	18797.4	1	0-6	97	18803.4	(*)
4972.8	20103.8	2	0-5	97	20110.2	(*)
4661.68	21445.47	5	0-4	224	21452.3	(*)
4380.27	22823.22	7	0-3	224	22830.5	121
4125.03	24235.42	7	0-2	224	24243.2	121
3893.20	25678.55	7	0-1	224	25686.1	121
3680.16	27165.01	4	0-0	224	27174.17	224

Data from Johnson and Asundi [121], Schmid and Gerö [224], and Herzberg [97]. Intensities are those of Herzberg [97].

\* Calculated from the band heads.



TABLE 9. Band heads of the Hopfield-Birge systems

Transition	$\lambda$	$\sigma$	$v'-v''$	Ref.
	$\text{\AA}$	$\text{cm}^{-1}$		
(a) $B^1\Sigma^+ \leftarrow X^1\Sigma^+$	1209.80	82658.3	0-2	209
	1179.59	84775.2	0-1	209
	1150.52	86917.3	0-0	209
	1123.6	89000	1-0	255
	1098.9	91000	2-0	255

Data from Read [209] and Tanaka, Jursa, and LeBlanc [255]. The identification of the 2-0 band is uncertain.

(b) $C^1\Sigma^+ \leftarrow X^1\Sigma^+$	<sup>a</sup> 1113.89	89775.9	0-1	209
	1087.91	<sup>b</sup> 91919.1	0-0	276
	1063.09	<sup>b</sup> 94065.5	1-0	

<sup>a</sup> Band head from data of Read [209].

<sup>b</sup> Band origins from [276].

(c) $E^1\Pi \leftarrow X^1\Sigma^+$	<sup>a</sup> 1076.08	92930.03	0-0	257
	<sup>b</sup> (1053)	94967	1-0	257

Data from Tilford, Vanderslice, and Wilkinson [257].

<sup>a</sup> origin: <sup>b</sup> head.

The 0-0 band head has been observed [255] at 1076.1  $\text{\AA}$ , 92928  $\text{cm}^{-1}$ .

(d) $F(^1\Sigma^+) \leftarrow X^1\Sigma^+$	1002.7	99730	0-0	104
	<sup>a</sup> (985.8)	101450	1-0	117
	<sup>a</sup> (973.1)	102770	2-0	117

Data from Hopfield and Birge [104] and Jevons [117].

<sup>a</sup> 1-0, 2-0 band positions inferred from vibrational constants listed by Jevons [117] and Herzberg [99].

(e) $G(^1\Pi) \leftarrow X^1\Sigma^+$	949.97	105266	0-0	104
---------------------------------------	--------	--------	-----	-----

Data from Hopfield and Birge [104].

(f) $b^3\Sigma^+ \leftarrow X^1\Sigma^+$	1193.1	83812	0-0	104
	<sup>a</sup> (1134.6)	88137	2-0	255

Data from Hopfield and Birge [104] and Tanaka, Jursa, and LeBlanc [255]. Herzberg [99] lists (1192.9  $\text{\AA}$ ), 83831  $\text{cm}^{-1}$  for 0-0.

<sup>a</sup> Uncertain identification of a very weak band. No other observation of  $b$ ,  $v=2$  has been reported.

TABLE 10. Band heads and origins of the  $a'^3\Sigma^+ \leftarrow X^1\Sigma^+$  Hopfield-Birge system (R)

$\lambda_H$	$\sigma_H$	$\sigma_0$	$I$	$v'-v''$
$\text{\AA}$	$\text{cm}^{-1}$	$\text{cm}^{-1}$		
1806.3	55363	-----	0	0-0
1767.7	56570	-----	0	1-0
1731.4	57756	57750.16	1	2-0
1697.2	58922	58916.49	2	3-0
1664.8	60068	60062.64	4	4-0
1634.2	61192	61187.65	8	5-0
1605.2	62299	62293.64	20	6-0
1577.7	63384	63379.44	40	7-0
-----	-----	<sup>a</sup> (64446.9)	-----	8-0
1526.8	65495	65492.28	80	9-0
1503.2	66523	66520.39	100	10-0
-----	-----	<sup>a</sup> (67529.5)	-----	11-0
1459.4	68522	-----	80	12-0
1439.0	69494	69491.25	100	13-0
-----	-----	-----	-----	14-0
-----	-----	-----	-----	15-0
1383.2	72298	72295.14	50	16-0
1366.2	73194	-----	50	17-0
-----	-----	-----	-----	18-0
1334.5	74934	74932.04	30	19-0
1319.7	75776	75773.91	20	20-0
-----	-----	-----	-----	21-0
1291.9	77408	-----	20	22-0
1279.0	78186	78184.8	10	23-0

Data from Herzberg and Hugo [101].

<sup>a</sup> Data from Gerö and Lörinczi [80].

TABLE 11. Band heads of the  $a^3\Pi - X^1\Sigma^+$  Cameron system (R)

(a) Emission											
Branch	$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.	$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.	
	$\text{\AA}$	$cm^{-1}$				$\text{\AA}$	$cm^{-1}$				
R <sub>1</sub>	2582.2	38745	1	4-8		2415.3	41390	1	2-5		
Q <sub>2</sub>	2580.2	38745	2			2413.5	41421	2			
R <sub>2</sub>	2579.1	38762	2			2412.55	41437.36	2		201	
Q <sub>3</sub>	2577.7	38783	3			2411.4	41457	1			
R <sub>3</sub>	2575.3	38819	0			2409.2	41495	1			
	2560.2	39048	1	3-7		2394.8	41744	1	1-4		
	2558.2	39078	2			2393.1	41774	2			
	2557.1	39095	2			2392.09	41789.86	2		201	
	2555.9	39113	1			2391.1	41809	1			
	2553.3	39153	0			2388.8	41849	1			
	2538.6	39380	1	2-6		2374.96	42093.18	0	0-3	201	
	2536.7	39409				2373.3	42123	1			
	2535.6	39427	2			2372.37	42139.06	1		201	
	2534.5	39444	1			2371.2	42160	0			
	2531.9	39484	0			2369.0	42199	0			
	2517.8	39705	1	1-5					1-3		
	2516.2	39731	2			2280.5	43837	0			
	2514.9	39751	2			2279.65	43852.82	0		201	
	2513.7	39770	1			2278.5	43875	0			
	2510.9	39814	0			2277.0	43904	0			
	2497.5	40028	1	0-4					0-2		
	2495.7	40057	1			2261.2	44211				
	2494.4	40078	1			2260.07	44228.94			201	
	2492.9	40102	1			2259.5	44244				
	2491	40132			97	2257.7	44279				
	2457.8	40674	1	4-7					0-1		
	2456.0	40704	2								
	2455.0	40721	2			2157.09	46344.40			201	
	2453.9	40739	1								
	2451.8	40774	0								
	2436.2	41035	0	3-6					0-0		
	2434.5	41064	1								
	2433.4	41082	1			2061.77	48486.40			201	
	2432.4	41099	0								
	2430.3	41135	0								

Most of the data are from Cameron [39] supplemented by data from Rao [201] and Herzberg [97]. Intensities are from Cameron [39]. Wavelengths of Cameron [39] are uncertain by 0.5 Å [119].

TABLE 11. Band heads of the  $a^3\Pi - X^1\Sigma^+$  Cameron system (R)—Continued

(b) Absorption <sup>a</sup>									
Branch	$\lambda$	$\sigma$	$I$	$v'-v''$	Branch	$\lambda$	$\sigma$	$I$	$v'-v''$
	$\text{\AA}$	$cm^{-1}$				$\text{\AA}$	$cm^{-1}$		
R <sub>1</sub> -----	2064.4	48440	1	0-0					
Q <sub>2</sub> -----	2063.2	48468	4						
R <sub>2</sub> -----	2062.6	48482	5						
Q <sub>3</sub> -----	2061.7	48504	1						
R <sub>3</sub> -----	2060.6	48530	2						
	1994.2	50145	1	1-0					
	1992.7	50183	5						
	1992.1	50198	5						
	1991.3	50218	1						
	1990.2	50246	3						
	1929.7	51822	1	2-0					
	1928.2	51862	3						
	1928.0	51867	5						
	1927.4	51883	1						
	1925.8	51926	2						
	1868.2	53527	3	3-0					
	1867.8	52539	3						
	1866.1	53588	1						
	1812.8	55163	1	4-0					
	<sup>b</sup> 1765.6	56638	0	5-0					

<sup>a</sup> All vacuum wavelengths.

<sup>b</sup> Possibly incorrect branch designation.

Data taken from Tanaka, Jursa, and LeBlanc [255].

TABLE 12. Band origins of the  $a^3\Pi - X^1\Sigma^+$  Cameron system

$\sigma_0$	$v'-v''$	Ref.
$cm^{-1}$		
50188.3	1-0	79
48473.97	0-0	201
46330.59	0-1	201
44213.85	0-2	201
43838.19	1-3	201
42123.58	0-3	201
41774.18	1-4	201

Data for 1-0 band from Gerö, Herzberg, and Schmid [79]; all other data from Rao [201].



TABLE 13. Band heads of the  $b^3\Sigma^+ \rightarrow a^3\Pi$  Third Positive system (V)

Branch	$\lambda$	$\sigma$	$I$	$v'-v''$	$\lambda$	$\sigma$	$I$	$v'-v''$
	$\text{\AA}$	$cm^{-1}$			$\text{\AA}$	$cm^{-1}$		
O <sub>3</sub> -----					3134.35	31895.3	7	
P <sub>3</sub> -----					3131.47	31924.6	10	
P <sub>2</sub> -----				1-7	3129.51	31944.6	9	0-2
P <sub>1</sub> -----					3127.68	31963.3	5	
	<sup>b</sup> 3825.1	26136			3124.94	31991.4	5	
	3699.54	27022.70			3079.90	32459.2	8	
	3696.48	27045.06			3078.40	32475.0	2	
	3693.59	27066.26		<sup>a</sup> 0-5	3075.73	32503.2	4	1-3
	3690.83	27086.50			3073.49	32526.9	6	
					2977.38	33576.8	7	
					2974.49	33609.5	10	
				1-6	2972.86	33628.1	9	0-1
					2971.16	33647.1	5	
	<sup>b</sup> 3612.7	27672			2968.57	33676.5	7	
	3493.31	28618.0	5		2930.76	34110.9	3	
	3490.44	28641.6	4		2929.25	34128.5	3	
	3487.73	28663.8	9	0-4				1-2
	3485.36	28683.2	3		2924.86	34179.7	2	
	3482.23	28709.0	5					
	3419.17	29238.5	6		2833.08	35287.0	7	
	3417.49	29252.9	6		2830.15	35323.4	10	
	3414.65	29277.2	4	1-5	2828.73	35341.4	9	0-0
	3411.52	29304.1	3		2827.28	35359.2	8	
					2824.86	35389.5	8	
	3305.66	30242.5	7		2793.07	35792.9	6	
	3302.76	30269.7	8					
	3300.51	30289.7	6	0-3	2789.11	35843.3	6	1-1
	3298.43	30308.7	5		2787.37	35865.5	8	
	3295.52	30335.6	5		2785.80	35885.8	6	
	3242.14	30835.0	8		<sup>b</sup> 2665.5	37515	---	
	3240.69	30848.8	6					
	3237.74	30876.9	6	1-4	2661.87	37556.4	8	1-0
	3235.18	30901.3	8		2660.42	37577.0	8	
					2658.80	37599.8	6	

Band heads mostly from Asundi [4].

<sup>a</sup>Data on 0-5 band from Beer [16].

<sup>b</sup>Data from Duffendack and Fox [57].

TABLE 14. Band heads of the  $c^3\Sigma^+ \rightarrow a^3\Pi$  3A system (V)

Branch	$\lambda$	$\sigma$	$I$	$v'-v''$
	$\text{\AA}$	$cm^{-1}$		
O <sub>3</sub> -----	2711.40	36871.0	2	
O <sub>2</sub> -----	2710.25	36886.0	2	
P <sub>3</sub> -----				<sup>b</sup> 0-4
P <sub>2</sub> -----	2708.09	36915.0	0	
P <sub>1</sub> -----	2706.59	36935.9	5	
	2596.99	38494.66	2	
	2595.97	38509.66	1	
	2595.74	38513.16	---	<sup>a</sup> 0-3
	2594.13	38536.98	0	
	2592.55	38560.60	5	
	2489.98	40148.77	2	
	2489.09	40163.09	0	
	2488.	40169.09	0	<sup>a</sup> 0-2
	2487.29	40192.33	1	
	2485.86	40215.47	5	
	2389.81	41831.64	2	
	2389.00	41845.71	1	
	2388.	41854.18	1	<sup>a</sup> 0-1
	2387.20	41877.25	2	
	2385.93	41899.61	5	
	2295.9	43542.5	0	
				<sup>b</sup> 0-0
	2292.22	43612.4	3	

<sup>a</sup>Data from the fine structure work of Gerö [77].

<sup>b</sup>Data from vibrational measurements of Asundi [4].

For the 0-3, 0-2, and 0-1 bands, both authors' data agree to within 0.3 Å. These bands appear double-headed under low dispersion.

TABLE 15. Band heads of the  $a'^3\Sigma^+ \rightarrow a^3\Pi$  Asundi system (R)

$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.	$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.
$\text{\AA}$	$\text{cm}^{-1}$				$\text{\AA}$	$\text{cm}^{-1}$			
			4-0		6445.3	15510.9		16-5	68
							6		
8592	11636.0		6-1	4	6405.0	15608.5		11-2	68
8281.0	12072.5			4	6397	15627			158
							6		
			5-0	4	6366.9	15702.0		8-0	4
8222.5	12158.5			4	6275.0	15931.8			4
7888.0	12674.0				6257.8	15975.6	6		68
			7-1	4	6244.7	16009.2		15-4	68
7833.9	12761.8			4					
7598.0	13157.7			158	6159.5	16230.6	5		68
7574	13200		9-2	4				10-1	
7552.5	13237.0			4	6135.0	16295.4			68
				158	6127.5	16315.4			68
7359.0	13585.1		6-0	4	6119.0	16338.0	6	9-0	68
7337	13626			4					
7314.0	13668.4			4	6105.2	16375.0			4
7257.0	13776.0		11-3	158	5888.7	16977.0		16-4	
7229	13829			4	5876.2	17013.1	5		68
7210.4	13864.5				5869.4	17032.8			
			8-1	158	5855.8	17072.4		11-1	
7158	13967			4					
7134	14014			158	5836.5	17128.8	5		68
7116.5	14043.8		10-2	4	5812.8	17198.7		20-6	
					5779.6	17297.4			68
			7-0	158	5773.5	17315.7		13-2	68
7007	14267			158	5769.7	17327.1	5		68
6983	14306			158	5761	17354			158
			9-1	4	5746.0	17398.6		19-5	68
							5		
6841.5	14612.7		10-1	4				13-1	
6820	14659			158	5711.1	17504.9			
6804.0	14693.2			4					
6726.3	14862.9		9-0	4	5671.4	17627.4	6	12-2	68
6704	14913			158	5658.8	17666.7			
6685.7	14953.2			4					
6551.0	15260.6		8-0	4				11-0	
6530	15309			158	5486.3	18222.2			
6513.5	15348.5			4					



TABLE 15. Band heads of the  $a^3\Sigma^+ \rightarrow a^3\Pi$  Asundi system (R)—Continued

$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.	$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.
$\text{\AA}$	$\text{cm}^{-1}$				$\text{\AA}$	$\text{cm}^{-1}$			
5447.4	18352.3	5	12-1	68	4671.0	21402.7	6	19-3	68
					4667.1	21420.6			
5394.8	18531.2		14-2	68				14-0	68
5388.3	18553.6								
5382.6	18573.2	6					8		
5376.4	18594.6				4541.2	22014.5			
5369.9	18617.2								
5256.2	19019.9		11-0	68	4342.0	23024.4		19-2	68
		7					7		
5237.8	19086.7		17-3	68	4330.8	23083.9		20-2	68
5095.7	19618.9								
5084.1	19663.7	6					7		
					4172.0	23962.6			
5057.4	19767.5								
5045.2	19815.3	5	21-5	68			5	25-4	68
5039.0	19839.7								
5033.8	19860.2		12-0	68	4069.5	24566.1		19-1	68
4998.5	20000.4								
4993.2	20021.8				4043.8	24722.2	7		
4990.4	20032.9	1			4038.9	24752.2			
4986.4	20049.0				4034.9	24776.8			
4979.0	20078.8		18-3	68	3925.7	25466.0		22-2	68
4872.8	20516.4				3922.6	25486.1			
		6			3915.9	25529.7	8		
4858.0	20578.9								
4853.6	20597.5								
			13-0	68				20-1	68
		8					8		
4747.6	21057.4		17-2	68	3902.9	25614.7			
4700.8	21267.0								
4693.5	21300.1	7							
4683.7	21344.7								

Data from Garg [68] Asundi [4], and McLennan, Smith, and Peters [158]. Intensities from Garg [68].

Pearse and Gaydon [187] use  $v'$  three units less than that used here. Wallace [266] lists vacuum wavelengths for these bands. Rosen et al. [213], p. 66 list the band at 6685.7  $\text{\AA}$  as 7-1; it is the  $^1R_{31}$  head of the 7-0 band.

TABLE 16. Band heads of the  $d^3\Delta_1 \rightarrow a^3\Pi$  Triplet system (R)

(a) Identified bands

$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.	$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.	$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.
$\text{\AA}$	$cm^{-1}$				$\text{\AA}$	$cm^{-1}$				$\text{\AA}$	$cm^{-1}$			
(7502.1)	13326				5140.3	19448				4494.4	22244			
7515	13303		1-0	7	5128.1	19495	2		160	4478.8	22321		15-3	96
(7483)	13360				5116.2	19540				4462.9	22401			
					5070.9	19715				4466	22385			
(7231.3)	13825		3-1	7	5052.7	19786	8	7-0	96	4452.2	22455		19-5	96
(7200)	13885				5033	19864								
6964.3	14355				4996.9	20007								
6925.2	14436		2-0	7	4979.0	20079	6	9-1	96	4460.4	22413		17-4	96
6901	14487				4959.0	20160				4444.7	22492			
6464.6	15465				4935.5	20256				4454.5	22443			
6433.1	15540	10	3-0	96	4917.2	20331	2	11-2	96	4445.5	22488	1		160
6401.0	15618				4897.5	20413				4437	22531			
6383.1	15662				4880.8	20483				4405.0	22695			
6348.7	15747	1		160	4869.3	20531		13-3	96	4390.9	22768		10-0	96
6319.8	15819									4374.0	22856			
6037.0	16560				4823.5	20726				4369.9	22877			
6010.5	16633	8	4-0	96	4806.7	20798	8	8-0	96	4339.4	23038		12-1	96
5982	16712				4787	20884								
5836.9	17128				4764.8	20981			96	4343.8	23015			
5812.1	17201	2		160	4747.5	21058	5	10-1	96	4328.7	23095		14-2	96
5779	17299				4729.1	21140			160	4314.1	23173			
5670.5	17630				4716.6	21196			96	4326	23110			
5647.6	17702	6	5-0	96	4702	21262		12-2	219	4312	23185		16-3	96
5624	17776				4678	21371			219	4297	23266			
5554.1	18000				4680.3	21360				4227.2	23650			
5532.5	18070	5	7-1	96	4646.7	21515		14-3	96	4213.7	23725		11-0	96
5508	18150									4198.9	23819			
5428.3	18417				4602.6	21721				4201.5	23794			
5414.5	18464	1		160	4586.4	21797	7	9-0	96	4188.4	23869		13-1	96
5402.5	18505				4571	21871				4174.6	23948			
5351.2	18682				4556.5	21941				4182.5	23902			
5330.5	18755	5	6-0	96	4541.0	22015	5	11-1	96	4171	23968		15-2	96
5308	18834				4524.0	22098				4157	24049			
5258.3	19012				4520.7	22114				4036.4	24768			
5238.4	19085	5	8-1	96	4505.5	22189		13-2	96	4023.7	24846		16-2	96
5216	19166				4488.4	22273				4011	24924			

Data from Herman and Herman [96, 92] supplemented by data from Merton and Johnson [160] and Asundi [7]. Intensities are from Merton and Johnson [160]. For the shortest wavelength heads in each triplet, Herman and Herman [96, 92] list values which are the average of double heads reported by Merton and Johnson [160]. Numerous other band positions of reference 96 are exactly those given in ref. [160].



TABLE 16. Band heads of the  $d^6\Delta_i \rightarrow a^3\Pi$  Triplet system (R)—Continued

(b) Unassigned bands <sup>a</sup>

$\lambda$	$\sigma$	$\lambda$	$\sigma$	$\lambda$	$\sigma$
$\text{\AA}$	$cm^{-1}$	$\text{\AA}$	$cm^{-1}$	$\text{\AA}$	$cm^{-1}$
6613	15118	4460.1	22415	4074	24540
6557	15247	4441.7	22508		
6515	15345				
6323	15811	4460	22415	4055	24655
6282	15914	4444	22495	4049	24690
6248	16001	4437	22531	4034	24780
6141	16279	4428	22575	4034	24780
6122	16330	4416	22640	4023	24850
6105	16375	4406	22690	4010	24930
5889	16976	4398	22730	3928	25450
5860	17060	4378	22835	3919	25510
5839	17121	4370	22875	3912	25555
5804	17225	4314	23175	3896	25660
5777	17305	4297	23265	3893	25680
5752	17380	4288	23315	3884	25740
5157	19386	4240	23580	3885	25735
5138	19457	4229	23640	3875	25800
5112	19556	4213	23730	3870	25830
5098	19610	4344	23015	3804	26280
		4339	23040	3795	26345
		4331	23085	3788	26390
4888	20453	4188	23870	3778	26460
4870	20528	4184	23895	3767	26540
4850	20613	4176	23940		
4663	21439	4161	24025	3755	26625
4647	21513	4156	24055	3740	26730
4625	21616				
4491	22260	4115	24295		
4480	22315	4108	24335		
4463	22400	4088	24455		

<sup>a</sup> Data from Sato [219] who assigned quantum numbers to these bands. Those assignments are considered uncertain.

TABLE 17. Band origins of the  $d^3\Delta_1 \rightarrow a^3\Pi$  Triplet system (R)

$\sigma_0$	$v'-v''$
$cm^{-1}$	
15538.5	3-0
16627.4	4-0
19785.8	7-0
21794.4	9-0

Data from Gerö and Szabó [82] with revised  $v'$  numbering of Carroll [40].

TABLE 18. Band heads and origins of the  $e^3\Sigma^- \leftarrow X^1\Sigma^+$  system (R)

$\lambda$	$\sigma_H$	$\sigma_0$	$I$	$v'-v''$
$\text{\AA}$	$cm^{-1}$	$cm^{-1}$		
1542.8	64816		50	1-0
1517.7	65891	65877.82	30	2-0
1493.7	66946	66933.26	20	3-0
1471.0	67982	67970.31	50	4-0
1428.9	69992	69986.39	30	6-0
1390.1	71925	71926.47	25	8-0
1355.2	73792	73792.28	20	10-0
1338.5	74709		20	11-0
1308.1	76448?		20	13-0
1293.6	77303?		15	14-0
1241.9	80523?		5	18-0

Data from Herzberg and Hugo [101]. All  $v'$  values are one unit larger than in the original paper [238a].

(?) Band heads seen at expected positions. Bands with  $v'$  up to 23 have been observed [238a] but detailed analysis is not yet available.

TABLE 19. Band heads of the  $C^1\Sigma^+ \rightarrow a'^3\Sigma^+$  Knauss system (V)

$\lambda$	$\sigma$	$v'-v''$
$\text{\AA}$	$cm^{-1}$	
3253	30732	0-5
3138	31858	0-4
3028	33015	0-3
2925	34178	0-2

Data from Knauss [130].

TABLE 20. Band heads of the Kaplan system (R) <sup>a</sup>

Transition	$\lambda$	$\sigma$	$v'-v''$	Ref.
	$\text{\AA}$	$cm^{-1}$		
$K-a^3\Pi$ -----	2750	36350	0-3	125
	2630	38010	0-2	125
	2518	39700	0-1	125
$?-X^1\Sigma^+$ -----	<sup>a</sup> (2575)	38820	-----	126

<sup>a</sup> The reality of this band as representing a transition in CO is highly questionable.

TABLE 21. Band heads of the  $d^3\Delta_i \leftarrow X^1\Sigma^+$  system (R)<sup>a</sup>

$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$cm^{-1}$		
1617.5	61824	0 <sup>+</sup>	1-0
1588.8	62941	1	2-0
1561.3	64049	3	3-0
1535.2	65138	3	4-0
1486.6	67268	4	6-0
1464.2	68297	5	7-0
1442.8	69310	6	8-0
1402.8	71286	10	10-0
1384.2	72244	10	11-0
1366.6	73174	9	12-0
1350.0	74074	7	13-0
1334.8	74918	5	14-0
1319.6	75781	5	15-0
1305.3	76611	3	16-0
1291.9	77405	3	17-0
1279.0	78186	2	18-0
1254.4	79719	1	20-0
1243.4	80425	0 <sup>+</sup>	21-0
1232.4	81142	0 <sup>+</sup>	22-0

Data from Tanaka, Jursa, and LeBlanc [255]. Their  $v'$  numbering has been raised one unit [40]; the upper state had formerly been considered  $^3\Pi_i$  [40, 171].

<sup>a</sup> Band intensities are almost  $10^3$  times weaker than the A-X bands.

TABLE 22. Band heads of the  $e^3\Sigma^- \rightarrow a^3\Pi$  Herman system (R)

$\lambda$	$\sigma$	$I$	$v'-v''$	Ref.
$(\text{\AA})$	$(cm^{-1})$			
5428.3	18417	1	3-0	160, 92
5414.5	18464			
5402.5	18505			
5140.3	19448	2	4-0	160, 92
5128.1	19495			
5116.2	19540			
4885.0	20465.1		5-0	92
(4657.2)	21465.9		6-0	12
4454.5	22442.9		7-0	160, 92
4445.5	22488			
4437	22531			
4270.8	23408.3		8-0	92

( )=interpolated.  
See also ref. [96]. The intensities are those given by Merton and Johnson [160]. Revised  $v'$  quantum numbering is that suggested by Simmons et al., [238a].

TABLE 23. Band heads of the  $F_0^1\Sigma \rightarrow X^1\Sigma^+$  system (R)

$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$cm^{-1}$		
1309.13	76386.4	1	0-7
1276.07	78365.9	2	0-6
1244.16	80375.5	3	0-5
1213.31	82419.2	4	0-4
1183.62	84486.6	5	0-3
1154.91	86586.9	4	0-2
1127.23	<sup>a</sup> 88713.0	2	0-1
1100.49	90868.6	2	0-0

Data from Tschulanovsky [260]. The formula fitting the term values of these bands deviates considerably from that of the ground state. This makes the reality of the  $F_0$  state somewhat uncertain. In this report  $v'=0$  has been assumed.

<sup>a</sup> Band also observed in absorption by Tanaka, Jursa, and LeBlanc [255].

TABLE 24. Band heads of the  $^1\Pi \rightarrow X^1\Sigma^+$  system (R)

$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$cm^{-1}$		
1132.52	88299	1	0-5
1106.44	90380	2	0-4
1081.77	92441		0-3
1057.60	94554	4	0-2
1034.65	96651	3	0-1
1011.75	98839	1	0-0

Data from Tschulanovsky [260].

TABLE 25. Band heads of the  $f^3\Sigma^+ \rightarrow a^3\Pi$  system (R)

$\lambda$	$\sigma$	$v'-v''$
$\text{\AA}$	$cm^{-1}$	
2979.95	33547.82	0-1
2670.14	37440.15	?-0

Data from Gerö [78]. The  $v'$  numbering is uncertain.



TABLE 26. *Band heads of Tanaka systems (absorption)*

(a) $P \leftarrow X^1\Sigma^+$			
$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$\text{cm}^{-1}$		
791.08	126410	5	0-0
781.50	127959	6	1-0
772.32	129480	6	2-0
763.68	130945	5	3-0
755.24	132408	4	4-0
747.30	133815	4	5-0
739.52	135223	2	6-0
732.00	136612	3	7-0

(b) $Q \leftarrow X^1\Sigma^+$			
$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$\text{cm}^{-1}$		
776.77	128738	2	0-0
767.60	130276	5	1-0
758.80	131787	3	2-0
750.22	133294	3	3-0
742.09	134755	3	4-0
<sup>a</sup> 734.15	136212	2	5-0

TABLE 26. *Band heads of Tanaka systems (absorption)—Continued*

(c) $R \leftarrow X^1\Sigma^+$			
$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$\text{cm}^{-1}$		
776.00	128866	3	0-0
766.71	130427	3	1-0
757.86	131950	3	2-0
749.48	133426	2	3-0
741.40	134880	2	4-0
733.36	136359	4	5-0

(d) $S \leftarrow X^1\Sigma^+$			
$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$\text{cm}^{-1}$		
690.92	144735	2	0-0
683.21	146368	1	1-0
675.73	147988	1	2-0
668.44	149602	0	3-0

(e) $T \leftarrow X^1\Sigma^+$			
$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$\text{cm}^{-1}$		
647.83	154362	3	0-0
642.27	155698	2	1-0
636.88	157016	1	2-0

Data taken from Tanaka [254].

<sup>a</sup> Misprint in original paper corrected.TABLE 27.  $B^2\Sigma^+(\text{CO}^+) \leftarrow X^1\Sigma^+$  " $\beta$ " Rydberg series

$n$	$\lambda$	Sharp series		$v'-v''$	$n$	Diffuse series		$I$	$v'-v''$
		$\sigma$	$I$			$\lambda$	$\sigma$		
	$\text{\AA}$	$\text{cm}^{-1}$				$\text{\AA}$	$\text{cm}^{-1}$		
4	<sup>a</sup> 725.9	137760	-----	0-0	4	<sup>a</sup> 716.7	139530	-----	0-0
5	672.44	148712	5	0-0	5	670.67	149105	4	0-0
	664.87	150405	4	1-0		663.32	150757	2	1-0
	657.67	152053	3	2-0		656.40	152346	1	2-0
	650.81	153655	0	3-0	6	653.58	153004	2	0-0
6	654.40	152812	4	0-0		646.53	154672	1	1-0
	647.22	154507	2	1-0		639.84	156239	1	2-0
	640.35	156165	0 <sup>+</sup>	2-0	7	645.40	154943	2	0-0
7	633.80	157779	0 <sup>+</sup>	3-0		638.56	156602	1	1-0
	645.91	154820	2	0-0		631.98	158233	0 <sup>+</sup>	2-0
	639.13	156463	2	1-0	8	640.81	156053	1	0-0
8	632.43	158120	0 <sup>+</sup>	2-0					
	641.26	155943	2	0-0					
	634.41	157627	1	1-0					
9	628.05	159223	-----	2-0	9	638.05	156728	1	0-0
	638.42	156637	2	0-0					
	631.72	158298	0 <sup>+</sup>	1-0					
	636.61	157082	1	0-0					
10				0-0					
11	635.26	157416	1	0-0					
12	634.41	157627	0	0-0					

Data taken from Tanaka [254].

<sup>a</sup> Data for  $n=4$  taken from Huffman, Larrabee, and Tanaka [107]; wave-lengths to  $\pm 0.5 \text{ \AA}$ .The sharp series is possibly  $(^1\Sigma) \leftarrow X^1\Sigma^+$  [254].The diffuse series is possibly  $(^1\Pi) \leftarrow X^1\Sigma^+$  [254]. $\sigma_{\infty}(\text{sharp}) = 158692 \pm 30 \text{ cm}^{-1} = 19.6748 \text{ eV} = 630.16 \text{ \AA}$ . $\sigma_{\infty}(\text{diffuse}) = 158745 \pm 30 \text{ cm}^{-1} = 19.6814 \text{ eV} = 629.94 \text{ \AA}$ .

TABLE 28.  $A^2\Pi_1(CO^+) \leftarrow X^1\Sigma^+ \text{ "}\alpha\text{" Rydberg series}$ 

$n$	$\lambda$	$\sigma$	$I$	$v'-v''$
	$\text{\AA}$	$cm^{-1}$		
5	810.80	123335	3	0-0
	800.94	124856	3	1-0
	791.64	126320	4	2-0
	782.47	127800	3	3-0
	773.76	129239	3	4-0
6	784.70	127436	3	0-0
	775.11	129014	3	1-0
	766.09	130533	4	2-0
7	757.35	132039	3	3-0
	772.61	129431	6	0-0
	763.50	130975	7	1-0
	754.70	132503	3	2-0
8	765.35	130659	2	0-0
	756.37	132210	2	1-0
	747.74	133736	2	2-0
9	761.48	131323	1	0-0
	752.60	132873	2	1-0
	744.03	134403	1	2-0

Data from Tanaka [254].

$$\sigma_{\infty} = 133380 \text{ cm}^{-1} = 16.537 \text{ eV} = 749.74 \text{ \AA}.$$

TABLE 29.  $X^2\Sigma^+(CO^+) \leftarrow X^1\Sigma^+ \text{ Rydberg series}$ 

$n$	0-0 series		1-0 series	
	$\lambda$	$\sigma$	$\lambda$	$\sigma$
	$\text{\AA}$	$cm^{-1}$	$\text{\AA}$	$cm^{-1}$
6	938.3	106576	919.5	108761
7	918.7	108849	900.7	111020
8	908.4	110084	890.8	112265
9	902.0	110865	884.8	113026
10	898.0	111357	880.8	113538
11	895.2	111707	878.1	113882
12	893.2	111957	876.2	114129
13	891.7	112145	874.8	114312
14	890.6	112284	873.8	114451

Data from Takamine, Tanaka, and Iwata [253].

$$\sigma_{\infty}(0-0) = 113029 \text{ cm}^{-1} = 14.013 \text{ eV} = 884.73 \text{ \AA} (\pm 0.004 \text{ eV}).$$

$$\sigma_{\infty}(1-0) = 115190 \text{ cm}^{-1} = 14.281 \text{ eV} = 868.13 \text{ \AA}.$$

TABLE 30.  $B^2\Sigma^+ \rightarrow X^2\Sigma^+ \text{ First Negative system of } CO^+ (R)$ 

$\lambda$	$\sigma$	$I$	$v'-v''$	$\lambda$	$\sigma$	$I$	$v'-v''$
$\text{\AA}$	$cm^{-1}$			$\text{\AA}$	$cm^{-1}$		
3152.71	31709.6	1	8-13	<sup>a</sup> 2325.17	42994.31	9	1-2
3107.45	32171.4	2	7-12	<sup>a</sup> 2299.80	43468.68	10	0-1
3064.02	32627.4	3	6-11	2298.18	43499.2	3	4-4
3022.97	33070.4	2	5-10	2293.67	43584.8	1	10-8
<sup>c</sup> 2984.16	33500.5	2	4-9	2268.58	44066.8	3	3-3
2947.56	33916.5	1	3-8	2255.75	44317.4	1	9-7
2938.54	34020.6	1	7-11	2254.28	44346.3	2	6-5
2913.18	34316.7	1	2-7	2240.43	44620.4	4	2-2
2897.16	34506.4	3	6-10	2222.66	44977.2	4	5-4
2882.23	34685.2	2	1-6	2220.27	45025.6	0	8-6
2874.46	34778.9	0	9-12	2214.46	45143.7	5	1-1
2858.08	34978.3	4	5-9	<sup>a</sup> 2189.85	45650.99	10	0-0
2820.82	35440.2	5	4-8	2185.13	45749.6	4	7-5
2785.81	35885.6	5	3-7	2164.34	46189.0	5	3-2
<sup>a</sup> 2752.89	36314.67	6	2-6	2154.15	46407.4	4	6-4
2745.10	36417.8	1	6-9	2137.77	46763.0	6	2-1
<sup>a</sup> 2722.32	36722.43	7	1-5	2123.77	47071.2	3	5-3
2707.95	36917.4	3	5-8	2112.38	47325.0	8	1-0
<sup>a</sup> 2693.87	37110.35	2	0-4	2095.27	47711.4	5	4-2
<sup>a</sup> 2672.34	37409.27	7	4-7	2090.95	47809.6	4	7-4
<sup>b</sup> 2638.72	37885.9	8	3-6	2067.92	48342.3	1	3-1
<sup>a</sup> 2607.16	38344.46	8	2-5	2067.81	48344.8	1	9-5
<sup>a</sup> 2577.78	38781.52	10	1-4	2061.03	48503.9	3	6-3
<sup>a</sup> 2550.32	39198.99	7	0-3	2042.29	48948.9	4	2-0
2534.79	39439.2	1	4-6	2034.32	49140.6	2	8-4
<sup>a</sup> 2504.47	39916.54	10	3-5	2032.27	49190.3	1	5-2
<sup>b</sup> 2474.21	40404.7	10	2-4	2004.73	49865.9	0	4-1
<sup>a</sup> 2445.82	40873.68	10	1-3	2003.11	49906.2	0	7-3
<sup>a</sup> 2419.48	41318.69	8	0-2	1983.92	50405.3	0	9-4
2412.43	41439.4	4	4-5	1975.28	50625.7	0	6-2
2381.52	41977.1	5	3-4	1831.05	54613.4	0	9-2
2362.46	42315.8	1	6-6	1802.83	55468.4	0	8-1
<sup>a</sup> 2352.56	42493.86	6	2-3				

Most of the band heads are taken from Biskamp [24].

<sup>a</sup> Data from Rao [202a].<sup>b</sup> Data from Schmid [237]. Intensities are those of Biskamp [24].<sup>c</sup> Listed by Biskamp as uncertain 8-12 or 4-9. Biskamp's measurements are probably good to 0.05  $\text{\AA}$ ; Rao claims accuracy to 1  $cm^{-1}$ . Johnson [118] has observed an additional band (unclassified) 2883.6  $\text{\AA}$  34668.7  $cm^{-1}$  ( $I=2$ ).



TABLE 31. *Band origins of the B<sup>2</sup>Σ<sup>+</sup>→X<sup>2</sup>Σ<sup>+</sup> First Negative system of CO<sup>+</sup>*

$v'$	$v''$							
	0	1	2	3	4	5	6	7
0	45633.44	43449.62	41295.97	39172.76	37079.87	-----	-----	-----
1	-----	-----	42975.40	40852.15	38759.31	36696.77	-----	-----
2	-----	-----	-----	42478.75	40385.72	38323.19	36291.06	-----
3	-----	-----	-----	-----	-----	39898.85	(37866.7)	-----
4	-----	-----	-----	-----	-----	-----	-----	37392.19

Data mostly from Rao [202a].

( ) = Origin of 3-6 band from Coster, Brons, and Bulthuis [45].

Data of ref. [45] agree with that of ref. [202a] within 0.5 cm<sup>-1</sup>.Misprint in ref. [202a], 2-5 origin should be 38323.19 cm<sup>-1</sup>.TABLE 32. *Band heads of the A<sup>2</sup>Π<sub>i</sub>→X<sup>2</sup>Σ<sup>+</sup> Comet Tail system of CO<sup>+</sup> (R)*

$\lambda$	$\sigma$	$I$	$v'-v''$	$\lambda$	$\sigma$	$I$	$v'-v''$	$\lambda$	$\sigma$	$I$	$v'-v''$	$\lambda$	$\sigma$	$I$	$v'-v''$	
$\text{\AA}$	$cm^{-1}$			$\text{\AA}$	$cm^{-1}$			$\text{\AA}$	$cm^{-1}$			$\text{\AA}$	$cm^{-1}$			
$\Pi_{3/2}-\Sigma$	$\left\{ \begin{array}{l} 8477.69 \\ 8465.00 \end{array} \right.$	$\left\{ \begin{array}{l} 11792.43 \\ 11810.10 \end{array} \right.$	-----	a 0-4	$\left\{ \begin{array}{l} 5075.56 \\ 5071.69 \end{array} \right.$	$\left\{ \begin{array}{l} 19696.76 \\ 19711.79 \end{array} \right.$	2	b 1-1	$\left\{ \begin{array}{l} 4274.38 \\ 4271.95 \end{array} \right.$	$\left\{ \begin{array}{l} 23388.62 \\ 23401.95 \end{array} \right.$	10	c 2.0	$\left\{ \begin{array}{l} 3526.95 \\ 3525.64 \end{array} \right.$	$\left\{ \begin{array}{l} 28345.01 \\ 28355.54 \end{array} \right.$	4	7-1
$\Pi_{1/2}-\Sigma$	$\left\{ \begin{array}{l} 8388.40 \\ 8373.97 \end{array} \right.$	$\left\{ \begin{array}{l} 11917.95 \\ 11938.49 \end{array} \right.$			$\left\{ \begin{array}{l} 5043.43 \\ 5039.19 \end{array} \right.$	$\left\{ \begin{array}{l} 19822.25 \\ 19838.92 \end{array} \right.$			$\left\{ \begin{array}{l} 4251.63 \\ 4248.92 \end{array} \right.$	$\left\{ \begin{array}{l} 23513.79 \\ 23528.75 \end{array} \right.$			$\left\{ \begin{array}{l} 3511.72 \\ 3510.27 \end{array} \right.$	$\left\{ \begin{array}{l} 28467.93 \\ 28479.69 \end{array} \right.$		
	$\left\{ \begin{array}{l} 7199.91 \\ 7191.74 \end{array} \right.$	$\left\{ \begin{array}{l} 13885.23 \\ 13901.01 \end{array} \right.$	-----	a 0-3	$\left\{ \begin{array}{l} 4913.64 \\ 4909.96 \end{array} \right.$	$\left\{ \begin{array}{l} 20345.82 \\ 20361.10 \end{array} \right.$	3	b 0-0	-----	-----	2	5-2	$\left\{ \begin{array}{l} 3429.20 \\ 3427.85 \end{array} \right.$	$\left\{ \begin{array}{l} 29152.96 \\ 29164.44 \end{array} \right.$	5	6-0
	$\left\{ \begin{array}{l} 7135.54 \\ 7125.55 \end{array} \right.$	$\left\{ \begin{array}{l} 14010.49 \\ 14030.13 \end{array} \right.$			$\left\{ \begin{array}{l} 4883.53 \\ 4879.44 \end{array} \right.$	$\left\{ \begin{array}{l} 20471.27 \\ 20488.45 \end{array} \right.$			$\left\{ \begin{array}{l} 4246.16 \\ 4244.10 \end{array} \right.$	$\left\{ \begin{array}{l} 23544.06 \\ 23555.49 \end{array} \right.$			$\left\{ \begin{array}{l} 3414.58 \\ 3413.29 \end{array} \right.$	$\left\{ \begin{array}{l} 29277.78 \\ 29288.85 \end{array} \right.$		
	$\left\{ \begin{array}{l} 6246.61 \\ 6239.30 \end{array} \right.$	$\left\{ \begin{array}{l} 16004.25 \\ 16023.01 \end{array} \right.$	5	a 0-2	$\left\{ \begin{array}{l} 4868.98 \\ 4865.83 \end{array} \right.$	$\left\{ \begin{array}{l} 20532.45 \\ 20545.74 \end{array} \right.$	3	3-2	$\left\{ \begin{array}{l} 4153.82 \\ 4151.92 \end{array} \right.$	$\left\{ \begin{array}{l} 24067.44 \\ 24078.45 \end{array} \right.$	2	7-3	$\left\{ \begin{array}{l} 3366.98 \\ 3366.02 \end{array} \right.$	$\left\{ \begin{array}{l} 29691.68 \\ 29700.15 \end{array} \right.$	4	8-1
	$\left\{ \begin{array}{l} 6196.38 \\ 6189.27 \end{array} \right.$	$\left\{ \begin{array}{l} 16133.98 \\ 16152.52 \end{array} \right.$			$\left\{ \begin{array}{l} 4839.55 \\ 4836.58 \end{array} \right.$	$\left\{ \begin{array}{l} 20657.31 \\ 20669.99 \end{array} \right.$			$\left\{ \begin{array}{l} 4132.53 \\ 4130.39 \end{array} \right.$	$\left\{ \begin{array}{l} 24191.43 \\ 24203.96 \end{array} \right.$			$\left\{ \begin{array}{l} 3352.94 \\ 3351.71 \end{array} \right.$	$\left\{ \begin{array}{l} 29816.00 \\ 29826.94 \end{array} \right.$		
	$\left\{ \begin{array}{l} 5905.66 \\ 5900.40 \end{array} \right.$	$\left\{ \begin{array}{l} 16928.22 \\ 16943.31 \end{array} \right.$	3	2-3	$\left\{ \begin{array}{l} 4714.56 \\ 4711.47 \end{array} \right.$	$\left\{ \begin{array}{l} 21204.96 \\ 21218.84 \end{array} \right.$	6	b 2-1	$\left\{ \begin{array}{l} 4140.42 \\ 4138.92 \end{array} \right.$	$\left\{ \begin{array}{l} 24145.33 \\ 24154.08 \end{array} \right.$	3	4-1	$\left\{ \begin{array}{l} 3315.50 \\ 3314.18 \end{array} \right.$	$\left\{ \begin{array}{l} 30152.69 \\ 30164.70 \end{array} \right.$	2	10-2
	$\left\{ \begin{array}{l} 5862.31 \\ 5856.51 \end{array} \right.$	$\left\{ \begin{array}{l} 17053.40 \\ 17070.29 \end{array} \right.$			$\left\{ \begin{array}{l} 4686.90 \\ 4683.51 \end{array} \right.$	$\left\{ \begin{array}{l} 21330.09 \\ 21345.53 \end{array} \right.$			$\left\{ \begin{array}{l} 4119.38 \\ 4117.28 \end{array} \right.$	$\left\{ \begin{array}{l} 24268.65 \\ 24281.03 \end{array} \right.$			$\left\{ \begin{array}{l} 3302.06 \\ 3300.66 \end{array} \right.$	$\left\{ \begin{array}{l} 30275.41 \\ 30288.25 \end{array} \right.$		
	$\left\{ \begin{array}{l} 5700 \\ 5695 \end{array} \right.$	$\left\{ \begin{array}{l} 17539 \\ 17554 \end{array} \right.$	1	-----	-----	-----	1	5-3	$\left\{ \begin{array}{l} 4019.86 \\ 4017.78 \end{array} \right.$	$\left\{ \begin{array}{l} 24869.49 \\ 24882.38 \end{array} \right.$	9	c 3-0	$\left\{ \begin{array}{l} 3275.24 \\ 3273.91 \end{array} \right.$	$\left\{ \begin{array}{l} 30523.32 \\ 30535.72 \end{array} \right.$	3	7-0
	$\left\{ \begin{array}{l} 5659 \\ 5654 \end{array} \right.$	$\left\{ \begin{array}{l} 17666 \\ 17682 \end{array} \right.$			$\left\{ \begin{array}{l} 4667.98 \\ 4664.54 \end{array} \right.$	$\left\{ \begin{array}{l} 21416.55 \\ 21432.34 \end{array} \right.$			$\left\{ \begin{array}{l} 3999.71 \\ 3997.45 \end{array} \right.$	$\left\{ \begin{array}{l} 24994.75 \\ 25008.92 \end{array} \right.$			$\left\{ \begin{array}{l} 3261.53 \\ 3260.35 \end{array} \right.$	$\left\{ \begin{array}{l} 30651.15 \\ 30662.71 \end{array} \right.$		
	$\left\{ \begin{array}{l} 5698.59 \\ 5693.59 \end{array} \right.$	$\left\{ \begin{array}{l} 17543.33 \\ 17558.74 \end{array} \right.$	3	1-2	$\left\{ \begin{array}{l} 4568.75 \\ 4565.97 \end{array} \right.$	$\left\{ \begin{array}{l} 21881.71 \\ 21895.02 \end{array} \right.$	8	b 1-0	$\left\{ \begin{array}{l} 3909.90 \\ 3908.00 \end{array} \right.$	$\left\{ \begin{array}{l} 25574.7 \\ 25581.3 \end{array} \right.$	5	5-1	$\left\{ \begin{array}{l} 3223.47 \\ 3222.42 \end{array} \right.$	$\left\{ \begin{array}{l} 31013.52 \\ 31023.62 \end{array} \right.$	2	9-1
	$\left\{ \begin{array}{l} 5658.09 \\ 5652.60 \end{array} \right.$	$\left\{ \begin{array}{l} 17668.91 \\ 17686.07 \end{array} \right.$			$\left\{ \begin{array}{l} 4542.91 \\ 4539.62 \end{array} \right.$	$\left\{ \begin{array}{l} 22006.16 \\ 22022.11 \end{array} \right.$			$\left\{ \begin{array}{l} 3890.53 \\ 3888.55 \end{array} \right.$	$\left\{ \begin{array}{l} 25696.2 \\ 25709.2 \end{array} \right.$			$\left\{ \begin{array}{l} 3210.67 \\ 3209.64 \end{array} \right.$	$\left\{ \begin{array}{l} 31137.15 \\ 31147.14 \end{array} \right.$		
	$\left\{ \begin{array}{l} 5504.46 \\ 5499.86 \end{array} \right.$	$\left\{ \begin{array}{l} 18162.04 \\ 18177.23 \end{array} \right.$	5	0-1	-----	-----	5	4-2	$\left\{ \begin{array}{l} 3797.70 \\ 3795.96 \end{array} \right.$	$\left\{ \begin{array}{l} 26324.27 \\ 26336.37 \end{array} \right.$	8	c 4-0	$\left\{ \begin{array}{l} 3181.34 \\ 3180.26 \end{array} \right.$	$\left\{ \begin{array}{l} 31424.21 \\ 31434.88 \end{array} \right.$	2	11-2
	$\left\{ \begin{array}{l} 5466.74 \\ 5461.42 \end{array} \right.$	$\left\{ \begin{array}{l} 18287.36 \\ 18305.17 \end{array} \right.$			$\left\{ \begin{array}{l} 4521.00 \\ 4517.99 \end{array} \right.$	$\left\{ \begin{array}{l} 22112.80 \\ 22127.53 \end{array} \right.$			$\left\{ \begin{array}{l} 3779.77 \\ 3777.87 \end{array} \right.$	$\left\{ \begin{array}{l} 26449.19 \\ 26462.46 \end{array} \right.$			$\left\{ \begin{array}{l} 3169.07 \\ 3168.03 \end{array} \right.$	$\left\{ \begin{array}{l} 31545.87 \\ 31556.23 \end{array} \right.$		
	$\left\{ \begin{array}{l} 5430.51 \\ 5426.44 \end{array} \right.$	$\left\{ \begin{array}{l} 18409.36 \\ 18423.17 \end{array} \right.$	1	3-3	-----	-----	1	6-3	$\left\{ \begin{array}{l} 3707.10 \\ 3705.31 \end{array} \right.$	$\left\{ \begin{array}{l} 26967.59 \\ 26980.62 \end{array} \right.$	6	6-1	$\left\{ \begin{array}{l} 3136.77 \\ 3135.51 \end{array} \right.$	$\left\{ \begin{array}{l} 31870.69 \\ 31883.50 \end{array} \right.$	2	8-0
	$\left\{ \begin{array}{l} 5393.92 \\ 5389.25 \end{array} \right.$	$\left\{ \begin{array}{l} 18534.24 \\ 18550.30 \end{array} \right.$			$\left\{ \begin{array}{l} 4381.4 \\ 4378.9 \end{array} \right.$	$\left\{ \begin{array}{l} 22817.3 \\ 22830.4 \end{array} \right.$			$\left\{ \begin{array}{l} 3689.84 \\ 3688.09 \end{array} \right.$	$\left\{ \begin{array}{l} 27093.74 \\ 27106.59 \end{array} \right.$			$\left\{ \begin{array}{l} 3124.47 \\ 3123.18 \end{array} \right.$	$\left\{ \begin{array}{l} 31996.15 \\ 32009.37 \end{array} \right.$		
	$\left\{ \begin{array}{l} 5248 \\ 5244 \end{array} \right.$	$\left\{ \begin{array}{l} 19050 \\ 19064 \end{array} \right.$	$\frac{1}{2}$	2-2	$\left\{ \begin{array}{l} 4406.24 \\ 4403.26 \end{array} \right.$	$\left\{ \begin{array}{l} 22688.72 \\ 22704.07 \end{array} \right.$	4	3-1	$\left\{ \begin{array}{l} 3602.34 \\ 3600.88 \end{array} \right.$	$\left\{ \begin{array}{l} 27751.84 \\ 27763.14 \end{array} \right.$	7	c 5-0	$\left\{ \begin{array}{l} 3094.57 \\ 3093.30 \end{array} \right.$	$\left\{ \begin{array}{l} 32305.29 \\ 32318.55 \end{array} \right.$	1	10-1
	$\left\{ \begin{array}{l} 5214 \\ 5209 \end{array} \right.$	$\left\{ \begin{array}{l} 19174 \\ 19192 \end{array} \right.$			$\left\{ \begin{array}{l} 4381.38 \\ 4378.93 \end{array} \right.$	$\left\{ \begin{array}{l} 22817.45 \\ 22830.22 \end{array} \right.$			$\left\{ \begin{array}{l} 3585.96 \\ 3584.33 \end{array} \right.$	$\left\{ \begin{array}{l} 27878.58 \\ 27891.25 \end{array} \right.$			$\left\{ \begin{array}{l} 3082.71 \\ 3081.46 \end{array} \right.$	$\left\{ \begin{array}{l} 32429.57 \\ 32442.73 \end{array} \right.$		

Data of Baldet [9] supplemented by data of (a) Rao [202b], (b) Schmid and Gerö [222], and (c) Coster, Brons, and Bulthuis [45]. Intensities are from Baldet [9]. Head-forming branches are Q<sub>1</sub>, R<sub>1</sub>, R<sub>2</sub>, R<sub>21</sub>, given in order of decreasing wavelength. Johnson [118] lists a double-headed band of intensity 2 (4139.0 Å, 4123.2 Å) which may belong to this system.

TABLE 33. Band origins of the  $A^2\Pi_i-X^2\Sigma^+$  Comet Tail system of  $CO^+$

$v'$	$v''$	0	1	2	3	4
0				16070.11	13946.74	11853.88 <sup>a</sup>
1		21942.73	19758.78 <sup>b</sup>			
2		23450.59	21266.74			
3		24931.68 <sup>c</sup>				
4		26386.23				
5						
6		29214.02				
7		30588.61				
8			29752.85			
9			31074.41			
10						
11				31484.55		

<sup>a</sup> Data from Rao [202b].

<sup>b</sup> Computed from the data of Schmid and Gerö [222].

<sup>c</sup> Computed from the data of Coster, Brons, and Bulthuis [45].

<sup>d</sup> Computed from the data of Bulthuis [34a].

Data compiled by Rao [202b].

TABLE 34. Band heads of the  $B^2\Sigma^+ \rightarrow A^2\Pi_i$  Baldet-Johnson system of  $CO^+$  (V)

$\lambda$	$\sigma$	$I$	$v'-v''$	Ref
$\text{\AA}$	$\text{cm}^{-1}$			
4236.28	23598.96	3	0-1	34b
4231.42	23626.09	9		
4212.98	23729.46	7		
4209.02	23751.82	8		
4201.5	23794	1	1-2	118
4182.6	23902	1		
4179.1	23922	1		
3978.13	25130.36	4	0-0	34b
3973.31	25160.84	9		
3957.37	25262.13	6		
3953.52	25286.77	9		
3729.8	26803	3	1-0	118
3725.0	26838	9		
3711.3	26937	9		
3707.6	26964	9		
3515.8	28435	2	2-0	118
3511.7	28468	7		
3500.4	28560	3		
3496.7	28590	4		
3331.9	30004	1	3-0	118
3329.0	30030	1		
3317.9	30131	1		
3314.8	30159	1		

Data from Johnson [118] and Bulthuis [34b]. Measurements of Baldet [8] differ from the above by up to 1  $\text{\AA}$ . Intensities are those given by Johnson [118]. Head forming branches are  $P_1$ ,  $Q_1$ ,  $P_2$ , and  $Q_2$  in order of decreasing wavelength.

TABLE 35. Miscellaneous unclassified bands

(a) Band heads observed together with the  $A^1\Pi-X^1\Sigma^+$  Fourth Positive system (emission)<sup>b</sup>

$\lambda$	$\sigma$	$I$	Ref.
$\text{\AA}$	$\text{cm}^{-1}$		
2485.8	40216.4	-----	277
2299.6	43472.4	-----	277
2209.0	45255.2	1	59
2098.19	<sup>a</sup> 47644.9	-----	89
2088.0	47877	-----	59
2064.0	48434	-----	59
2042.3	48949.5	1	90
2040.0	49004	-----	59
1953.0	51203.3	5	L
1933.6	51717.0	2	23
1837.2	54430.7	1	L
1835.47	54481.9	-----	89
1827.6	54715.7	1	90
1801.9	55497.0	2	L
1772.9	56403.8	4	90
1698.8	58865.1	1	L
1688.5	59224.2	1	L
1666.7	59998.8	1	L
1492.6	66997.2	7	90
1438.7	69507	1	L
1405.5	71149	1	L
1404.0	71225	1	L
1386.4	72129	1	L
1363.3	73350.4	1	90
1343.0	74460	1	L
1228.2	81422.6	1	90
1152.9	86733.3	3	90
1025.7	97496.3	-----	90

L=Lyman (see ref. 23).

<sup>a</sup> Original measurement was in vacuum. Air wavelengths listed above.

<sup>b</sup> See sec. 3.1.

(b) Band heads observed in absorption<sup>a</sup>

$\lambda$	$\sigma$	$I$
$\text{\AA}$	$\text{cm}^{-1}$	
1569.5	63 715	1
1462.3	68 385	3
1433.3	69 769	2
1432.3	69 818	2
1422.4	70 304	8
1394.3	71 721	1
1369.5	73 019	2
1366.6	73 174	8
1352.4	73 943	1
1333.2	75 008	1
1329.3	75 228	3
1284.8	77 833	1
1218.6	82 061	0
1209.2	82 699	0
1181.8	84 617	0
1135.3	88 082	1
1134.6	88 137	1
1130.4	88 464	2
1115.4	89 654	1
1115.1	89 678	1
1111.6	89 960	1

Taken from Tanaka, Jursa, and LeBlanc [255].

<sup>a</sup> Intensities of these are approximately  $10^3$  times weaker than the strong  $A-X$  bands.

"The above bands are probably due to intercombination bands of  $CO$ . The strongest two bands appeared even at a pressure of  $\sim 1$  mm Hg" [255].



TABLE 35. *Miscellaneous unclassified bands*—Continued(c) Band heads observed with the  $a^3\Sigma^+ \rightarrow a^3\Pi$  Asundi system

$\lambda$	$\sigma$	$I$
$\text{\AA}$	$\text{cm}^{-1}$	
6960	14364	1
6939	14407	0
* 6921	14444	0
6461	15472	2
6450	15500	2
6431	15544	2
6349	15747	1
6316	15829	1
6213	16091	0
6195	16137	0
6180	16177	1
5980	16716	2
5687	17579	1
5668	17636	2
5659	17666	3

Data from McLennan, Smith, and Peters [158].

\* Band head given as 6982  $\text{\AA}$ . It is assumed the correct wavenumber was given.

(d) Band heads observed in emission by Tschulanowsky and Gassilewitsch

$\lambda$	$\sigma$
$\text{\AA}$	$\text{cm}^{-1}$
1183.62	84486.6
1154.83	86592.8
1123.63	88997.3
1107.77	90271.4
1083.24	92315.6
1066.19	93791.9
1057.60	94553.7
1052.44	95017.3
1034.68	96648.2
1032.78	96826.6
1024.70	97589.5
1018.54	98179.7
1011.78	98835.7
1008.80	99127.7
1007.77	99229.0
1007.05	99299.9
1005.26	99476.8
983.84	101643
980.65	101973
962.47	103899
961.27	104029
952.45	104992
* 948.42	105439
* 945.54	105760
943.24	106017
* 935.71	106871

Data from Tschulanowsky and Gassilewitsch [261].

\* Observed in absorption by Huffman, Larrabee, and Tanaka [107].

TABLE 35. *Miscellaneous unclassified bands*—Continued

(e) Triplet bands observed in emission

	$\lambda$	$\sigma$		$\lambda$	$\sigma$
	$\text{\AA}$	$\text{cm}^{-1}$		$\text{\AA}$	$\text{cm}^{-1}$
(a)	4100	24383	(b)	*4028	24819
	4093	24425		4024	24844
	4086	24467		*4021	24862
	*4074	24539		4017	24887
	4070	24563		*4013	24912
	*4068	23575		*4007	24949
	4063	24605		*4002	24980
	4056	24648		3997	25012
	4052	24672		*3995	25024
	4046	24709		*3993	25037
	4040	24745		3991	25049
	*4033	24788		*3987	25074
	4030	24807		*3983	25100

Data taken from Herman [95].

\* Observed in comets; band positions less accurate than the others.

(f) Unclassified absorption maxima

$\lambda$	$\sigma$	$\lambda$	$\sigma$	$\lambda$	$\sigma$
$\text{\AA}$	$\text{cm}^{-1}$	$\text{\AA}$	$\text{cm}^{-1}$	$\text{\AA}$	$\text{cm}^{-1}$
1,000.0	100,000	910.8	109,790	788.7	126,790
995.8	100,420	909.3	109,970	776.1	128,850
990.9	100,920	908.5	110,070	771.2	129,670
990.1	101,000	905.2	110,470	768.8	130,070
989.4	101,070	903.7	110,660	760.5	131,490
986.0	101,420	903.2	110,720	752.2	132,940
985.5	101,470	900.1	111,100	749.2	133,480
982.5	101,780	898.8	111,260	745.6	134,120
977.5	102,300	895.4	111,680	737.3	135,630
973.4	102,730	894.7	111,770	728.3	137,310
972.9	102,785	891.8	112,130	722.6	138,390
971.8	102,900	889.7	112,470	721.5	138,600
970.5	103,040	889.1	112,470	719.2	139,040
969.8	103,110	888.5	112,550	715.6	139,740
968.9	103,210	888.1	112,600	714.5	139,960
968.1	103,295	885.4	112,940	712.3	140,390
964.6	103,670	867.9	115,220	707.8	141,280
964.1	103,720	864.5	115,670	705.2	141,800
960.4	104,120	863.8	115,770	702.1	142,430
959.5	104,220	861.4	116,090	701.6	142,530
956.3	104,570	859.7	116,320	700.0	142,860
954.1	104,810	858.3	116,510	698.8	143,100
946.3	105,670	857.2	116,660	697.0	143,470
941.2	106,250	854.5	117,030	695.2	143,840
940.0	106,380	853.2	117,210	681.6	146,710
933.2	107,160	852.4	117,320	680.6	146,930
931.9	107,310	851.6	117,430	663.3	150,760
930.1	107,515	849.4	117,730	650.0	153,850
928.5	107,700	846.0	118,200	646.6	154,660
925.9	108,000	841.4	118,850	634.4	157,630
924.5	108,170	837.7	119,370	631.6	158,330
922.6	108,390	834.5	119,830	630.6	158,580
921.1	108,570	833.9	119,920	605.8	165,070
919.6	108,740	832.5	120,120	604.8	165,340
917.5	108,990	830.2	120,450	604.1	165,540
915.7	109,210	821.4	121,740	601.9	166,140
913.6	109,460	819.4	122,040		
912.2	109,630	795.6	125,690		

Meas. of Huffman et al., [107]. Good to  $\pm 0.5 \text{ \AA}$ .

TABLE 35. *Miscellaneous unclassified bands—Continued*

(g) Unclassified absorption bands<sup>c</sup>

$\lambda$	$\sigma$
$\text{\AA}$	$\text{cm}^{-1}$
881.3	113470
871.7 <sup>b</sup>	114720
848.6 <sup>*</sup>	117840
844.3 <sup>*</sup>	118440
839.3	119150
836.1	119600
827.2 <sup>*</sup>	120890
824.7	121260
817.8	122280
814.6	122760
807.3 <sup>*</sup>	123870
804.8 <sup>*</sup>	124250
801.1	124830
797.7	125360
794.8 <sup>*</sup>	125820
791.6	126330
789.7	126630
786.3	127180
783.1 <sup>a</sup>	127700
779.5 <sup>*</sup>	128290
774.0 <sup>*</sup>	129200
765.8	130580
757.2	132070
750.4	133260
734.2 <sup>*</sup>	136200
725.7	137800

Data of Henning [91].  
<sup>\*</sup>These bands also seen in absorption by Huffman, Larrabee, and Tanaka [107].  
<sup>a</sup> Observed also by Tanaka [254] as a closely spaced, triple-headed, and red degraded band.  
<sup>b</sup> Seen in emission by Anand [2].  
<sup>c</sup> See section 3.20.

TABLE 35. *Miscellaneous unclassified bands—Continued*

(i) Band heads observed together with the  $A^2\Pi_1 \rightarrow X^2\Sigma^+$  Comet Tail system of  $\text{CO}^+$

$\lambda$	$\sigma$	$I$
$\text{\AA}$	$\text{cm}^{-1}$	
6412	15591	1/2
6405	15608	
6361	15716	
6354	15734	
6021	16604	1/2
6015	16621	
5976	16729	
5970	16746	
5812	17201	1/2
5806	17219	
5769	17329	
5764	17344	
5321	18788	3/2
5317	18802	
5286	18913	
5281	18931	

Data of Baldet [9]; bands are red-degraded.

(j) Band heads observed together with the  $B^2\Sigma^+ \rightarrow A^2\Pi_1$  system of  $\text{CO}^+$

$\lambda$	$\sigma$	$I$
$\text{\AA}$	$\text{cm}^{-1}$	
3674.6	27206	3
3662.6	27295	4
3545.7	28195	1
3534.1	28288	1
3377.7	29597	5
3370.3	29662	7
3254.5	30718	4
3247.5	30784	5
3204.8	31194	0
3199.6	31245	0

Data from Johnson [118].

(h) Emission and absorption bands of Anand

$\lambda$	$\sigma$	$\lambda$	$\sigma$
$\text{\AA}$	$\text{cm}^{-1}$	$\text{\AA}$	$\text{cm}^{-1}$
(a) 906.2	110350	(b) 897.7	111400
886.3	112830	880.7	113550
875.7	114190	*871.6	114730
869.4	115020	866.5	115410
865.3	115570		
862.5	115940		
Emission series		Absorption series	

<sup>\*</sup>Observed previously by Henning [91].  
Data of Anand [2]; most of these bands have been observed recently, in absorption, by Huffman, Larrabee, and Tanaka [107]. See section 3.20.



TABLE 36. Rotational constants for the  $X^1\Sigma^+$  state

Isotope	$v$	$B_v$	$D_v$ ( $10^{-6}$ )
$C^{12}O^{16}$	0	1.922521	6.1193
	1	1.905014	6.1184
	2	1.887513	6.1174
	3	1.870010	6.1165
	4	1.852513	6.1155
	5	1.83514	
	6	1.81766	
	7	1.80018	
	8	1.7829	
	9	1.7629	
	10	1.7475	
	11	1.7296	
	12	1.7130	
	13	1.6957	
	14	1.6778	
	15	1.6599	
	16	1.6434	
	17	1.6256	
	18	1.6090	
	19	1.5910	
	20	1.5732	
	21	1.5561	
	22	1.5390	
	23	(1.5220)	
	24	1.5048	
$C^{13}O^{16}$	0	1.8380	
	1	1.8216	
$C^{12}O^{18}$	0	1.8319	
	1	1.8156	

- (a)  $v=0$  through 4, data from Rank et al. [198, 196] and Wiggins [273].  
 $H_0 \sim 5.8 \times 10^{-12} \text{ cm}^{-1}$  [198, 196].  $D$  values obtained from Dunham constants.  
 $D_0$  (exptl.)  $= 6.117 \times 10^{-6}$ .  
Uncertainty in  $B_0$  is  $\pm 0.0000035 \text{ cm}^{-1}$ .  
(b)  $v=5$  through 7, data from Goldberg and Møller [85].  
(c)  $v=8$  through 24, data from Schmid and Gerö [228].  
(d)  $v=23$ ,  $B$  (calc).  
(e) Data for  $v=0$  through 7 derived from vibration-rotation spectra; data for  $v \geq 8$  derived from A-X system.  
(f)  $C^{13}O^{16}$  from Plyler, Blaine, and Tidwell [193].  
(g)  $C^{12}O^{18}$  from Mills and Thompson [162].

TABLE 37. Rotational constants for the  $A^1\Pi$  state

Isotope	$v$	$B_v$
$C^{12}O^{16}$	0	1.6001
	1	1.5775
	2	1.5561
	3	1.5329
	4	1.5089
	5	1.4861
	6	<sup>a</sup> 1.4616
	7	1.4384
	8	1.4138
	9	1.3900
	10	1.3651
	11	1.3421
	12	1.3167
	13	1.2915

Data from Schmid and Gerö [228]. These values obtained from  $F(J+\frac{1}{2})$ . Work currently in progress will provide more accurate values in the near future [238a].

<sup>a</sup> Data from Onaka [182].

$C^{13}O^{16}$	1	1.509
	2	1.489
	3	1.468
	4	1.446
	5	1.424

Data from McCulloh and Glockler [157b].

TABLE 38. Rotational constants for the  $B^1\Sigma^+$  state

Isotope	$v$	$B_v$	$D_v$ ( $10^{-6}$ )
$C^{12}O^{16}$ -----	0	1.9475	$\sim 6.1$
	1	1.921	
	2	<sup>a</sup> (1.893)	

Data from Schmid and Gerö [224].

<sup>a</sup> Calculated; The identification of the  $v=2$  level is uncertain [25b].

$C^{13}O^{16}$ -----	0	1.863	5.4
	1	1.837	-----

Data from Douglas and Møller [56].

TABLE 39. Rotational constants for the  $C^1\Sigma^+$  state

$v$	$B_v$	$D_v$ ( $10^{-6}$ )
0	<sup>a</sup> 1.9436	<sup>b</sup> $\sim 5.7$
1	<sup>a</sup> 1.924	

<sup>a</sup> Data from ref. [276].

<sup>b</sup> Data from ref. [224].

TABLE 40. Rotational constants for the  $E^1\Pi$  state

Isotope	$v$	$B_0$	$D_0$ ( $10^{-6}$ )
$C^{12}O^{16}$ -----	0	1.9644	6.50
$C^{13}O^{16}$ -----	<sup>a</sup> 0	1.8773	5.87

Data from Tilford, Vanderslice, and Wilkinson [257].

<sup>a</sup>  $T_{00}=92929.62\text{ cm}^{-1}$ .TABLE 41. Rotational constants for the  $a'^3\Sigma^+$  state

$v$	$B_0$
2	1.2995
3	1.2823
4	1.2658
5	1.2478
6	1.2307
7	1.2140
8	<sup>a</sup> 1.1950
9	1.1816
10	1.1655
11	<sup>a</sup> (1.1470)
13	1.1176
16	1.0692
19	1.0206
20	1.0037
23	0.958

Data from Herzberg and Hugo [101].

<sup>a</sup> Data from Gerö and Lörinczi [80].TABLE 42. Rotational constants for the  $a^3\Pi$  state

$v$	$B_0$	$D$ ( $10^{-6}$ )	$Y$	$A$	Ref.
0	1.6803	6.1	24.7	41.5	32
1	1.6627	6.4	25.0	41.5	32
2	1.6433	6.4	25.1	41.2	32
3	1.6231	6.2	25.3	41.0	32
4	1.6026	6.3	-----	-----	54, 16
5	1.5840	6.3	-----	-----	54, 16
6	<sup>a</sup> (1.5644)	-----	-----	-----	-----
7	<sup>a</sup> (1.5449)	-----	-----	-----	-----

Data from Budó [32], Dieke and Mauchly [54], and Beer [16].

<sup>a</sup> Calculated. $A$  (average) = 41.3.TABLE 43. Rotational constants for the  $b^3\Sigma^+$  state

$v$	$B_0$
0	1.965
1	1.923

Data from Stepanov [246]. Schmid and Gerö [72, 226] list  $B_0=2.058$ ,  $B_1=2.025$ ; but see [246] and footnote  $b^3\Sigma^+$  to table 1.TABLE 44. Rotational constants for the  $c^3\Sigma^+$  state

$v$	$B_0$
0	1.9563

Data from Gerö [77].

TABLE 45. Rotational constants for the  $d^3\Delta_i$  state

$v$	$B_0$	$D_0$ ( $10^{-6}$ )	$A$
3	1.2512	8.2	-16.17
4	1.2344	9.0	-16.63
7	1.1840	11.6	-16.9
9	1.1506	10.1	-17.1

Data from Carroll [40].

TABLE 46. Rotational constants for the  $e^3\Sigma^-$  state

$v$	$B_0$
2	1.2399
3	1.2221
4	1.2048
6	1.1720
8	1.1360
10	1.1052

Data from Herzberg and Hugo [101] with  $v'$  numbering one unit larger than in that reference [238a].TABLE 47. Rotational constants for the  $E_0^1\Sigma^+$  state

$v$	$B_0$
0	1.182

Data from Tschulanovsky [260].

TABLE 48. Rotational constants for the  $^1\Pi$  state

$v$	$B_0$
0	1.139

Data from Tschulanovsky [260].

TABLE 49. Rotational constants for the  $f^3\Sigma^+$  state

$v$	$B_0$
<sup>a</sup> 0	0.83
?	0.75

Data from Gerö [78].

<sup>a</sup>  $v=0$  is assumed, but uncertain.Stepanov [247] obtained  $B_0=0.800$ ,  $B_1=0.711$ .

Both sets of values are uncertain.



TABLE 50. Rotational constants for the  $X^2\Sigma^+$  state of  $\text{CO}^+$ 

$v$	$B_v$	$D_v$ ( $10^{-6}$ )
0	1.96773	(6.36)
1	1.94871	(6.37)
2	1.92960	(6.38)
3	1.91039	6.39
4	1.89107	(6.40)
5	1.87178	(6.41)
6	1.85244	(6.42)
7	1.8328	(6.43)

Data from Rao [202a]. All  $D_v$  values are estimated, except for  $v=3$ .TABLE 51. Rotational constants for the  $A^2\Pi_i$  state of  $\text{CO}^+$ 

$v$	$B_v$
0	1.57971
1	1.56022
2	1.54070
3	1.52152
4	1.50203
5	<sup>a</sup> (1.48259)
6	1.46319
7	1.44375
8	1.42409
9	1.40509
10	<sup>a</sup> (1.38549)
11	1.36515

Data from Rao [202b].  $D$  is estimated as  $\sim 6.60 \times 10^{-6}$ ;  $A = -117.5$ .

• Calculated.

TABLE 52. Rotational constants for the  $B^2\Sigma^+$  state of  $\text{CO}^+$ 

$v$	$B_v$	$D_v$ ( $10^{-6}$ )
0	1.78480	7.89
1	1.75460	8.06
2	1.72424	8.27
3	1.69409	8.56
4	1.6636	8.5

Data from Rao [202a].

TABLE 53. Doublet splitting constants for states of  $\text{CO}^+$ 

B-X transition		$v$	$\gamma$	
$v'-v''$	$ \gamma'-\gamma'' $		( $X^2\Sigma^+$ )	( $B^2\Sigma^+$ )
0-1	0.0093	0	-----	0.0192
0-2	.0081	1	0.0285	
0-3	.0105	2	.0273	
1-2	.0098	3	.0297	
1-3	.0084			
1-4	.0100			
1-5	.0093			
2-4	.0090			
2-5	.0090			
2-6	.0088			
3-5	.0079			

Data on  $|\gamma'-\gamma''|$  from Rao [202a].Data on  $\gamma$  from Rao and Sarma [205].TABLE 54. Perturbations of the electronic states of CO and  $\text{CO}^+$ 

(a) $A^1\Pi$ state					
$v$	Max. pert. (J)	Perturbing term	Perturber $B_v$	Ref.	Comments
0	$\sim 1$	$^1\Pi$ (or $^1\Delta$ )	$cm^{-1}$	44, 71	See fig. 1 of ref. [44] <sup>a</sup> . Reality of perturbation disputed [140]. Extra lines $J=11$ to 14; B-A, 0-0 band [216]. <sup>c</sup> Q branch line magnetically sensitive. Schmid and Gerö [225] called perturber $^3\Pi$ at 65002 $cm^{-1}$ , $B_v \sim 1.37$ . Not singlet [268]. Only Q branch perturbed.
	8, 10, 17	$^3\Pi$	-----	268-9 <sup>b</sup>	
	9, 12, 16	$e^3\Sigma^-, v=1$	1.247-1.254	216, 44, 225, 262	
	25	-----	-----	50	
	27, 31, 34-5	$d^3\Delta_i, v=4$	-----	44, 216	
	28-9	$^3\Pi$ or $^3\Delta_i$	-----	268, 216	
	>30	$^3\Pi$	-----	44, 71	
1	$\sim 1$	$^1\Pi$	-----	225, 71, 216	All branches crossed at the same J. Perturber not $\Sigma$ . Q branch, A-X band. Perturber probably $a^3\Sigma^+$ . Perturbing term estimated at 66499 $cm^{-1}$ ; $a^3\Sigma^+, v=9$ is at 66575 $cm^{-1}$ . Ref. [43] estimates perturber $B_v$ to be $>1.59$ . [44] lists this perturbation, Gerö [74] identified the perturber. The reality of this perturbation is not established. It is overlapped by a $^3\Sigma^+$ perturbation. [262] believes this to be a $^3\Sigma$ perturbation with $B_v \sim 1.155$ . Q branch; see also ref. [52]. P, R branch; extra line R-X <sup>a</sup> =5.63 $cm^{-1}$ . P branch. Q branch; extra line, R-X=-4.07 $cm^{-1}$ . Q branch; A-X. See also [262, 44, 52, 234b]. P, R branches.
	$\sim 1$	$^3\Pi$ (or $^3\Delta$ )	-----	216	
	8	$^2\Sigma$	1.155	262, 269	
	23-4, 26, 29	$a^3\Sigma^+, v=9$	1.15-1.18	216, 44, 43	
	25-6 <sup>b</sup>	$^1\Sigma^-$	1.50	262	
				262, 44	
				225, 74	
	26	$^1\Sigma^+$	-----	43, 209	
	26	$^3\Sigma$	-----	44	
	28	-----	-----	44	
	29 <sup>b</sup>	-----	-----	43	
	34-5	$I^1\Sigma^-$	1.46	225, 43	
	>35	$a^3\Pi$	-----	216	
	44	$e^3\Sigma^-$	-----	74	

<sup>a</sup> R-X=regular-extra.<sup>b</sup> Extra lines R(26), Q(29), B-A [43,44]; P(26), Q(29), A-X [262].

TABLE 54. *Perturbations of the electronic states of CO and CO<sup>+</sup>*—Continued

$v$	Max. pert. (J)	Perturbing term	Perturber B.	Ref.	Comments
2	7	$1^1\Sigma^-$	$cm^{-1}$	73	Q branch; line doubled; no further details given. See also [52]). P branch shows trace of a perturbation. Perturber estimated at 67953 $cm^{-1}$ (far from the experimental value [101]).
	19			262	
	25, 28-9, 33,	$e^3\Sigma^-, v=4$	1.22	216, 44, 262, 225	
	35			216	
	Extra lines:		P, R(25) P, R(33) P(25, 26)	[216] [216] [262]	
3	27-8, 31, 34	$a' ^3\Sigma^+, v=12$	1.14	216, 44, 262, 225	Q, P, Q branches crossed by $^3\Sigma$ sublevels $J=N+1, N, N-1$ . P branch. Q branch. Schmid and Gerö [225] refer to this as the $^1\Sigma$ perturbation of Coster and Brons. See also [234b, 52].
	29	$^3\Sigma$	1.120	262	
	38	$1^1\Sigma^-$		44, 262	
	40	$1^1\Sigma^-$		52	
	50	$e^3\Sigma^-$		74	
4	Extra lines:		Q (27, 28) in A-X band [262].		
	0	$^3\Sigma^+$		73, 5, 71, 225	P, R, branches. P branch; very small perturbation (may not be real). McCay refers to a marked decrease in intensity of lines in B-A, 0-4 near Q(20). C-A, 0-4; Q branch only. See also [52]. Intensity drop; see also [87]. B-A, 0-4; reality disputed by [5]. Perturber 70878 $cm^{-1}$ above X. Coster and Brons [44] give 70572 $cm^{-1}$ which is in error.
	18			5	
	20			156	
	23	$1^1\Sigma^-$		73, 225	
	27			5	
5	31			44	A-X, Q branch. $^3\Sigma^+$ and $^3\Pi$ cross A $^1\Pi$ close to one another, largely overlap. See also [229]. A-X, R(35), Q(36), P(37). Perturber may be $d^3\Delta, v=14$ . See Ref. [81]. A-X, Q branch. See ref. [52]. A-X, R(50), P(52) begins $a^3\Sigma^-$ perturbation. See [234b, 229].
	34, 37, 40	$e^3\Sigma^-, v=7$	1.17	44, 73	
	44, 46, 49	$a' ^3\Sigma^+$		73, 229	
	50	$^3\Pi$		73	
	30, 33, 35	$a' ^3\Sigma^+$		73, 71	
6	36	$^3\Pi$		73	A-X. A-X; $^3\Pi, ^3\Sigma$ overlap; destroyed order of lines below $J=16$ . A-X, P, R (7, 8); Q(7, 8). A-X, Q, P, Q; similar perturbation found at $v=0$ by Coster and Brons [44] A-X, Q branch (perturbing state originally identified as $^3\Pi$ ).  $e^3\Sigma^-, \sim 73777 \text{ cm}^{-1}$ .  Q branch; probably the beginning of a $^3\Pi$ perturbation; other branches not followed to such high $J$ values, so identification of perturbation is uncertain. Extra lines A-X    P   Q   R J { 7   7 11 13 13 16
	44	$1^1\Sigma^-$	1.17	73	
	50-1, 53, 56	$e^3\Sigma^-, v=8$	1.13	73	
	0	$\Pi$ or $\Delta$	$>1.46$	182	
	Near head			73	
7	6-8	$d^3\Delta, v=12$		182	A-X, 7-1. A-X, Q branch.  Intensity drop; possible perturbation.
	11, 14, 16	$a' ^3\Sigma^+, v=17$		182	
	14	$^3\Delta_i$		182	
	29	$^3\Sigma^-$		209, 73	
	30	$1^1\Sigma^-$	1.11	73, 52	
8	41, 44, 47	$e^3\Sigma^-$	1.12	229	A-X. A-X. A-X. Intensity drop; possible perturbation.
	45, 47, 50	$a' ^3\Sigma^+$	1.07	73, 229	
	52	$^3\Pi(?)$		73	
				182	
9	26, 29, 32	$e^3\Sigma^-$	1.14	209, 73	Extra line Q (13). Intensity breakoff.
	27			209	
	33, 36, 38	$a' ^3\Sigma^+$	1.06	73, 229	
	39	$^3\Pi$		73	
	48			74, 87	
10	16, 18, 21	$a' ^3\Sigma^+$		73, 229	A-X. A-X. A-X. Intensity drop; possible perturbation.
	27	$^3\Pi$		73	
	34	$1^1\Sigma^-$	0.93	73, 52	
	37			74, 87	
11	13	$1^1\Sigma^-$	0.86	73, 74, 52	A-X. A-X. A-X. Intensity drop; possible perturbation.
	22			74, 87	
	28			74	
	37	$^3\Sigma^+$		74	



TABLE 54. *Perturbations of the electronic states of CO and CO<sup>+</sup>*—Continued

<i>v</i>	Max. pert. ( <i>J</i> )	Perturbing term	Perturber <i>B</i> <sub>0</sub>	Ref.	Comments
<hr/>					
10	16, 18, 21	<i>e</i> <sup>3</sup> Σ <sup>-</sup>	<i>cm</i> <sup>-1</sup>	73	A-X, P(17), Q(18); overlapped by <sup>3</sup> Σ <sup>+</sup> . See [31, 229].
	19, 21, 23	<i>a'</i> <sup>3</sup> Σ <sup>+</sup>	-----	229	
	32	<sup>3</sup> Π	-----	73	A-X.
11	16	<sup>3</sup> Π, <sup>1</sup> Σ <sup>-</sup>	0.62	74, 234b	A-X.
12	22, 24, 27	<i>a'</i> <sup>3</sup> Σ <sup>+</sup>	-----	74, 229	A-X.
13	16	<sup>3</sup> Π(?)	-----	74, 87	A-X, analysis till <i>J</i> =15, probably <sup>3</sup> Π perturbation in all branches at the same <i>J</i> .
<hr/>					
(b) E <sup>1</sup> Π state					
0	31	<sup>1</sup> Σ <sup>+</sup>	-----	257	
<hr/>					
(c) a <sup>3</sup> Π state					
0, 1, 2, 3	-----	-----	-----	16	Unperturbed. Gerö says that the <i>b</i> -component ( <i>d</i> -component) of the Π-doublet of a <sup>3</sup> Π, <i>v</i> =0 for <i>N</i> =13 is perturbed by an unknown state, possible <sup>1</sup> Σ <sup>+</sup> .
	-----	-----	-----	81, 76	
4	-----	<i>a'</i> <sup>3</sup> Σ <sup>+</sup> , <i>v</i> =0	1.31	16	<div style="display: flex; align-items: center;"><div style="margin-right: 10px;">P<sub>1</sub>, R<sub>1</sub> at <i>N</i>=28, 38-9 P<sub>2</sub>, R<sub>2</sub> at <i>N</i>=28 P<sub>3</sub>, R<sub>3</sub> at <i>N</i>=26-8 Q<sub>1</sub> at <i>N</i>=33 Q<sub>2</sub> at <i>N</i>=22-3, 32-3 Q<sub>3</sub> at <i>N</i>=22</div><div style="font-size: 3em; margin-right: 10px;">}</div><div>eight perturbations.</div></div>
5	-----	-----	-----	16	
6	-----	-----	-----	16	
	-----	-----	-----	16	
<hr/>					
(d) d <sup>3</sup> Δ <sub>5</sub> state					
6	8	} a <sup>3</sup> Π, <i>v</i> =12	-----	82, 81	<i>d</i> - <i>a</i> , 6-0.
	13-14		-----		
	14-15		-----		
<hr/>					
(e) f <sup>3</sup> Σ <sup>+</sup> state					
Higher term	0	<sup>3</sup> Σ <sup>+</sup> <i>b</i> <sup>3</sup> Σ <sup>+</sup> , <i>v</i> =0	-----	78	The <i>v</i> numbering is uncertain for the <i>f</i> state
	N=13		-----	78	
	low <i>J</i> N=13	<sup>3</sup> Σ <sup>+</sup> <i>b</i> <sup>3</sup> Σ <sup>+</sup> , <i>v</i> =1	-----	78	
			-----	78	
<hr/>					
(f) b <sup>3</sup> Σ <sup>+</sup> state					
0	N=19	<i>f</i> <sup>3</sup> Σ <sup>+</sup> , <i>v</i> =0	-----	54	<i>b</i> - <i>a</i> .
	N=55		-----	54	Large perturbation in <i>b</i> - <i>a</i> bands; probably Π or Δ.
	20-1		-----	} 72	
	29		-----		N values. Perturber is <i>f</i> or <i>a'</i> .
	36		-----		
	41		-----		
	44-5	-----	} 72, 247		
	48	-----			
	50	-----			
	52	-----			
1	17	<i>f</i> or <i>a'</i>	-----	} 72, 247	N values. [247] states that "there are no unperturbed levels in the b <sup>3</sup> Σ <sup>+</sup> state". Lines deviate from their normal positions by as much as 36 cm <sup>-1</sup> for <i>v</i> =0 and 130 cm <sup>-1</sup> for <i>v</i> =1
	25		-----		
	30		-----		
	34		-----		
	40		-----		

TABLE 54. *Perturbations of the electronic states of CO and CO<sup>+</sup>—Continued*

<i>v</i>	Max. pert. ( <i>J</i> )	Perturbing term	Perturber B.	Ref.	Comments
(g) A <sup>2</sup> Π <sub>i</sub> state of CO <sup>+</sup>					
5	14.5 22.5	X <sup>2</sup> Σ <sup>+</sup> , <i>v</i> =14 X <sup>2</sup> Σ <sup>+</sup> , <i>v</i> =14	----- -----	45 45	<sup>2</sup> Π <sub>1/2</sub> component perturbed by Σ, <i>J</i> = <i>N</i> -½. <sup>2</sup> Π <sub>1/2</sub> (other A-component) perturbed by Σ, <i>J</i> = <i>N</i> +½. Successive extra lines in both Q <sub>1</sub> and Q <sub>2</sub> branches. <sup>2</sup> Π <sub>3/2</sub> level is unperturbed.
10	-----	X <sup>2</sup> Σ <sup>+</sup> , <i>v</i> =18	-----	35	Numerous levels perturbed ( <i>J</i> =3½ to 21½).

TABLE 55. *Observed predissociations in C<sup>12</sup>O<sup>16</sup> and C<sup>13</sup>O<sup>16</sup> <sup>a</sup>*

Molecule	Term	<i>v</i>	<i>J</i> for last strong line	<i>J</i> for first weak line	Rotational energy	Energy above X, <i>v</i> =0, <i>J</i> =0	Ref.
C <sup>12</sup> O <sup>16</sup>	B <sup>1</sup> Σ <sup>+</sup>	0	37	----- 38	2738.2 2886.2	89655.9 89803.9	56, 43.
			17	----- 18	537.8 657.0	89587.6 89656.8	
		1	42	----- 43	3472.9 3638.3	89488 89653	74, 224, 74.
			55	----- 56	6052.2 6273.3	89884.7 90104.8	
	b <sup>3</sup> Σ <sup>+</sup>	0	-----	~ 27	(1479.0)	(93555.1)	77, 221.
	C <sup>1</sup> Σ <sup>+</sup>	0	28	----- 29	1577.1 1689.7	93497.8 93610.4	74, 224b 221.
				-----	-----	-----	
	C <sup>3</sup> Σ <sup>+</sup>	0	-----	~ 27	(1479.0)	(93555.1)	77, 221.
C <sup>13</sup> O <sup>16</sup>	B <sup>1</sup> Σ <sup>+</sup>	0	39	----- 40	2893.1 3040.8	89808.4 89956.1	56 (Using observed ZPE from ref. 157b).
			19	----- 20	697.7 771.0	89650.3 89723.6	
		1	-----	-----	-----	-----	-----

<sup>a</sup> Hagstrum [87] has summarized the predissociations using slightly different data.



TABLE 56. Vibration-rotation bands of CO in the infrared region

Isotope	Origin $\sigma_0$		Band	Ref.
	Obs.	Calc. <sup>b</sup>		
	$cm^{-1}$			
$C^{12}O^{16}$	2116.80	-----	2-1	193
	2143.274	0.2731	1-0	192, 193
	3996.88	.9466	7-5	85
	4049.24	.2958	6-4	85
	4101.73	.7820	5-3	85
	4154.404	.4056	4-2	273
	4207.168	.1664	3-1	273
	4260.064	.0646	2-0	<sup>a</sup> 197
	6350.436	.4404	3-0	197
	8414.458	.4702	4-0	102
$C^{13}O^{16}$	2096.071	-----	1-0	193
$C^{12}O^{18}$	2092.12	-----	1-0	162

<sup>a</sup> Absolute accuracy claimed is 1 part in  $5 \times 10^6$   
Relative accuracy claimed is 1 part in  $10^7$  } measurement

Calculated line frequencies accurate to 1 part in  $10^6$

Similar accuracies claimed for 1-0 band [196].

<sup>b</sup> Calculated origins from Wiggins [273]; some data previously given by Rank et al., [198].

TABLE 57. Dipole moments for states of CO and CO<sup>+</sup>

Molecule	State	$\mu$ (Debye)
CO	X <sup>1</sup> $\Sigma$	<sup>a</sup> 0.114 $\pm$ 0.005
	A <sup>1</sup> $\Pi$	.03 $\pm$ 0.08
	B <sup>1</sup> $\Sigma$	.15 $\pm$ 0.15
	a <sup>3</sup> $\Pi$	<sup>b</sup> 1.38
CO <sup>+</sup>	X <sup>2</sup> $\Sigma$	.1 $\pm$ 0.1
	A <sup>2</sup> $\Pi$	.1 $\pm$ 0.1

<sup>a</sup> This is the value of Burrus [36] to which has been added a polarizability correction of 0.002D [163] (1D= $10^{-18}$  esu).

<sup>b</sup> A preliminary value determined from the radio frequency spectrum of the a<sup>3</sup> $\Pi$  state [67(a)].  
Data from [134].

TABLE 58. Rotational frequencies, rotational constants, and magnetic moments of carbon monoxide

Isotope	Frequency $\nu$	Transition $J' \leftarrow J''$	Wave length	$B_e^c$	$B_e^d$	$\mu_J (=g)^a$	Ref.
	MHz		mm	MHz	MHz	Nuclear magnetons	
$C^{12}O^{16}$	691472.60 $\pm$ 0.60	6 $\leftarrow$ 5	0.43	-----	-----	-----	122
	576267.75 $\pm$ 0.10	5 $\leftarrow$ 4	0.52	-----	-----	-----	
	461040.68 $\pm$ 0.06	4 $\leftarrow$ 3	0.65	-----	-----	-----	
	345795.900 $\pm$ 0.090	3 $\leftarrow$ 2	0.87	-----	-----	-----	86
	230537.974 $\pm$ 0.030	2 $\leftarrow$ 1	1.30	-----	-----	-----	
	115271.204 $\pm$ 0.005	1 $\leftarrow$ 0	2.60	<sup>e</sup> 57898.462	57907.863	-0.26910	
$C^{12}O^{17}$	<sup>b</sup> 112359.276 $\pm$ 0.060	1 $\leftarrow$ 0	-----	56432.560	56441.491	-.26227	215, 47
$C^{13}O^{16}$	110201.370 $\pm$ 0.008	1 $\leftarrow$ 0	-----	55346.453	55355.038	-.25704	
$C^{12}O^{18}$	109782.182 $\pm$ 0.008	1 $\leftarrow$ 0	-----	55135.337	55143.861	-.25622	
$C^{14}O^{16}$	105871.110 $\pm$ 0.004	1 $\leftarrow$ 0	-----	53166.929	53174.841	<sup>b</sup> -.24664	
$C^{13}O^{18}$	104711.416 $\pm$ 0.008	1 $\leftarrow$ 0	-----	52583.285	52591.033	<sup>b</sup> -.24418	

<sup>a</sup> The error of  $\pm 0.0005$  MHz to these values is the absolute error; the relative errors are about three times smaller.

<sup>b</sup> Calculated.

<sup>c</sup> Uncorrected values derived from frequency measurement with values of  $\alpha_e = 525.03$  MHz,  $\gamma_e = 0.0887$  MHz,  $D_e = 0.18347$  MHz [17]. These  $B_e$  values differ slightly from those given in ref. [215] primarily because of the different value of  $\alpha_e$ , which still introduces an uncertainty in the  $B_e$  values of about 0.2 MHz.

<sup>d</sup> Corrected for nonspherical distribution of the electrons, wobble stretching, and the Dunham corrections [215, 47].

<sup>e</sup>  $B_0 = 57635.970 \pm 0.003$  MHz;  $D_0 = 0.18390 \pm 0.00014$  MHz [122, 86]. This value of  $D_0$  is not compatible with the preferred value of 0.18345 MHz derived from IR data [198, 196].

Table 60. (a) Franck-Condon factors for the  $\Lambda^1\Pi - \chi^1\Sigma^+$  Fourth Positive system

$v''v'$	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	
0	1.1319-1	2.0077-1	2.8477-1	1.8299-1	9.6040-2	3.5535-2	1.0340-2	2.4282-3	4.6859-4	7.5264-5	1.0151-5	1.1563-6	1.1817-7	9.1869-9	6.4215-10	3.7783-11	1.8874-12	9.0394-14	1.9632-15	1.7459-16	6.9579-16	8.4461-18	4.7739-16	2.2078-16	2.3198-17	
1	2.1612-1	1.5487-1	1.0487-1	3.0508-3	7.6435-2	1.9313-1	1.8569-1	1.0829-1	4.4465-2	1.3804-2	3.3745-3	6.6637-4	1.0812-4	1.4579-5	1.6462-6	1.5635-7	1.2525-8	8.4721-10	4.8179-11	2.2816-12	8.6633-14	3.5241-15	5.5781-16	1.8542-16	1.2590-16	7.7682-16
2	2.2977-1	1.2179-2	9.0126-2	1.1607-1	5.0830-3	5.7179-2	1.6502-1	1.5661-1	9.9629-2	4.1314-2	1.2831-2	3.1192-3	6.0724-4	9.6710-5	1.2731-5	1.3957-6	1.2800-7	9.8388-9	6.3500-10	3.4137-11	1.5954-12	5.3403-14	8.4261-16	2.4194-16	5.7173-19	
3	1.8128-1	2.0493-2	1.1706-1	6.4536-4	8.9595-2	8.4157-2	4.7128-4	6.7660-2	1.5881-1	1.4862-1	8.4545-2	3.3725-2	1.0112-2	2.3717-3	4.4621-4	6.8445-5	8.6524-6	9.0748-7	7.9258-8	5.7731-9	3.5084-10	1.7696-11	0.0676-13	2.3445-14	2.0941-15	
4	1.8919-1	8.7205-2	3.4433-2	5.7630-2	6.6561-2	6.1711-3	2.8034-2	5.1406-2	2.7202-3	8.9110-2	1.2941-1	1.2741-1	6.7610-2	2.5283-2	7.1696-3	1.5970-3	2.8572-4	4.1658-4	4.9963-6	4.9572-7	4.0797-8	2.7899-9	1.5695-10	7.2888-12	3.0316-13	
5	6.8846-2	1.2301-1	3.2357-4	9.4254-2	1.3841-5	8.2355-2	2.2684-2	3.1553-2	2.4377-2	2.1092-2	1.8021-2	1.2991-1	1.5095-1	1.0784-1	5.1173-2	1.7696-2	3.6707-2	1.1654-2	2.8628-3	5.5919-4	8.7798-5	1.1256-5	1.1833-6	1.0277-7	7.2653-9	
6	3.6675-2	1.1603-1	3.2109-2	4.2201-2	4.2898-2	4.2285-2	2.1776-2	7.2895-2	5.9304-4	6.3519-2	7.2827-2	2.4331-3	4.5471-2	1.3211-1	1.3895-1	8.5849-2	1.2017-1	6.4922-2	2.4980-2	7.2532-3	1.6441-3	2.9713-4	4.3394-5	1.1631-6	5.0652-7	
7	1.8423-2	8.8097-2	7.3465-2	2.2336-3	7.3886-2	4.2668-5	6.9877-2	2.9021-3	5.8181-2	3.7199-2	1.0104-2	8.2806-2	4.0479-2	2.7402-2	7.8643-2	1.4142-1	1.0892-1	1.3911-1	9.8285-2	4.6607-2	1.6441-3	4.2794-3	8.9197-4	1.4877-4	1.0027-5	
8	8.8774-3	5.8371-2	5.0972-2	9.9994-3	4.4622-2	3.2779-2	2.9742-2	3.0612-2	4.2672-2	9.8388-3	5.8218-2	6.0762-3	4.0732-2	7.7972-2	1.4842-2	2.2424-2	1.8307-4	5.7010-2	6.0094-2	2.2404-3	3.7782-2	1.6611-1	1.2977-1	8.5100-2	3.8334-2	
9	4.1842-3	3.5370-2	8.4527-2	4.0627-2	7.9687-3	6.0808-2	3.8286-5	7.8344-2	1.6689-5	5.8819-2	5.5671-3	2.4513-2	2.4546-3	6.2762-2	5.2619-2	2.6032-2	5.6287-2	1.8132-2	1.6792-2	1.0029-2	8.9341-2	1.3590-1	1.0721-1	5.5381-2	2.0615-2	5.8141-3
10	1.9367-3	2.0162-2	6.6321-2	6.4755-2	1.4618-3	4.9952-2	2.4181-2	2.7371-2	3.2532-2	2.3741-2	2.4797-2	4.1697-2	5.1872-2	6.2178-2	1.2255-2	2.6032-2	1.8390-5	5.4755-2	1.2915-1	1.2656-1	7.5863-2	3.1782-2	9.9277-3	2.3977-3	4.9850-4	
11	8.8947-4	1.4024-2	4.6641-2	4.6641-2	1.4024-2	1.4504-2	1.4504-2	1.1030-5	2.9442-2	1.0943-3	5.7511-2	4.1697-2	5.1872-2	6.2178-2	1.2255-2	2.6032-2	1.3307-4	5.7010-2	6.0094-2	2.2404-3	3.7782-2	1.6611-1	1.2977-1	8.5100-2	3.8334-2	
12	4.0787-4	5.8643-3	3.0446-2	6.5986-2	4.1768-2	1.7143-4	4.4180-2	1.6298-2	2.1592-2	3.0204-2	1.4383-2	3.1435-2	2.0623-2	2.0759-2	4.0979-2	1.9786-3	1.1148-2	2.5596-2	2.0612-2	4.4282-4	5.0304-2	2.4094-2	6.4819-2	2.4538-2	1.0409-1	
13	1.8777-4	3.0666-3	1.8881-2	5.2041-2	6.5313-2	6.7201-3	2.0995-2	3.1909-2	4.2287-4	1.2877-2	1.7179-3	4.1004-2	2.1272-3	4.4591-2	6.1213-6	4.6549-2	1.6848-2	4.0913-2	4.4129-2	5.0304-2	2.4094-2	6.4819-2	2.4538-2	1.0409-1		
14	8.7084-5	1.5931-3	1.1304-2	3.8223-2	5.7629-2	2.3102-2	3.3642-3	4.1680-2	9.2001-3	2.1882-2	2.5544-2	1.0672-2	3.1590-2	1.0167-2	2.9406-2	1.9778-2	4.1117-2	1.5983-4	4.1675-2	1.4535-2	1.8417-2	5.1503-2	2.7593-3	3.9879-2	3.6755-2	
15	4.0646-5	8.2177-4	6.6118-3	2.6471-2	5.1936-2	3.1973-2	7.3801-4	2.6615-2	2.8617-2	1.9365-3	3.1613-2	1.2081-3	3.4831-2	4.0014-3	3.5446-2	2.3548-3	8.6642-3	2.6530-2	2.0329-2	1.5822-2	1.1357-2	2.4062-5	1.2985-2	3.0498-2	3.4517-3	
16	1.9425-5	4.2525-4	3.8121-3	1.1597-2	4.2528-2	4.5927-2	9.4788-3	9.4220-3	3.6896-2	3.5232-3	2.3565-2	1.9857-2	1.0192-2	2.8332-2	6.0848-3	1.0450-2	4.8212-3	3.2361-2	1.7042-3	3.6887-2	8.6115-4	3.8212-2	1.5083-2	1.1515-2	5.1596-2	
17	9.3867-6	2.2157-4	2.1816-3	1.1373-2	3.2596-2	4.6731-2	2.1795-2	6.1450-4	3.0143-2	1.8049-2	5.1274-3	3.2762-2	2.6912-4	3.1077-2	3.7993-3	2.9198-2	2.7723-2	4.6775-3	2.7916-2	9.2564-3	2.2094-2	2.2721-2	7.8327-3	2.1710-2	1.2774-3	
18	4.6176-6	1.1653-4	1.2459-3	7.2199-3	2.3851-2	4.2574-2	3.1818-2	1.6937-3	1.6811-2	2.9379-2	3.2516-4	2.5475-2	1.2637-2	1.1864-2	2.3360-2	5.3000-3										

Data taken from Nicholls [178].

TABLE 59. Hyperfine structure

(a) Hyperfine structure caused by  $O^{17}$  nuclear quadrupole moment

$J=1-0$	Transition	$\nu$
	$\Delta F$	MHz
	0	112360.016 $\pm$ 0.015
	+1	112383.930 $\pm$ 0.015
	-1	112358.720 $\pm$ 0.100

(b)  $\nu_0$  MHz <sup>a</sup>112359.276 $\pm$ 0.060(c) eq Q = b = +4.43 $\pm$ 0.040 MHz <sup>b</sup>(d) Q = -0.026 $\pm$ 0.009 barn

Data taken from Jarmain, Ebisuzaki, and Nicholls [113].

<sup>a</sup> Undisplaced center of triplet calculated from known value of nuclear masses with corrections [214].<sup>b</sup> For  $O^{17}O^{17}$  [214].0.1 barn = 10<sup>-28</sup> cm<sup>2</sup>; above value of nuclear quadrupole moment is derived from microwave measurements [249]; value from paramagnetic studies is -0.0265 $\pm$ 0.003 barn [123].



TABLE 61. *Franck-Condon factors, band strengths, r-centroids, and  $I_{\infty}$  for the  $B^1\Sigma^+ - A^1\Pi$  Ångström system*

$v' \backslash v''$	0	1	2	3	4	5	6	7	8	9	10	11	12	13
0	0.08898	0.18159	0.21056	0.18339	0.13399	0.08706	0.06211	0.20941	0.01591	0.00835	0.00429	0.00217	0.00109	0.00055
	0.088	.171	.154	.139	.110	.091								
	1.177	1.158	1.139	1.122	1.106	1.091								
	6.8	10.0	6.7	4.5	2.6	1.5								
1	0.25053	.17569	.03039	.00420	.05214	.09553	.10665	.09311	.07008	.04781	.03046	.01849	.01085	
	0.402	.202	.028	.002	.034	.070	.093							
	1.205	1.185	1.166	1.148	1.131	1.115	1.100							
	4.4	1.7	0	0	0	0.2	0.2							

 $q$  Ref. [183] (also includes  $q$  for  $v' \geq 2$ ). $r$ -centroid ref. [211]. $p$  Ref. [211];  $p$  and  $q$  are made equal for 0-0;  $I_{\infty}$  for the strongest band in  $v'=0$  progression is normalized to 10.0. $I_{\infty}$  ref. [211].TABLE 62. *Franck-Condon factors for the  $C^1\Sigma^+ - A^1\Pi$  Herzberg system*

$v' \backslash v''$	0	1	2
0	0.072	0.165	0.214
1	.211	.192	.031
2	.293	.034	-----

Data from Leskov (ref. [150]).

TABLE 63. *Franck-Condon factors, band strength, r-centroids, and  $I_{\infty}$  for the  $b^3\Sigma^+ - a^3\Pi$  Third Positive system*

$v' \backslash v''$	0	1	2	3	4	5	6	7	8	9	10	11	12	13
0	0.05810	0.13468	0.17897	0.17905	0.14996	0.11112	0.07531	0.04771	0.02869	0.01656	0.00925	0.00503	0.00268	0.00140
	0.054	0.082	0.066	0.035	0.013	0.002								
	1.149	1.131	1.115	1.100	1.086	1.072								
	8.0	10.0	6.6	2.8	0.8	0.1								
1	0.21702	.19811	.06809	.00213	.01660	.06808	.08657	.09191	.08060	.06238	.04421	.02935	.01853	
	0.342	0.225	0.054	0.001	0.003	0.006	0.003							
	1.173	1.154	1.136	1.120	1.105	1.091	1.077							
	64.0	35.0	6.9	0.1	0.3	0.4	0.2							
2	0.32971	0.02800	.03677	.10572	.07610	.01909	.00015	.01803	.04558	.06352	.06708	.06000		

 $q$  Ref. [183]. $I_{\infty}$  Ref. [211].  $I_{\infty}$  for the strongest band in the  $v'=0$  progression is normalized to 10.0. $p$  Ref. [211].  $p$  and  $q$  are made  $\approx$  equal for the 0-0 band. $r$ -centroid ref. [211].TABLE 64. *Franck-Condon factors for the  $a'^3\Sigma^+ - a^3\Pi$  Asundi system*

$v'$	$v''$			$v'$	$v''$		
	0	1	2		0	1	2
0.-----	0.03 <sub>7</sub>	0.14 <sub>3</sub>	0.24 <sub>3</sub>	7.-----	0.05 <sub>9</sub>	0.10 <sub>2</sub>	0.00 <sub>1</sub>
1.-----	.10 <sub>3</sub>	.19 <sub>5</sub>	.09 <sub>7</sub>	8.-----	.03 <sub>6</sub>	.09 <sub>0</sub>	.00 <sub>6</sub>
2.-----	.16 <sub>2</sub>	.10 <sub>9</sub>	.00 <sub>0</sub>	9.-----	.02 <sub>1</sub>	.08 <sub>2</sub>	.02 <sub>3</sub>
3.-----	.18 <sub>1</sub>	.01 <sub>9</sub>	.05 <sub>0</sub>	10.-----	.01 <sub>0</sub>	.06 <sub>4</sub>	.04 <sub>1</sub>
4.-----	.16 <sub>5</sub>	.00 <sub>3</sub>	.05 <sub>9</sub>	11.-----	.01 <sub>2</sub>	.05 <sub>0</sub>	.05 <sub>2</sub>
5.-----	.12 <sub>9</sub>	.03 <sub>9</sub>	.06 <sub>7</sub>	12.-----	.00 <sub>7</sub>	.03 <sub>8</sub>	.05 <sub>5</sub>
6.-----	.09 <sub>1</sub>	.08 <sub>1</sub>	.02 <sub>4</sub>				

Data from Jarman, Fraser, and Nicholls [114].

Table 65. Franck-Condon factors for the  $a^3\Pi - X^1\Sigma^+$  Cameron system

$v''$	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	
0	2.6500-1	3.7800-1	2.3951-1	8.9476-2	2.2038-2	3.7861-3	4.2355-3	2.8555-6	1.3893-7	5.0012-9	1.2989-10	2.3326-12	3.3047-14	9.1193-16	2.9180-16	3.9476-17	2.6795-16	4.9718-16	1.6200-17	2.0220-16	1.6200-17	2.0220-16	1.6200-17	2.0220-16	1.6200-17	2.0220-16
1	3.1901-1	2.8712-2	1.0300-1	2.6517-1	9.1600-1	7.2368-2	1.7063-2	2.7157-3	3.0355-4	2.4313-5	1.4060-6	5.8594-8	1.7433-9	3.6439-11	4.9358-13	3.2162-15	1.3872-16	5.8010-17	7.9066-18	1.1789-16	2.5255-17	2.3684-16	3.1879-16	3.4180-17	7.1301-16	
2	2.889-1	5.1400-2	1.5924-1	7.5855-2	1.4638-1	2.3381-1	1.3634-1	4.3693-2	8.8478-3	1.2007-3	1.1517-4	7.7721-6	3.7269-7	1.2612-8	2.9550-10	4.6008-12	5.4893-14	1.4047-15	1.7788-16	1.1072-16	2.5792-16	2.9580-18	2.6741-16	2.6022-17	1.0951-16	
3	1.1396-1	1.5951-1	9.7775-2	1.2910-1	5.8927-2	1.2955-1	1.8913-1	8.2714-2	2.1208-2	3.5178-3	3.8616-4	1.7019-5	1.7019-5	1.7019-5	1.7200-9	3.0529-11	3.3733-13	1.4595-15	6.2277-21	1.7772-16	4.3996-17	1.6740-17	4.6659-17	2.2711-17		
4	5.0350-2	1.6609-1	3.5793-2	3.6110-2	1.2121-1	4.2124-4	1.2000-1	2.0982-1	4.2758-1	4.1463-2	8.3235-3	1.1032-3	9.9758-5	6.2393-6	2.6989-7	7.9777-9	1.5732-10	2.0134-12	1.8275-14	1.5746-17	1.6023-18	1.9767-17	1.0614-16	3.5360-16		
5	2.0031-2	1.1062-1	1.1183-1	1.1718-3	1.1270-1	8.5459-5	1.1653-1	3.9092-2	4.3675-2	1.9101-1	1.6783-1	1.6915-2	2.6322-3	2.7411-4	1.9480-5	9.4838-7	3.1326-8	6.8959-10	9.7081-12	7.8353-14	7.4792-17	8.5007-16	1.2108-16			
6	7.036-3	6.0231-2	1.2681-1	3.5975-2	4.1084-2	7.2436-2	3.0659-2	6.2812-2	3.1855-2	1.4106-1	1.9248-1	1.0345-1	3.0490-2	5.5668-3	6.6637-4	5.3683-7	2.939-6	1.0813-7	2.6360-9	4.0840-10	3.3264-13	4.3673-15	1.0726-16	1.5510-16		
7	2.6402-3	2.8609-2	9.4632-2	9.0858-2	2.8742-4	8.1935-2	1.8761-2	7.5386-2	1.3024-2	1.1088-1	7.7275-3	7.9310-2	1.9388-1	1.3828-1	4.9750-2	1.0668-2	1.4657-3	1.3379-4	8.1980-6	3.3621-7	9.0660-9	1.5554-10	1.5277-12	1.0778-14	7.2723-16	

Data from Nicholls [178].

Table 66. Franck-Condon factors for the  $b^3\Sigma^+ - X^1\Sigma^+$  Hopfield-Birge system

$v''$	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	
0	6.4628-1	2.5972-1	6.8320-2	1.7672-2	5.0662-3	1.3807-3	5.1937-4	1.7758-4	6.3613-5	9.1523-5	3.6588-6	1.5093-6	6.4423-7	2.8017-7	1.2580-7	5.7736-8	2.7171-8	1.3078-8	6.4324-9	3.2290-9	1.6529-9	8.6180-10	4.5726-10	2.4669-10		
1	2.4628-1	1.6322-1	2.7984-1	1.5516-1	6.0316-2	2.2891-2	8.9289-3	3.5504-3	1.4335-3	5.9018-4	2.4870-4	1.0736-4	4.7436-5	2.1437-5	9.9013-6	4.6714-6	2.2498-6	1.1053-6	5.5355-7	2.8241-7	1.4668-7	7.7518-8	4.1660-8	2.2754-8	1.2623-8	

Data from Nicholls [178].

Table 67. Franck-Condon factors for the  $a^3\Sigma^+ - X^1\Sigma^+$  Hopfield Birge system

$v''$	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24
0	1.5697-4	1.6277-3	8.4812-3	2.6351-2	6.4009-2	1.0798-1	1.5170-1	2.7343-1	1.6411-1	1.3008-1	8.7089-2	4.9521-2	2.3997-2	9.9259-3	3.5040-3	1.0540-3	2.6929-4	5.8162-5	1.0549-5	1.5927-6	1.9798-7	1.9963-8	1.5991-9	9.9531-11	4.7857-12
1	1.0767-3	8.8725-3	3.3690-2	7.6925-2	1.1168-1	1.1118-1	6.1653-2	9.7384-3	5.9000-3	5.4248-2	1.4221-1	1.3682-1	1.1960-1	8.0765-2	4.3671-2	1.9270-2	7.0007-3	2.1102-3	5.2609-4	1.0828-4	1.8202-5	2.5083-6	2.7590-7	2.3830-8	1.5862-9
2	3.9034-3	2.4998-2	6.9140-2	1.0331-1	8.1031-2	2.4228-2	2.3255-2	4.6942-2	8.2793-2	1.4975-2	6.9467-3	1.0517-2	6.6862-2	1.4860-2	1.2495-1	3.3309-2	5.3002-2	2.3701-2	8.4947-3	2.4580-3	5.7598-4	1.0895-4	1.6515-5	1.9818-6	1.8495-7
3	9.9030-3	4.8226-2	9.2253-2	7.7504-2	1.6123-2	6.9151-3	5.5939-2	5.9960-2	1.1284-2	9.0294-3	5.9203-2	6.8427-2	2.1514-2	1.6357-3	4.7716-2	1.0768-1	1.2388-1	9.5266-2	5.3726-2	2.3231-2	7.8706-3	2.1097-3	4.3823-4	7.5157-5	9.8435-6
4	1.9710-2	7.4369-2	8.6776-2	2.8151-2	2.8185-3	4.9502-2	4.9084-2	3.0263-2	2.2514-2	5.9330-2	2.6219-2	1.1286-3	4.4512-2	6.0395-2	2.6367-2	6.9983-4	4.5300-2	1.0718-1	1.2226-1	9.0538-2	4.8245-2	1.9331-2	5.9642-3	1.4283-3	2.6541-4
5	3.3075-2	8.5232-2	5.7084-2	2.2752-2	6.6924-3	5.0536-2	4.5541-3	2.4252-2	4.6830-2	1.1566-2	1.4839-2	1.3552-2	9.1162-3	5.0762-2	2.6398-2	1.3949-3	4.6857-2	2.1733-2	5.4351-2	1.1833-1	1.1864-1	8.0196-2	3.8632-2	1.3832-2	3.7504-3
6	4.8277-2	8.4912-2	2.2752-2	2.4015-3	3.1103-3	3.6692-2	3.8977-2	2.0914-2	5.3343-3	4.4913-2	1.1796-2	4.4448-2	9.7984-3	1.3195-2	5.0551-2	1.7414-2	6.4357-3	5.6548-2	3.397-2	8.6147-3	2.2666-2	1.2255-1	1.0898-1	6.5020-2	2.7306-2
7	6.3177-2	7.0732-2	2.4015-3	2.3681-3	4.7645-2	9.9733-3	1.8500-2	3.5116-2	1.4081-1	3.3619-2	1.8494-2	3.9222-2	7.4098-3	1.6262-2	4.1100-2	3.5445-3	2.2850-2	4.7619-2	6.2729-3	1.9316-2	6.4157-2	3.2199-2	6.1712-2	1.2571-1	9.4086-2
8	7.6099-2	4.8694-2	2.3681-3	1.6517-2	4.0495-2	9.2551-3	3.5952-2	9.7616-3	1.6205-2	3.3398-2	1.2204-2	1.0978-2	3.6331-2	1.8255-3	2.5710-2	3.3234-2	1.1195-2	3.6537-2	3.6733-2	8.4435-2	4.0196-2	6.8523-2	5.3937-2	1.5505-2	
9	8.3998-2	2.6149-2	1.6517-2	2.7446-2	1.1011-2	3.2677-2	3.3262-4	3.3303-2	5.4679-3	1.9683-2	2.2929-2	2.6866-3	4.4098-2	2.0484-2	2.8555-2	3.8342-4	3.6135-2	1.8934-2	6.0007-3	4.7110-2	1.7748-2	8.1978-3	6.0122-2	4.0075-2	
10	8.7426-2	9.1030-3	3.3341-2	2.7446-2	1.1011-2	3.2677-2	3.3262-4	3.3303-2	5.4679-3	1.9683-2	2.2929-2	2.6866-3	4.4098-2	2.0484-2	2.8555-2										

Data from Nicholls [178].



TABLE 68. *Franck-Condon factors, r-centroids and  $I_{\infty}$  for the  $d^3\Delta_1 - a^3\Pi$  Triplet system*

$v' \backslash v''$	0	1	2	3
1	0.007 1.298 1	0.043 1.321 3.4	0.120 1.346 5.1	0.201 1.371 4.0
2	0.031 1.282 5.8	0.114 1.305 14.9	0.162 1.328 10.2	0.091 1.351 3.0
3	0.068 1.267 17.0	0.145 1.289 22.8	0.074 1.311 6.9	0.000 ----- -----
4	0.107 1.252 35.1	0.113 1.274 23.7	0.004 1.286 0.5	0.057 1.323 4.4
5	0.133 1.238 56.3	0.051 1.259 14.3	0.018 1.287 3.3	0.079 1.305 8.7
6	0.141 1.225 75.3	0.008 1.243 3.4	0.061 1.269 14.9	0.030 1.289 4.7
7	0.132 1.212 87.7	0.002 1.241 0.8	0.073 1.254 22.7	0.000 ----- -----
8	0.113 1.200 90.2	0.022 1.220 130	0.045 1.252 18.4	0.020 1.267 5.2
9	0.086 1.186 83.2	0.055 1.211 38.4	0.017 1.238 8.4	0.043 1.240 14.6

$q$  Ref. [240, 239].

$r$ -centroids Ref. [240].

$I_{\infty}$  ref. [239].

The quantum number  $v'$  listed above are one unit higher than that used in refs. [239, 240]. The tabulated quantities are therefore somewhat uncertain.

Table 69. (a) Franck-Condon factors, r-centroids, and wavelengths for the  $A^2\Pi_1 - X^2\Sigma^+$  Comet Tail system of  $CO^+$

$v'$	0	1	2	3	4	5	6	7	8	9	10	11	12	13
				$\nu''$										
0	4.2745-2 1.177 4910.9	1.5201-1 1.201 5499.9	2.4987-1 1.227 6238.7	2.5157-1 1.253 6222.7	1.7335-1 1.280 7170.1	8.6541-2 1.310 8436.1	3.2330-2 1.340	9.2074-3 1.373	2.0181-3 1.407	3.4136-4 1.445	4.4433-5 1.486	4.4130-6 1.531	3.2957-7 1.581	1.8103-8 1.639
1	1.1506-1 1.160 4565.8	1.9325-1 1.184 5072.1	8.0462-2 1.208 5693.6	5.6766-4 1.234 5900.4	9.1665-2 1.260	1.9053-1 1.288	1.7505-1 1.317	9.9924-2 1.347	3.8971-2 1.380	1.0969-2 1.415	2.2812-3 1.453	3.5380-4 1.494	4.0879-5 1.539	3.4847-6 1.590
2	1.6979-1 1.144 4272.0	9.7458-2 1.167 4711.2	3.2276-3 1.191 5000.4	1.0830-1 1.216 5900.4	8.4514-2 1.241	3.5046-4 1.268	7.2252-2 1.295	1.7243-1 1.324	1.6130-1 1.355	8.8296-2 1.388	3.2140-2 1.423	8.2237-3 1.460	1.5168-3 1.502	2.0329-4 1.547
3	1.8244-1 1.129 4017.7	1.2721-2 1.152 4403.3	7.4829-2 1.175 4865.8	6.8836-2 1.198 5420	4.1855-3 1.223	9.7419-2 1.248	6.2005-2 1.275	1.0539-3 1.303	9.1449-2 1.332	1.7241-1 1.362	1.3941-1 1.395	6.6706-2 1.430	2.1102-2 1.468	4.6305-3 1.510
4	1.6004-1 1.114 3795.8	5.7963-3 1.136 4138.9	9.6177-2 1.159 4518.8	5.7808-4 1.182	7.7853-2 1.206	3.5317-2 1.230	1.9930-2 1.256	9.8372-2 1.282	2.9637-2 1.310	1.6384-2 1.339	1.2684-1 1.370	1.6791-1 1.403	1.0874-1 1.438	4.2944-2 1.476
5	1.2187-1 1.100 3600.8	4.5371-2 1.122 3908.0	4.7189-2 1.144 4244.1	3.3357-2 1.166	5.5935-2 1.190	9.2749-3 1.213	7.8567-2 1.238	5.5694-3 1.263	5.3829-2 1.290	8.1631-2 1.318	2.9760-3 1.347	5.7725-2 1.378	1.5869-1 1.411	1.4625-1 1.446
6	8.3692-2 1.087 3427.9	8.3590-2 1.108 3705.3	4.9733-3 1.130	7.1968-2 1.152	2.7299-3 1.174	6.3553-2 1.197	1.3161-2 1.221	4.3797-2 1.245	5.0731-2 1.271	4.9493-3 1.297	8.5397-2 1.325	4.0134-2 1.354	8.7396-3 1.385	1.1564-1 1.419
7	5.3165-2 1.074 3273.9	9.8536-2 1.095 3525.6	5.1549-3 1.116	5.6486-2 1.137	1.6807-2 1.159	4.5990-2 1.182	1.3848-2 1.205	5.5377-2 1.228	2.2886-3 1.253	6.9170-2 1.278	5.5381-3 1.305	4.3949-2 1.333	7.9079-2 1.362	3.0458-3 1.393
8	3.1817-2 1.061 3135.5	9.1846-2 1.082 3366.1	3.2115-2 1.103	1.8802-2 1.124	5.2375-2 1.145	4.0655-3 1.167	5.4195-2 1.189	3.7020-3 1.212	5.2171-2 1.236	1.3426-2 1.261	3.8243-2 1.286	4.6432-2 1.313	5.7134-3 1.341	8.3059-2 1.370
9	1.8182-2 1.049 3222.4	5.9270-2 1.090	2.3453-4 1.110	2.3453-4 1.110	5.4271-2 1.131	9.0051-3 1.153	3.8654-2 1.175	1.6939-2 1.197	3.6070-2 1.220	1.4865-2 1.243	4.6950-2 1.268	4.2012-3 1.294	6.4832-2 1.321	4.4651-3 1.349
10	1.0025-2 1.037 3093.3	5.3529-2 1.057 3093.3	7.2508-2 1.078 3314.2	8.8302-3 1.098	2.8246-2 1.118	3.8194-2 1.139	4.8691-3 1.161	4.7003-2 1.182	4.9856-4 1.205	4.9408-2 1.228	5.7364-4 1.252	5.1822-2 1.276	5.7995-3 1.302	4.6167-2 1.329
11														

LEGEND:  $q_{v',v''}; r_{v',v''}(\text{\AA}); \lambda_{v',v''}(\text{\AA}; R_2 \text{ head})$ .

Data from Nicholls [177]



TABLE 69. Continued—(b) *Smoothed band strengths and  $I_{\infty}$  for the  $A^2\Pi_i-X^2\Sigma^+$  Comet Tail system of  $CO^+$*

$v' \backslash v''$	0	1	2	3
0	0.07 <sub>1</sub> 1.6	0.21 <sub>9</sub> 3.1	0.25 <sub>0</sub> 2.2	
1	0.18 <sub>8</sub> 5.6	0.31 <sub>2</sub> 6.2	0.10 <sub>8</sub> 1.3	
2	0.25 <sub>5</sub> 10.0	0.16 <sub>2</sub> 4.3	0.00 <sub>5</sub> —	0.13 <sub>1</sub> 1.2
3	0.23 <sub>8</sub> 11.9	0.02 <sub>1</sub> .7	.12 <sub>1</sub> 2.9	0.10 <sub>3</sub> 1.6
4	0.16 <sub>9</sub> 10.6	.00 <sub>9</sub> .4	0.16 <sub>8</sub> 5.3	
5	0.09 <sub>5</sub> 7.4	.05 <sub>4</sub> 3.0	0.07 <sub>3</sub> 2.9	
Order of entry: $p_{v',v'',b}$ $I_{\infty}^b = p_{v',v'',b} / \lambda^4_{v',v''}$				

<sup>a</sup>  $p_{v',v''}$  and  $q_{v',v''}$  are made equal for the (0,2) band.

<sup>b</sup>  $I_{\infty}$  for the strongest band in  $v'=2$  progression is normalized to 10.0.

Data from Robinson and Nicholls [212]. See also Nicholls [177] who lists absolute band strengths.

Table 70. Franck-Condon factors, r-centroids, and wavelengths for the  $B^2\Sigma^+ - X^2\Sigma^+$  First Negative system of  $CO^+$

$v'$	0	1	2	3	4	5	6	7	8	9	10	11	12	13
0	5.3146-1 1.144 2189.8	3.3935-1 1.194 2299.6	1.0391-1 1.244 2419.4	2.1425-2 1.287 2550.3	3.3829-3 1.326 2693.9	4.3283-4 1.364	4.6493-5 1.400	4.2800-6 1.436	3.4197-7 1.471	2.3951-8 1.504	1.4753-9 1.540	8.0798-11 1.573	3.8512-12 1.611	1.5198-13 1.662
1	3.1251-1 1.106 2112.4	5.6508-2 1.155 2214.5	3.1801-1 1.204 2325-2	2.1702-1 1.251 2445.8	7.4905-2 1.294 2577.7	1.7485-2 1.333 2722-3	3.0871-3 1.370 2888.2	4.3778-4 1.406	5.1617-5 1.441	5.1685-6 1.475	4.4619-7 1.508	3.3470-8 1.543	2.1985-9 1.577	1.2678-10 1.611
2	1.1150-1 1.069 2042.3	2.5777-1 1.121 2137.8	1.7629-2 1.171 2240.4	1.4638-1 1.213 2352.5	2.5497-1 1.261 2474.2	1.4707-1 1.303 2607.2	5.0060-2 1.340 2752.9	1.2034-2 1.377 2913.2	2.2278-3 1.412	3.3332-4 1.447	4.1505-5 1.480	4.3877-6 1.515	3.9856-7 1.547	3.1397-8 1.581
3	3.2473-2 1.033	1.9840-1 1.085 2067.9	9.1270-2 1.134 2164.3	1.2209-1 1.181 2268.6	1.7280-2 1.224 2381.5	1.9272-1 1.268 2504.6	2.0093-1 1.308 2638.8	1.0112-1 1.348 2785.8	3.3106-2 1.384 2947.6	7.9746-3 1.419	1.5079-3 1.853	2.3253-4 1.486	2.9988-5 1.519	3.2803-6 1.553
4	8.7613-3 1.001	9.3336-2 1.051 2004.7	1.8641-1 1.100 2095.3	2.0368-3 1.151	1.6458-1 1.192 2298.2	1.3055-2 1.235 2412.4	8.0761-2 1.276 2530.8	1.9766-1 1.320 2672.4	1.5481-1 1.356 2850.8	7.0092-2 1.393 2981.2	2.2074-2 1.426	5.2794-3 1.495	1.0075-3 1.494	1.5830-4 1.525
5	2.3512-2 0.972	3.5568-2 1.018	1.4450-1 1.067	9.6430-2 1.116 2123.8	3.3589-2 1.161 2240.7	9.9646-2 1.203	8.4073-2 1.253	5.9474-3 1.289	1.3060-1 1.328 2707.9	1.8110-1 1.365 2858.1	1.1773-1 1.400 3023	4.9196-2 1.433	1.4979-2 1.467	3.5478-3 1.499
6	6.5481-4 0.949	1.2423-2 0.989	7.7652-2 1.036	1.4383-1 1.085 2061.0	1.4131-2 1.132 2154.1	1.0278-1 1.175 2254.3	1.6521-2 1.216 2362.5	1.2332-1 1.263	1.6654-2 1.302	4.2878-2 1.337 2745.1	1.5470-1 1.374 2897.2	1.5653-1 1.408 3064.0	8.9430-2 1.441	3.5151-2 1.474
7	1.9382-4 0.930	4.2548-3 0.964	3.5179-2 1.007	1.1345-1 1.055 2003.1	8.6423-2 1.105 2091.0	7.6920-3 1.141	1.1023-1 1.188 2185.1	7.1716-3 1.233	8.1655-2 1.273	7.6130-2 1.313	1.7244-4 1.347	8.3111-2 1.383 2938.5	1.5900-1 1.417 3107.5	1.3163-1 1.450
8	6.1606-5 0.916	1.4863-3 0.944	1.4785-2 0.982	6.7715-2 1.027	1.1987-1 1.075 2034.3	1.9286-2 1.112	6.0578-2 1.160 2220.3	4.8978-2 1.203	6.0654-2 1.245	1.4760-2 1.286	1.0067-1 1.320	2.8675-2 1.359	1.5114-2 1.394 2984-2	1.1196-1 1.427 3152-7
9	2.1063-5 0.906	5.4116-4 0.928	6.0866-3 0.960	3.5219-2 1.002	7.0725-2 1.086	7.0725-2 1.086	2.1280-3 1.134 2067.8	9.2383-2 1.177 2255.7	1.1706-3 1.217	8.7535-2 1.264	5.9808-3 1.301	5.6825-2 1.337	7.9908-2 1.371	4.2805-3 1.405 2874.5
10	7.7224-6 0.896	2.0752-4 0.916	2.5347-3 0.941	1.7188-2 0.978	6.1765-2 1.023	9.0888-2 1.061	1.6576-2 1.120	4.1251-2 1.157	5.7720-2 1.193 2293.7	2.3196-2 1.241	4.6205-2 1.275	5.2402-2 1.318	4.7300-3 1.350	8.0689-2 1.384

LEGEND:  $q_{v'v''}; r_{v'v''}(\bar{R}); \lambda_{v'v''}(\bar{R}; R_2 \text{ head})$ .

Data from Nicholls [177].



Table 71. Franck-Condon factors, r-centroids, and wavelengths for the  $B^2\Sigma^+ - A^2\Pi_1$

Baldet-Johnson system of  $CO^+$

$v'$	0	1	2	3	4	5	6	7	8	9	10
0	4.1496-1 • 1.208 3953.6	3.1478-1 • 1.173 4209.1	1.6116-1 • 1.143	6.8711-2 1.117	2.6328-2 1.092	9.4074-3 1.068	3.1995-3 1.045	1.0485-3 1.020	3.3355-4 1.026	1.0344-4 1.032	3.1333-5 1.036
1	4.3277-1 • 1.258 3707.4	1.1762-2 1.212	1.5698-1 1.177 4179.1	1.7617-1 • 1.147	1.1671-1 1.121	6.0549-2 1.096	2.7238-2 1.073	1.1160-2 1.050	4.2796-3 1.026	1.5611-3 1.032	5.4720-4 1.036
2	1.3906-1 • 1.345 3496.7	3.8077-1 • 1.261	2.6187-2 1.215	4.0375-2 1.180	1.2166-1 • 1.150	1.2300-1 • 1.124	8.3855-2 1.100	4.6281-2 1.077	2.2368-2 1.055	9.8660-3 1.032	4.0686-3 1.006
3	1.3129-2 1.414 3314.8	2.6016-1 1.348 •	2.6178-1 1.263	8.0269-2 1.217	2.0501-3 1.182	6.4335-2 1.153	1.0282-1 • 1.127	9.0829-2 1.103	6.0397-2 1.081	3.3872-2 1.059	1.6944-2 1.036
4	7.0428-5 1.435	3.2447-2 1.435	3.4332-1 1.348 •	1.7224-1 1.264	1.0488-1 1.219	4.0252-3 1.184	2.6454-2 1.155	7.4452-2 1.129	8.5003-2 1.106 •	6.719-2 1.084	4.3142-2 1.062
5	2.6197-5 1.459	4.3439-5 1.459	5.0155-2 1.459	3.9982-1 1.346	1.2009-1 1.264	1.0358-1 1.219	1.7832-2 1.185	7.5257-3 1.157	4.8716-2 1.131	7.2391-2 • 1.108	6.7323-2 • 1.087
6	1.7089-7 1.6746-7	1.4408-4 1.7060-8	2.7978-5 4.0853-4	6.0829-2 1.488	4.4210-1 1.340	9.5917-2 1.263	8.8811-2 1.219	3.0953-2 1.186	8.1668-4 1.158	2.9484-2 1.133	5.7954-2 1.110
7	1.6746-7	1.7060-8	4.0853-4	7.9766-4	6.1242-2 1.525	4.7689-1 1.333 •	9.1824-2 1.261	6.8927-2 1.219	4.0185-2 1.186	1.6487-4 1.158	1.6742-2 1.134
8	2.1895-9	1.0854-6	6.0635-6	7.5515-4	3.8175-3 1.579	5.0416-2 1.579	5.0427-1 1.324 •	1.0460-1 1.258	4.8088-2 1.217	4.6275-2 • 1.185	1.5661-3 1.158
9	6.5173-10	7.1522-8	2.8422-6	5.4601-5	9.2520-4	1.0692-2	3.0690-2	5.1784-1 1.314 •	1.3454-1 1.254 •	2.8395-2 1.215 •	5.1256-2 • 1.184
10	1.9671-10	8.7212-10	6.2907-7	2.8034-6	2.2658-4	6.1653-4	2.1837-2	9.4667-3	5.0573-1 1.304 •	1.8267-1 1.249 •	1.1825-2 1.212

LEGEND:  $q_{v'v''}; r_{v'v''}(\lambda); \lambda_{v'v''}(\lambda; R_2 \text{ head})$ .

Data from Nicholls [177]

TABLE 72. *Franck-Condon factors for ionizing transitions*(a)  $X^2\Sigma^+(CO^+) \leftarrow X^1\Sigma^+(CO)$  ( $v''=0$ )

$v'$	$q$	$\Sigma q$
0	0.96355	0.96355
1	.03634	.99989
2	.00011	1.00000
3	.00000	1.00000
4	.00000	1.00000
5	.00000	1.00000
6	.00000	1.00000
7	.00000	1.00000
8	.00000	1.00000
9	.00000	1.00000
10	.00000	1.00000

(b)  $A^2\Pi_i(CO^+) \leftarrow X^1\Sigma^+(CO)$  ( $v''=0$ )

$v'$	$q$	$\Sigma q$
0	0.07937	0.07937
1	.17618	.25555
2	.21517	.47072
3	.19212	.66284
4	.14062	.80346
5	.08979	.89320
6	.05187	.94507
7	.02786	.97293
8	.01416	.98709
9	.00690	.99399
10	.00326	.99725

(c)  $B^2\Sigma^+(CO^+) \leftarrow X^1\Sigma^+(CO)$  ( $v''=0$ )

$v'$	$q$	$\Sigma q$
0	0.68853	0.68853
1	.24825	.93678
2	.05248	.98926
3	.00895	.99821
4	.00146	.99967
5	.00026	.99993
6	.00005	.99998
7	.00001	.99999
8	.00000	.99999
9	.00000	.99999
10	.00000	.99999

Data from Wacks [264]. Halmann and Laulicht [88a] have calculated  $q$  for isotopes  $C^{12}O^{16}$ ,  $C^{13}O^{16}$ , and  $C^{12}O^{18}$ .



TABLE 73. (a) *Lifetimes, Einstein coefficients, and oscillator strengths*

Molecule	Transition	Band	Radiative lifetime $\tau$	$A_{v'}$	$A_{v'v''}$	Absorption $f$ -value	Ref.
CO	$A^1\Pi-X^1\Sigma^+$	(a)	$\mu s$	$10^8 s^{-1}$	$10^6 s^{-1}$	0.24	142
		0-0	$11.2(\pm 0.5) \times 10^{-3}$	9.0			102a
		1-0	$10.0(\pm 0.5) \times 10^{-3}$	10.0			102a
		2-0	$10.0(\pm 0.5) \times 10^{-3}$	10.0			102a
	$B^1\Sigma^+-X^1\Sigma^+$	(a)				0.034	142
		0-0	$17 \times 10^{-3}$			0.012	102b
	$C^1\Sigma^+-X^1\Sigma^+$	(a)				$> 0.28$	142
	$b^3\Sigma^+-a^3\Pi$	(b)	$8.6(\pm 0.9) \times 10^{-2}$			$8.9(\pm 1.5) \times 10^{-3}$	237a
CO <sup>+</sup>	$A^2\Pi_i-X^2\Sigma^+e$	1-0	$2.61 \pm 0.20$				19
		1-1	$2.95 \pm 0.22$				
		2-0	$2.48 \pm 0.21$				
		2-1	$2.74 \pm 0.21$				
		3-0	$2.36 \pm 0.15$				
		4-0	$2.22 \pm 0.13$				
		5-0	$2.11 \pm 0.13$				
		Average	2.6			$2.2(\pm 0.5) \times 10^{-3}$	
	$B^2\Sigma^+-X^2\Sigma^+$	0-0			11.4	$8.2 \times 10^{-3}$	143
		0-1			8.0	$6.4 \times 10^{-3}$	
		0-2			2.7	$2.4 \times 10^{-3}$	
		0-3			0.45	$0.4 \times 10^{-3}$	
		( $v'=0$ A v.)	$39.5(\pm 3) \times 10^{-3}$	22.8			
		1-0			8.2	$5.5 \times 10^{-3}$	
		1-1			1.8	$1.3 \times 10^{-3}$	
		1-2			7.0	$5.7 \times 10^{-3}$	
		1-3			4.3	$3.9 \times 10^{-3}$	
		1-4			1.1	$1.1 \times 10^{-3}$	
		( $v'=1$ A v.)	$39.5(\pm 3) \times 10^{-3}$	22.8			237a
		( $v'=2$ A v.)	$38.8(\pm 4) \times 10^{-3}$	23.2			
		(c)	$39.5(\pm 8) \times 10^{-3}$			$18(\pm 3.6) \times 10^{-3}$	
		(o)	$1.01(\pm 0.05) \times 10^{-1}$			$7.9(\pm 0.7) \times 10^{-3}$	
		(d)				2.8	251

<sup>a</sup> Integrated  $f$ -value for the system from inelastic electron scattering. Table 73 (c) lists  $f$ -values for bands of this system.

<sup>b</sup>  $\tau$  and  $f$  for  $v' = 0$ .

\* Average for the B-X transition. The discrepancy between the two values is as yet unexplained.

<sup>d</sup> Continuum, 876–374 Å.

<sup>e</sup> Decay of emission from each band observed. Lifetimes of upper vibrational levels measured.

(b) Einstein coefficients, absolute band strengths, and band oscillator strengths for the  $A^2\Pi_1 - X^2\Sigma^+$  Comet Tail system of  $CO^+$

$v''$				
$v'$	0	1	2	3
0	4.37+4	9.47+4	7.63+4	
	1.01-2	3.11-2	3.66-2	
	1.58-4	4.30-4	4.45-4	
1	1.42+5	1.75+5	3.90+4	
	2.65-2	4.50-2	1.21-2	
	4.44-4	6.74-4	1.70-4	
2	2.37+4	1.13+5		4.33+4.
	3.65-2	2.33-2		1.76-2
	6.50-4	3.75-4		2.26-4
3	2.66+5	1.71+4	7.84+4	4.83+4.
	3.51-2	2.88-3	1.78-2	1.52-2
	6.64-4	4.98-5	2.78-4	2.13-4
4	2.24+5	9.19+3	1.33+5	
	2.14-2	1.29-2	2.42-2	
	4.83-4	2.36-4	4.07-4	
5	1.48+5	6.51+4	6.84+4	
	1.36-2	7.67-2	1.03-2	
	2.88-4	1.49-4	1.85-4	

LEGEND:  $A_{\mathbf{v}'\mathbf{v}''}$  ( $\text{sec}^{-1}$ );  $p_{\mathbf{v}'\mathbf{v}''}$  ( $a_0^2 e^2$ );  $f_{\mathbf{v}'\mathbf{v}''}$ .

Data from Nicholls [177]

TABLE 73—Continued—(c) *Absolute f-values for the  $A^1\Pi - X^1\Sigma^+$  Fourth Positive system*

$\pi' \backslash \pi''$	0	1	2	3	4	5	6	7	8	9	$\Sigma_{\pi''} f_{\pi'\pi''}$
0	9.0	20.0	21.2	14.1	6.7	2.4	-----	-----	-----	-----	73.5
1	20.2	14.0	0.3	6.5	15.8	14.7	8.2	3.3	-----	-----	83.0
2	22.6	1.2	8.3	10.3	0.4	4.8	13.2	12.9	7.4	3.0	84.1

Data from Hesser and Dressler [102a].

TABLE 74. Potential energy of the electronic states of CO and CO<sup>+</sup> \*Potential energy of the X<sup>1</sup>Σ<sup>+</sup> state of CO ( $r_e=1.128322$  Å)<sup>a, c</sup>

<i>v</i>	V	V	<i>r</i> <sub>min</sub>	<i>r</i> <sub>max</sub>
	cm <sup>-1</sup>	eV	Å	Å
0	1081.590	0.1341	1.083	1.179
1	3224.864	.3998	1.053	1.220
2	5341.654	.6623	1.034	1.250
3	7432.026	.9214	1.020	1.276
4	9496.048	1.1773	1.008	1.300
5	11533.76	1.4300	0.997	1.322
6	13545.29	1.6794	.988	1.343
7	15530.64	1.9255	.980	1.363
8	17489.9	2.1684	.972	1.383
9	19423.5	2.4081	.965	1.402
10	21330.7	2.6446	.959	1.420
11	23212.4	2.8779	.953	1.438
12	25068.5	3.1080	.947	1.456
13	26898.4	3.3349	.942	1.474
14	28703.2	3.5587	.937	1.492
15	30482.7	3.7793	.932	1.509
16	32235.4	3.9966	.928	1.526
17	33963.9	4.2109	.923	1.544
18	35666.8	4.4220	.919	1.561
19	37344.4	4.6300	.916	1.579
20	38997.5	4.8349	.912	1.596
21	40625.7	5.0368	.908	1.614
22	42208	5.2330	.905	1.631
23	43776	5.4274	.901	1.649
<sup>b</sup> 24	45307	5.6172	.898	1.668
<sup>b</sup> 25	46803	5.8027	.894	1.686

\*Data for the e<sup>2</sup>Σ<sup>-</sup> state [138] are not included because of a revision in vibrational quantum numbering [238a]. The newly revised numbering is definitive (sec. 3.10).

<sup>a</sup> Data for *v* = 0 to 7 from vibration-rotation spectra.

<sup>b</sup> Extrapolated to highest observed level.

<sup>c</sup> *r*<sub>e</sub>: See footnote X<sup>1</sup>Σ<sup>+</sup>, table 1.

Data taken from Krupenie and Weissman [138].

Potential energy of the a<sup>3</sup>Π state of CO ( $T_e=48687.40$  cm<sup>-1</sup>=6.0363 eV,  $r_e=1.2058$  Å)

<i>v</i>	V	V	T <sub>e</sub> +V	T <sub>e</sub> +V	<i>r</i> <sub>min</sub>	<i>r</i> <sub>max</sub>
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	868.16	0.1076	49555.56	6.1440	1.157	1.263
1	2582.77	.3202	51270.17	6.3565	1.124	1.310
2	4266.7	.5290	52954.1	6.5653	1.103	1.344
3	5921.2	.7341	54608.6	6.7704	1.087	1.375
4	7546.8	.9357	56234.2	6.9720	1.074	1.403
5	9143.3	1.1336	57830.7	7.1699	1.063	1.429
<sup>a</sup> 6	10710	1.3278	59398	7.3642	1.052	1.454
<sup>a</sup> 7	12246	1.5183	60934	7.5547	1.043	1.478

<sup>a</sup> Extrapolated to highest observed level.

TABLE 74. Potential energy of the electronic states of CO and CO<sup>+</sup>—ContinuedPotential energy of the a'<sup>3</sup>Σ<sup>+</sup> state of CO ( $T_e=55822.92$  cm<sup>-1</sup>=6.9210 eV,  $r_e=1.3519$  Å)

<i>v</i>	V	V	T <sub>e</sub> +V	T <sub>e</sub> +V	<i>r</i> <sub>min</sub>	<i>r</i> <sub>max</sub>
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	612.581	0.0759	56435.50	6.9969	1.296	1.423
1	1821.44	.2258	57644.36	7.1468	1.257	1.478
2	3008.83	.3730	58831.75	7.2940	1.232	1.520
3	4175.16	.5176	59998.08	7.4386	1.214	1.557
4	5321.31	.6597	61144.23	7.5807	1.198	1.591
5	6446.32	.7992	62269.04	7.7202	1.185	1.622
6	7552.31	.9363	63375.23	7.8573	1.174	1.653
7	8638.11	1.0710	64461.03	7.9920	1.164	1.682
8	9705.57	1.2033	65528.49	8.1243	1.155	1.711
9	10751.15	1.3329	66574.07	8.2539	1.147	1.739
10	11779.26	1.4604	67602.18	8.3814	1.139	1.766
11	12788.37	1.5855	68611.29	8.5065	1.132	1.793
12	13778.58	1.7083	69601.50	8.6293	1.126	1.820
13	14750.12	1.8287	70573.33	8.7497	1.120	1.847
14	15703.24	1.9469	71526.16	8.8679	1.114	1.874
15	16637.78	2.0628	72460.70	8.9838	1.108	1.900
16	17554.01	2.1764	73376.93	9.0974	1.103	1.927
17	18451.41	2.2876	74274.33	9.2086	1.099	1.954
18	19330.47	2.3966	75153.39	9.3176	1.094	1.981
19	20190.91	2.5033	76013.83	9.4243	1.090	2.009
20	21032.78	2.6077	76855.70	9.5287	1.086	2.036
21	21855.93	2.7097	77678.85	9.6307	1.082	2.064
22	22660.14	2.8094	78483.06	9.7304	1.079	2.092
23	23443.67	2.9066	79266.59	9.8276	1.075	2.121

Potential energy of the d<sup>3</sup>Δ<sub>i</sub> state of CO ( $T_e=61154.06$  cm<sup>-1</sup>=7.5820 eV,  $r_e=1.3700$  Å)

<i>v</i>	V	V	T <sub>e</sub> +V	T <sub>e</sub> +V	<i>r</i> <sub>min</sub>	<i>r</i> <sub>max</sub>
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	574.46	0.0712	61728.52	7.6532	1.309	1.440
1	1712	.2123	62866	7.7942	1.270	1.498
2	2834	.3513	63988	7.9333	1.245	1.541
3	3940	.4885	65094	8.0705	1.226	1.579
4	5029	.6235	66183	8.2055	1.211	1.613
5	6100	.7563	67254	8.3383	1.197	1.645
6	7153	.8869	68307	8.4688	1.186	1.676
7	8187	1.0151	69341	8.5970	1.176	1.707
8	9202	1.1409	70356	8.7228	1.167	1.736
9	10196	1.2641	71350	8.8461	1.158	1.765
10	11169	1.3848	72323	8.9667	1.150	1.794
11	12121	1.5027	73275	9.0847	1.143	1.823
12	13050	1.6179	74204	9.1999	1.136	1.852
13	13956	1.7303	75110	9.3123	1.130	1.881
14	14839	1.8397	75993	9.4217	1.124	1.910
15	15697	1.9461	76851	9.5281	1.118	1.940
16	16530	2.0494	77684	9.6314	1.112	1.970
17	17337	2.1495	78492	9.7315	1.106	2.000
18	18118	2.2463	79273	9.8283	1.101	2.032
19	18873	2.3398	80027	9.9218	1.095	2.064
20	19599	2.4299	80753	10.0118	1.090	2.097
21	20297	2.5164	81451	10.0984	1.084	2.132
22	20966	2.5993	82120	10.1813	1.079	2.167



TABLE 74. Potential energy of the electronic states of CO and CO<sup>+</sup>—Continued

Potential energy of the A<sup>1</sup>Π state of CO ( $T_e=65074.6$  cm<sup>-1</sup>=8.0680 eV,  $r_e=1.2352$  Å)

<i>v</i>	V	V	$T_e+V$	$T_e+V$	$r_{min}$	$r_{max}$
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	753.49	0.0934	65828.1	8.1614	1.183	1.297
1	2242.3	.2780	67316.9	8.3460	1.148	1.348
2	3685.2	.4569	68759.8	8.5249	1.126	1.387
3	5097.9	.6320	70172.5	8.7001	1.110	1.422
4	6476.1	.8029	71550.7	8.8709	1.096	1.454
5	7818.2	.9693	72892.8	9.0373	1.085	1.484
6	9125.0	1.1313	74199.6	9.1993	1.075	1.514
7	10401.8	1.2896	75476.4	9.3576	1.066	1.542
8	11641.0	1.4433	76715.6	9.5113	1.058	1.571
9	12846.9	1.5928	77921.5	9.6608	1.051	1.599
10	14018.8	1.7381	79093.4	9.8061	1.044	1.627
11	15155.6	1.8790	80230.2	9.9470	1.038	1.656
12	16262	2.0162	81336	10.0841	1.032	1.685
13	17333	2.1490	82407	10.2169	1.027	1.714
14	18371	2.2777	83445	10.3456	1.022	1.743
15	19366	2.4010	84440	10.4690	1.017	1.773
16	20330	2.5205	85404	10.5885	1.013	1.804
17	21257	2.6355	86331	10.7034	1.009	1.837
18	22138	2.7447	87212	10.8126	1.005	1.871
a 19	22970	2.8478	88045	10.9159	1.001	1.909
a 20	23762	2.9460	88837	11.0141	0.996	1.951

a Extrapolated to highest observed level.

Potential energy of the I<sup>1</sup>Σ<sup>-</sup> state of CO ( $T_e=66185$  cm<sup>-1</sup>=8.2057 eV,  $r_e=1.416$  Å)

<i>v</i>	V	V	$T_e+V$	$T_e+V$	$r_{min}$	$r_{max}$
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	530	0.066	66715	8.271	1.35	1.48
1	1580	.196	67760	8.401	1.29	1.53
2	2600	.324	68790	8.528	1.25	1.56
3	3610	.448	69790	8.653	1.22	1.59
4	4600	.571	70780	8.776	1.19	1.61
5	5560	.690	71750	8.896	1.16	1.64
6	6510	.808	72700	9.013	1.14	1.66
7	7450	.924	73630	9.129	1.12	1.68
8	8360	1.037	74540	9.242	1.09	1.69
9	9250	1.148	75440	9.353	1.07	1.71
10	10120	1.257	76310	9.461	1.05	1.73
11	10980	1.363	77160	9.567	1.03	1.75
12	11810	1.467	78000	9.670	1.01	1.77

Potential energy of the B<sup>1</sup>Σ<sup>+</sup> state of CO ( $T_e=86928$  cm<sup>-1</sup>=10.7774 eV,  $r_e=1.120$  Å)<sup>a</sup>

<i>v</i>	V	V	$T_e+V$	$T_e+V$	$r_{min}$	$r_{max}$
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	1072	0.1329	87998	10.9101	1.08	1.17
1	3154	.3910	90080	11.1682	1.05	1.22
2	5154	.6390	92080	11.4162	1.03	1.25

a Curve taken from ref. [258] but with different Z.P.E.

TABLE 74. Potential energy of the electronic states of CO and CO<sup>+</sup>—Continued

Potential energy of the X<sup>2</sup>Σ<sup>+</sup> state of CO<sup>+</sup> ( $T_e=113007$  cm<sup>-1</sup>=14.0108 eV,  $r_e=1.11516$  Å)

<i>v</i>	V	V	$T_e+V$	$T_e+V$	$r_{min}$	$r_{max}$
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	1103.33	0.1368	114110.33	14.1475	1.071	1.165
1	3287.22	0.4076	116294.22	14.4183	1.041	1.206
2	5440.80	0.6746	118447.80	14.6853	1.023	1.236
3	7564.01	0.9378	120571.01	14.9486	1.008	1.263
4	9656.94	1.1973	122663.94	15.2080	0.996	1.286
5	11719.47	1.4530	124726.47	15.4637	0.986	1.309
6	13751.63	1.7049	126758.63	15.7157	0.977	1.330
7	15753.47	1.9531	128760.47	15.9639	0.968	1.350
8	17724.95	2.1976	130731.95	16.2083	0.961	1.370
9	19666.6	2.4383	132676.6	16.4490	0.954	1.390
10	21578.2	2.6753	134585.2	16.6860	0.948	1.409
11	23459.2	2.9085	136466.2	16.9193	0.943	1.428
12	25310.0	3.1380	138317.0	17.1487	0.938	1.447
13	27102.3	3.3602	140109.3	17.3709	0.933	1.466

Potential energy of the A<sup>2</sup>Π<sub>i</sub> state of CO<sup>+</sup> ( $T_e=133740$  cm<sup>-1</sup>=16.5813 eV,  $r_e=1.24378$  Å)

<i>v</i>	V	V	$T_e+V$	$T_e+V$	$r_{min}$	$r_{max}$
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	777.65	0.0964	134517.65	16.6777	1.191	1.304
1	2312.88	.2868	136052.88	16.8680	1.157	1.354
2	3820.74	.4737	137560.74	17.0550	1.136	1.391
3	5301.83	.6573	139041.83	17.2386	1.119	1.424
4	6756.38	.8377	140496.38	17.4189	1.105	1.453
5	8184.11	1.0147	141924.11	17.5959	1.094	1.481
6	9585.20	1.1884	143325.20	17.7697	1.083	1.508
7	10959.79	1.3588	144699.79	17.9401	1.074	1.534
8	12307.90	1.5259	146047.90	18.1072	1.066	1.560
9	13629.46	1.6898	147369.46	18.2711	1.059	1.585
10	14924.72	1.8504	148664.72	18.4317	1.052	1.610
11	16193.82	2.0077	149933.82	18.5890	1.046	1.634

Potential energy of the B<sup>2</sup>Σ<sup>+</sup> state of CO<sup>+</sup> ( $T_e=158384$  cm<sup>-1</sup>=19.6987 eV,  $r_e=1.16878$  Å)

<i>v</i>	V	V	$T_e+V$	$T_e+V$	$r_{min}$	$r_{max}$
	cm <sup>-1</sup>	eV	cm <sup>-1</sup>	eV	Å	Å
0	860.15	0.1066	159744.15	19.8053	1.120	1.227
1	2539.55	.3149	161423.55	20.0135	1.089	1.277
2	4166.05	.5165	163050.05	20.2152	1.069	1.315
3	5741.62	.7119	164625.62	20.4105	1.054	1.349
4	7268.23	.9011	166152.23	20.5998	1.042	1.381
5	8749.8	1.0848	167633.8	20.7835	1.032	1.411
6	10186.9	1.2630	169070.9	20.9616	1.023	1.441
7	11588.7	1.4368	170472.7	21.1354	1.015	1.471
8	12920.7	1.6019	171804.7	21.3066	1.008	1.500
9	14193.8	1.7598	173077.8	21.4584	1.001	1.531
a 10	15390.1	1.9081	174274.1	21.6067	0.993	1.564

a Extrapolated to highest observed level.

FIGURE 1. Potential energy curves for CO and CO<sup>+</sup>.

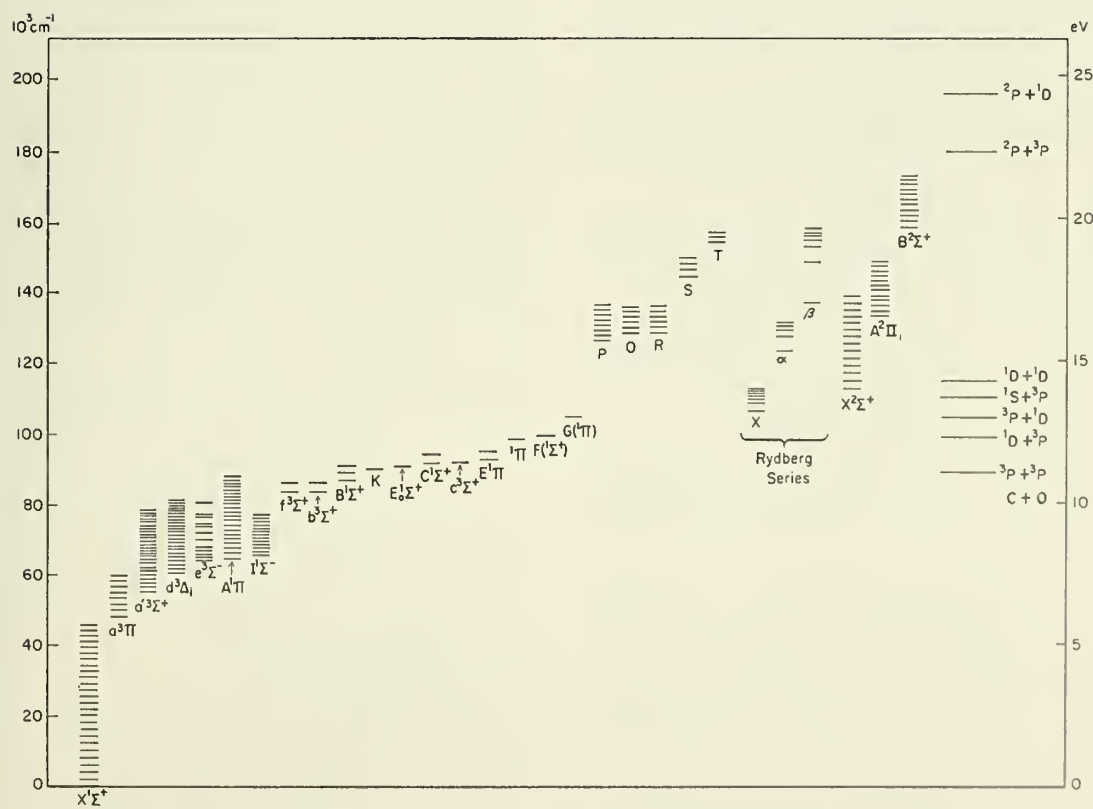
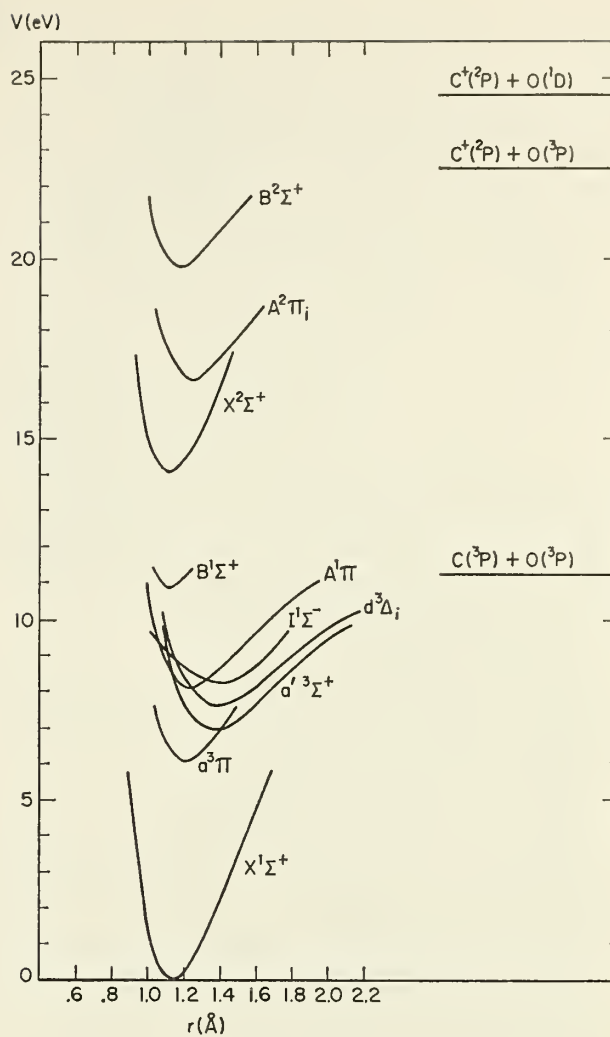


FIGURE 2. Energy level diagram for CO and CO<sup>+</sup>.



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## Appendix A. Notation and Terminology

The spectroscopic notation used in this report is that adopted in Herzberg's book [99] as modified by recommendations of the Triple Commission on Spectroscopy (*J. Opt. Soc. Am.* **43**, 425-30 (1953); **52**, 476-7 (1962); **53**, 883-5 (1963)). A number of specific conventions used are itemized below.

- (1) Wavenumber in  $cm^{-1}$  is denoted by  $\sigma$ ;  $\nu$  is reserved for frequency in  $s^{-1}$ .
- (2)  $N$  is total angular momentum of electrons and nuclei exclusive of spin (case  $b$ ,  $b'$ ,  $d$ ), formerly denoted by  $K$ .
- (3) Rotational angular momentum of the nuclei,



formerly denoted by N, is now denoted by O.

- (4) Dissociation energy is written as  $D^0$  or  $D^e$ ; rotational constants (for the zero level or equilibrium value, respectively) are denoted as usual by  $D_0$  and  $D_e$ .
- (5) A transition is always represented with a dash, as  ${}^2\Pi-{}^2\Sigma$  transition. The upper state is always written first.  $\leftarrow$  means absorption;  $\rightarrow$  means emission, for an electronic, vibration-rotation, or rotational (here, microwave) transition.
- (6) A perturbation by one state of another is indicated as e.g., ( ${}^1\Delta^1\Pi$ ) perturbation, following an early notation of Kovács. (Conventions used in some early papers include  ${}^1\Delta\times{}^1\Pi$  or  ${}^1\Delta$ ,  ${}^1\Pi$ .)
- (7) A progression of bands is indicated as follows; e.g.,
  - (a)  $v''=0$  progression
  - (b)  $(v'-0)$  progression.
- (8) Reciprocal dispersion is given in  $\text{\AA}/\text{mm}$ . However, following the colloquial use of many spectroscopists this quantity is referred to as dispersion.
- (9) In the tables, wavelengths above 2000  $\text{\AA}$  are air wavelengths unless otherwise specified; below 2000  $\text{\AA}$  vacuum wavelengths are listed.
- (10) The known band degradation is indicated by R (red-degraded) or V (violet-degraded) in the headings of section 3 as well as in tables 1 and 3 to 35.
- (11) Rotational constants in tables 36 to 52 are given in units of  $\text{cm}^{-1}$ .
- (12) Abbreviations used are listed on p. 2. In addition, zero-point energy is abbreviated by ZPE.

The following items apply to table 1.

- (1) Vibrational constants and term values T are assumed to be derived from data on band origins. (Herzberg [99] denotes these by the letter Z.) In conformity with Herzberg, H denotes constants derived from head measurements.
- (2)  $[r_e]$  means  $r_0$ ;  $[B_e]$  means  $B_0$ ;  $[\omega_e]$  means  $\Delta G(\frac{1}{2})$ , as in Herzberg's book.
- (3) ( ) means uncertain.
- (4)  $T_0$  is the mean height (in case of multiplets) above X,  $v=0$ ,  $J=0$ .
- (5) States which are predissociated have Pr written in the column for dissociation

products. (For CO all known predissociations are due to  ${}^3P+{}^3P$  states.)

- (6) All numerical data are in units of  $\text{cm}^{-1}$  unless otherwise indicated.
- (7) References cited include only those from which the numerical data have been extracted. Other pertinent references are cited in the appropriate sections of this report.
- (8) The tabulated molecular constants have been taken from the references cited in table 1 but are not necessarily those which best fit the individual vibrational term values or rotational constants where these have been compiled from various sources.
- (9) To avoid confusion of sign conventions several formulas are listed below:

Vibrational terms:

$$G(v) = \omega_e(v + \frac{1}{2}) - \omega_ex_e(v + \frac{1}{2})^2 + \omega_ey_e(v + \frac{1}{2})^3 + \omega_ez_e(v + \frac{1}{2})^4$$

i.e., a negative value of  $\omega_ex_e$  from the table would mean a positive anharmonic term.

Rotational terms:

$$F_v(J) = B_vJ(J+1) - D_vJ^2(J+1)^2 + H_vJ^3(J+1)^3$$

where  $(-D_v)$  is always  $<0$ , and

$$B_v = B_e - \alpha_e(v + \frac{1}{2}) + \gamma_e(v + \frac{1}{2})^2 + \delta_e(v + \frac{1}{2})^3$$

$$D_v = D_e + \beta_e(v + \frac{1}{2}) \text{ [for CO, } X^1\Sigma^+, \beta < 0]$$

$$H_v \sim H_e.$$

- (10) Footnotes which give supplementary information pertaining to the individual electronic states are indicated at the end of table 1 and are identified by the electronic state. Table 1 has been left free of superscripts.
- (11) The tabulated ZPE do not include the Dunham correction which would add an amount given by

$$Y_{00} = \frac{B_e}{4} + \frac{\alpha_e\omega_e}{12B_e} + \frac{\alpha_e^2\omega_e^2}{144B_e^3} - \frac{\omega_ex_e}{4}.$$

(For the ground state this correction is  $0.600 \text{ cm}^{-1}$ .)

## Appendix B. Physical Constants<sup>10-12</sup> and Conversion Factors

$$\begin{aligned}
 c &= 2.997925 \times 10^{10} \text{ cm-s}^{-1} \\
 h &= 6.6256 \times 10^{-27} \text{ erg-s} \\
 N_0 &= 6.02252 \times 10^{23} \text{ mole}^{-1} \\
 1 \text{ eV} &= 8065.73 \text{ cm}^{-1} = 23.0609 \text{ k-cal-mole}^{-1} \\
 \mu_A(\text{C}^{12}\text{O}^{16}) &= 6.85621 \\
 \mu_A(\text{CO}^+) &= 6.85603 \\
 \mu &= \mu_A/N_0: \mu(\text{CO}) = 1.13843 \times 10^{-23} \text{ g} \\
 &\quad \mu(\text{CO}^+) = 1.13840 \times 10^{-23} \text{ g}
 \end{aligned}$$

<sup>10</sup> The molecular reduced masses are calculated from the data of Everling et al., Relative Nuclidic Masses, Nuclear Phys. 18, 529-69 (1960) which are based on the unified atomic weight scale with C<sup>12</sup>=12.

<sup>11</sup> The mass  $\mu_A(\text{CO}^+)$  was calculated by assuming  $5.49 \times 10^{-4}$  a.m.u. for the mass of the electron and assuming that ionization removes an electron from the C atom (1 a.m.u.=1/12 mass of C<sup>12</sup>).

<sup>12</sup> The universal constants and conversion factors are those recommended by the NAS-NRC (Phys. Today, pp. 48-9 (Feb. 1964)).

$$\begin{aligned}
 &\text{Atomic weights:} \\
 &\quad \text{C} = 12 \\
 &\quad \text{O} = 15.994915 \\
 &k_e(\text{CO}) = 40.3930 \times 10^{-2} \omega_e^2 \\
 &\quad (\omega_e \text{ in cm}^{-1}; k_e \text{ in dyne-cm}^{-1}) \\
 &k_e(\text{CO}^+) = 40.3920 \times 10^{-2} \omega_e^2 \\
 &r_e(\text{CO}) = 1.568031 \sqrt{\frac{1}{B_e}} \\
 &\quad (B_e \text{ in cm}^{-1}; r_e \text{ in \AA}) \\
 &r_e(\text{CO}^+) = 1.568052 \sqrt{\frac{1}{B_e}}
 \end{aligned}$$

The "Tabelle der Schwingungszahlen" of Kayser has been superseded by NBS Monograph 3, "Table of Wavenumbers," by Coleman, Bozman, and Meggers (1960) which is based on the 1953 formula for dispersion in standard air of Edlen. For low resolution work the older tables are adequate.





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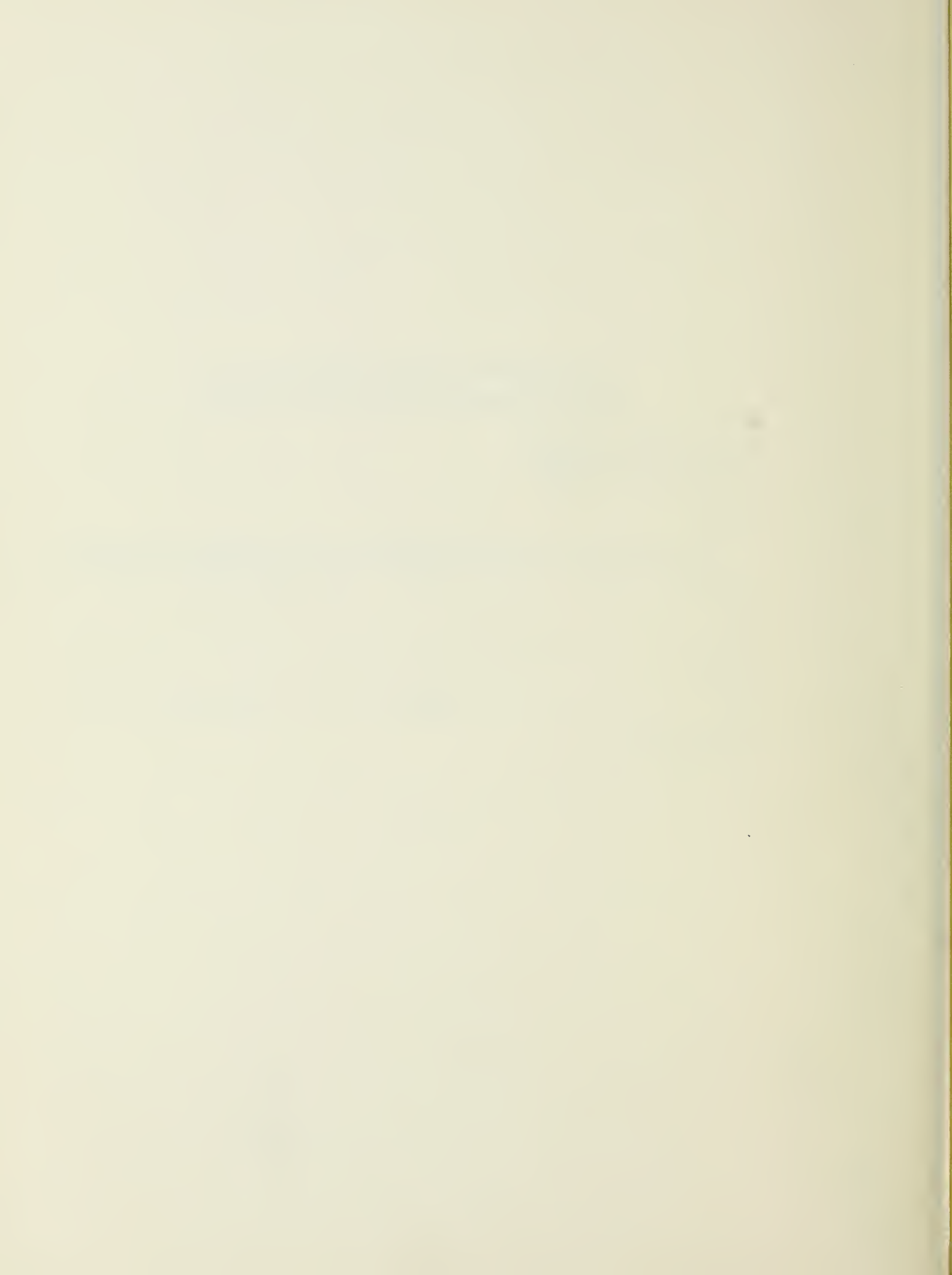
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